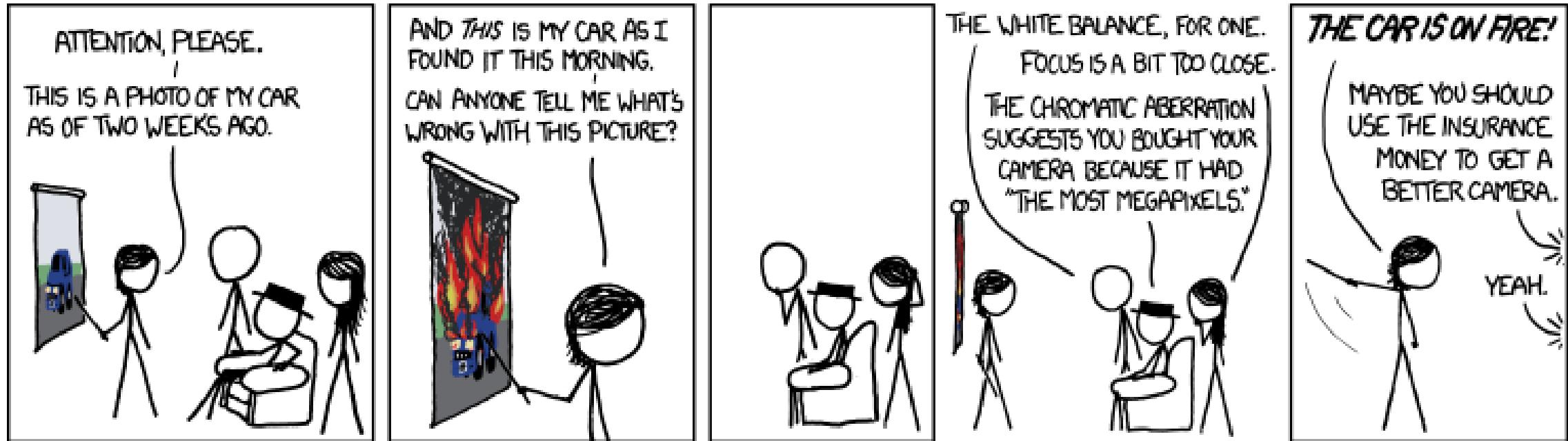


# Digital photography



# Course announcements

- Homework 4 has been posted.
  - Due Friday March 23<sup>rd</sup> (one-week homework!)
  - Any questions about the homework?
  - How many of you have looked at/started/finished homework 4?
- Talk this week: Katie Bouman, “Imaging the Invisible”.
  - Wednesday, March 21st 10:00 AM GHC6115.
  - How many of you attended this talk?

# Overview of today's lecture

- Leftover from color lecture.
- Imaging sensor primer.
- Color sensing in cameras.
- In-camera image processing pipeline.
- Some general thoughts on the image processing pipeline.
- Radiometric calibration (a.k.a. HDR imaging)
- Color calibration.

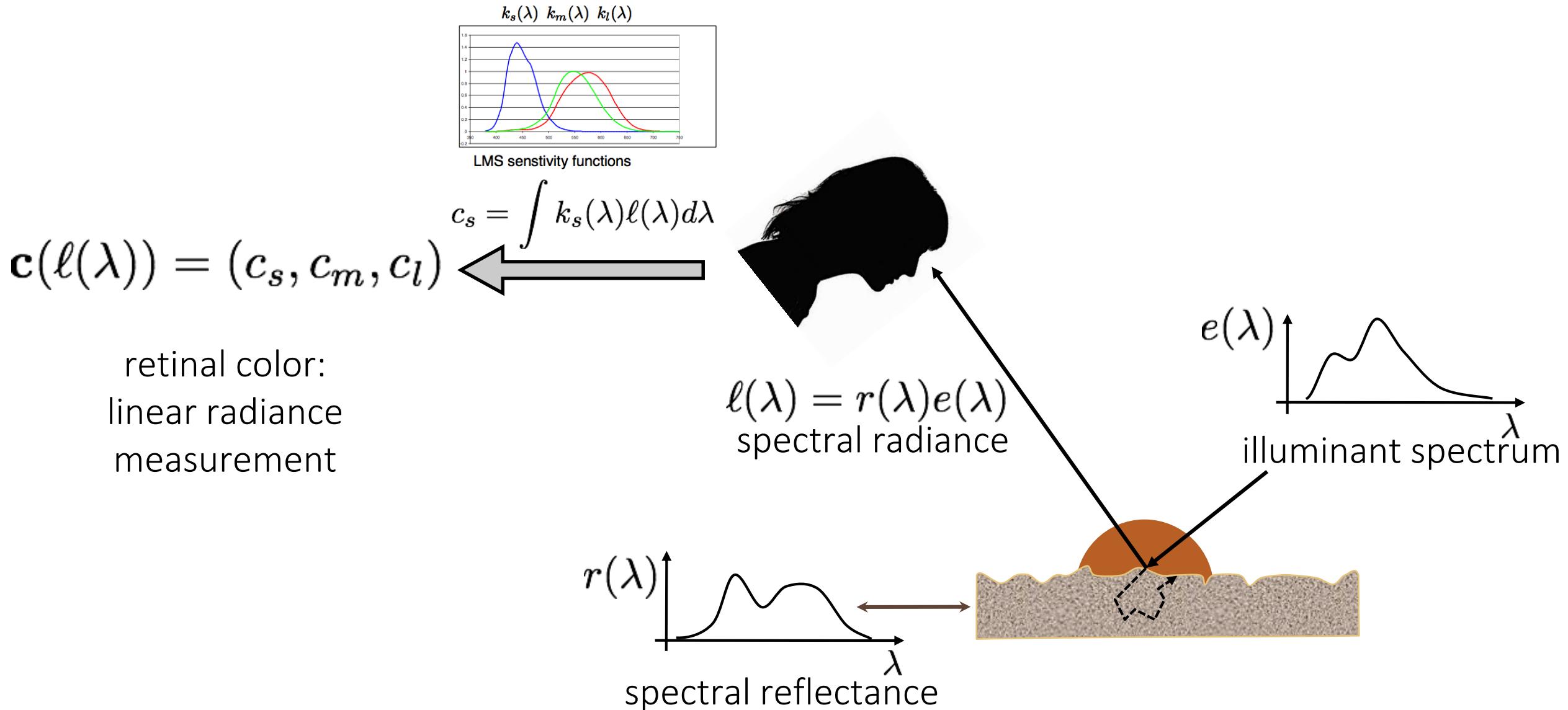
Take-home message: The values of pixels in a photograph and the output of your camera's sensor are two very different things.

# Slide credits

A lot of inspiration and quite a few examples for these slides were taken directly from:

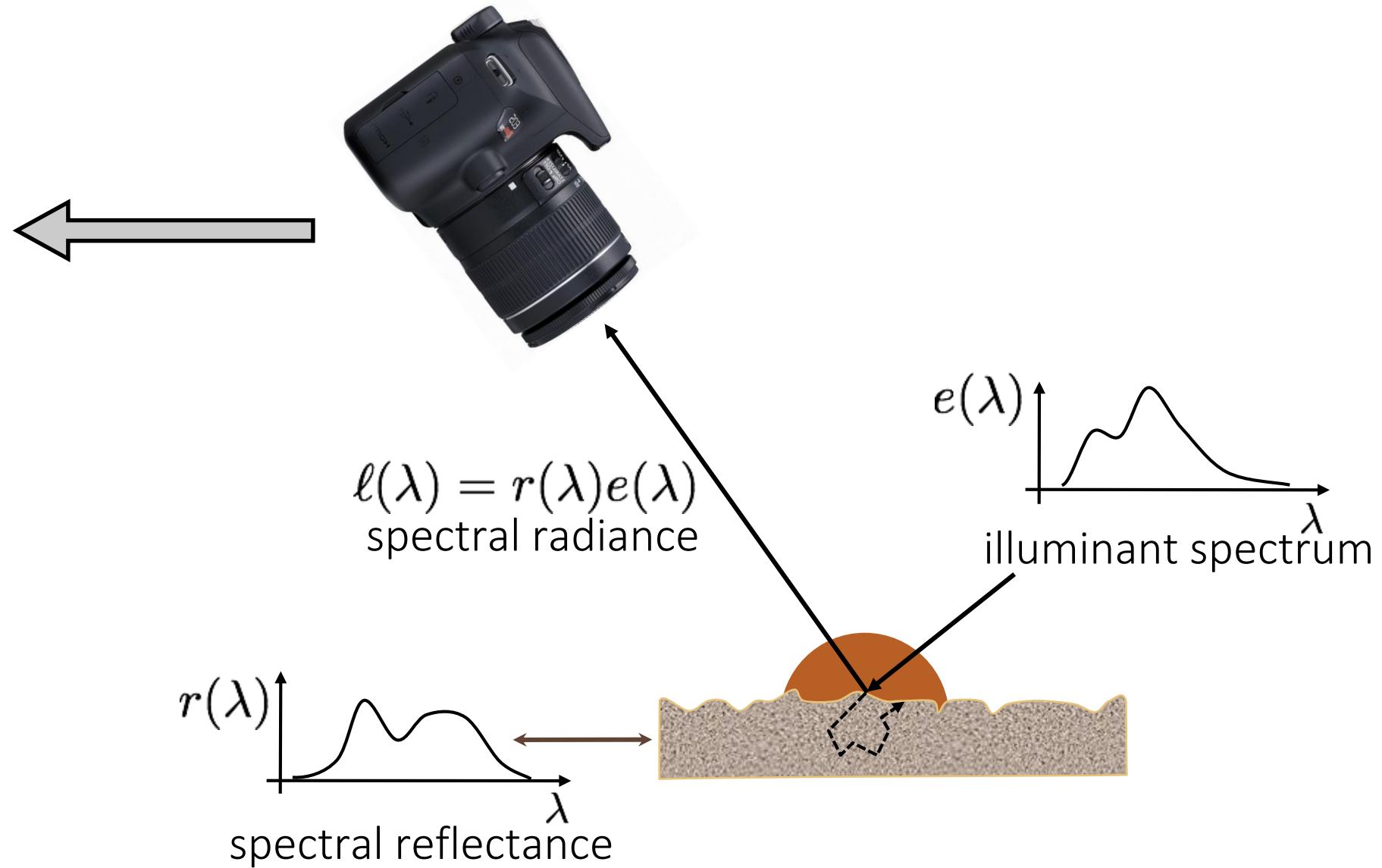
- Kayvon Fatahalian (15-769, Fall 2016).
- Michael Brown (CVPR 2016 Tutorial on understanding the image processing pipeline).

# Human visual system



# Digital imaging system

What functional of  
radiance are we  
measuring?



# The modern photography pipeline

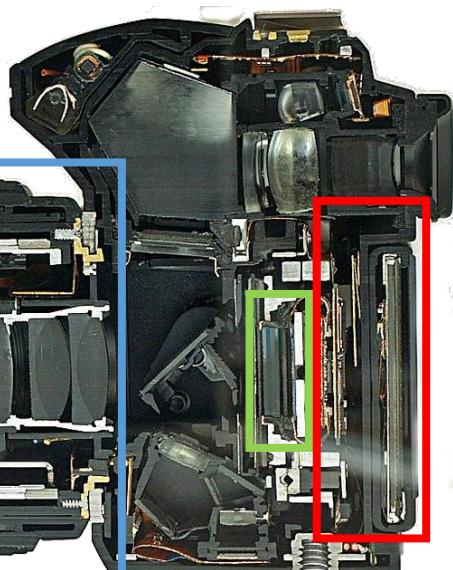




# The modern photography pipeline



post-capture processing  
(16-385, 15-463)



optics and  
optical controls

(15-463)

sensor, analog  
front-end, and  
color filter array

(this lecture)

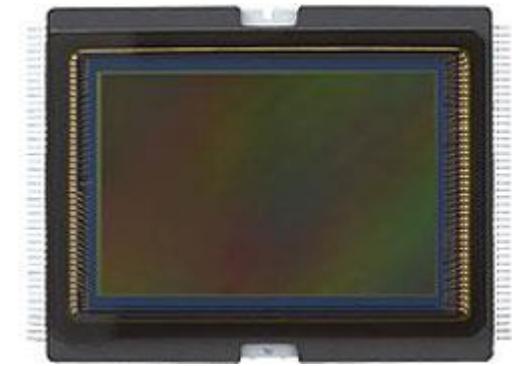
in-camera image  
processing  
pipeline

(this lecture)

# Imaging sensor primer

# Imaging sensors

- Very high-level overview of digital imaging sensors.
- We could spend an entire course covering imaging sensors.

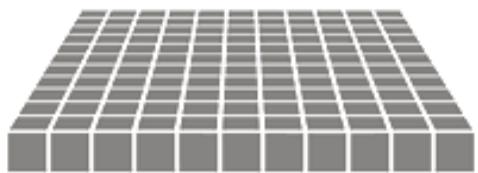


Canon 6D sensor  
(20.20 MP, full-frame)

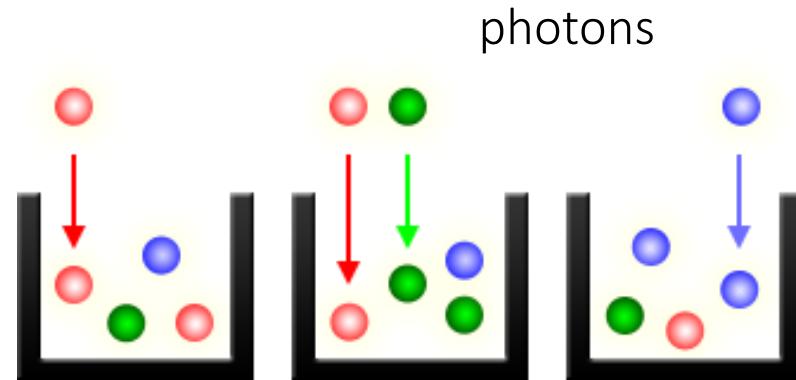
# What does an imaging sensor do?

When the camera shutter opens...

... exposure begins...



array of photon buckets

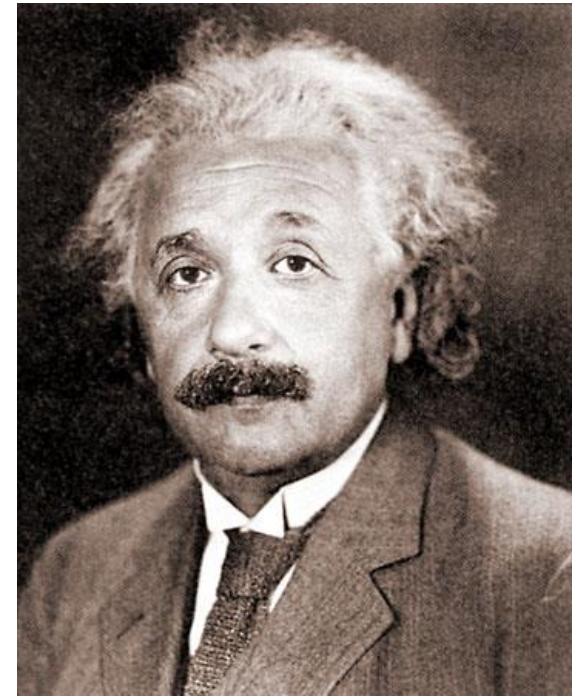


close-up view of photon buckets

... photon buckets begin to store photons...

... until the camera shutter closes. Then, they convert stored photons to intensity values.

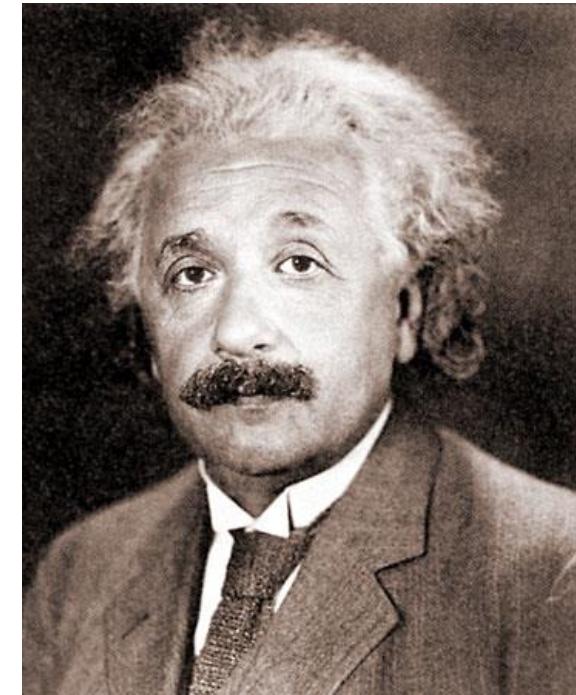
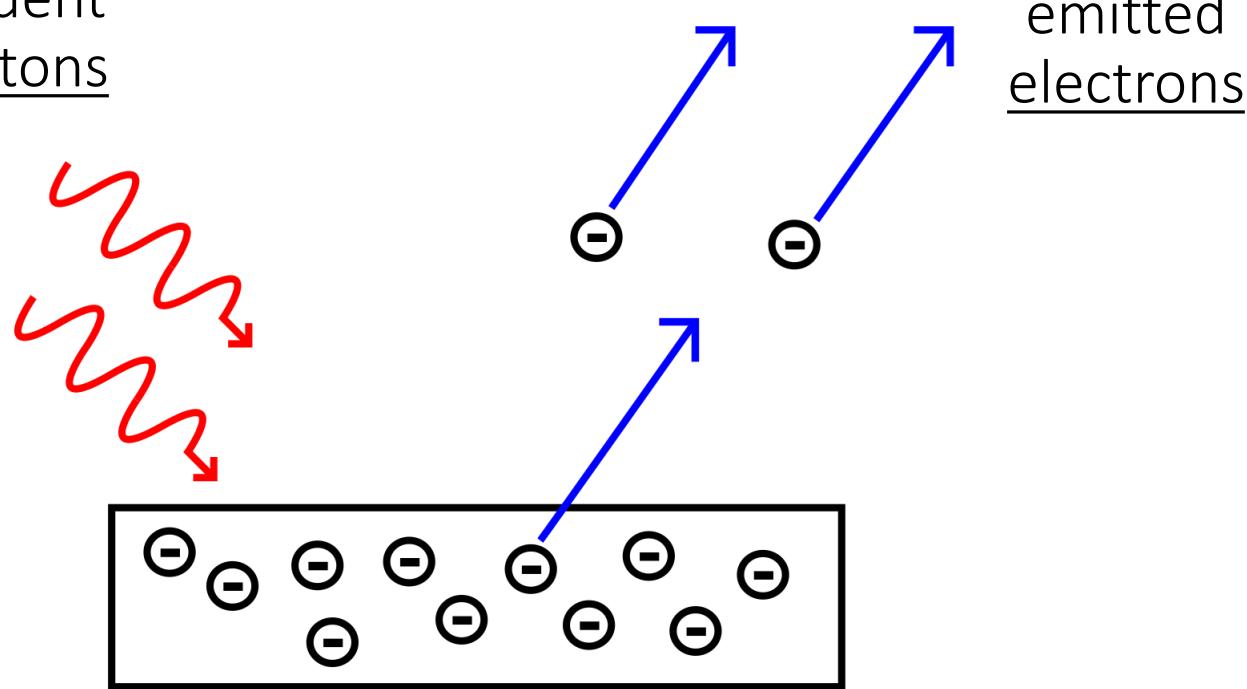
# Nobel Prize in Physics



Do you know who this is?

# Photoelectric effect

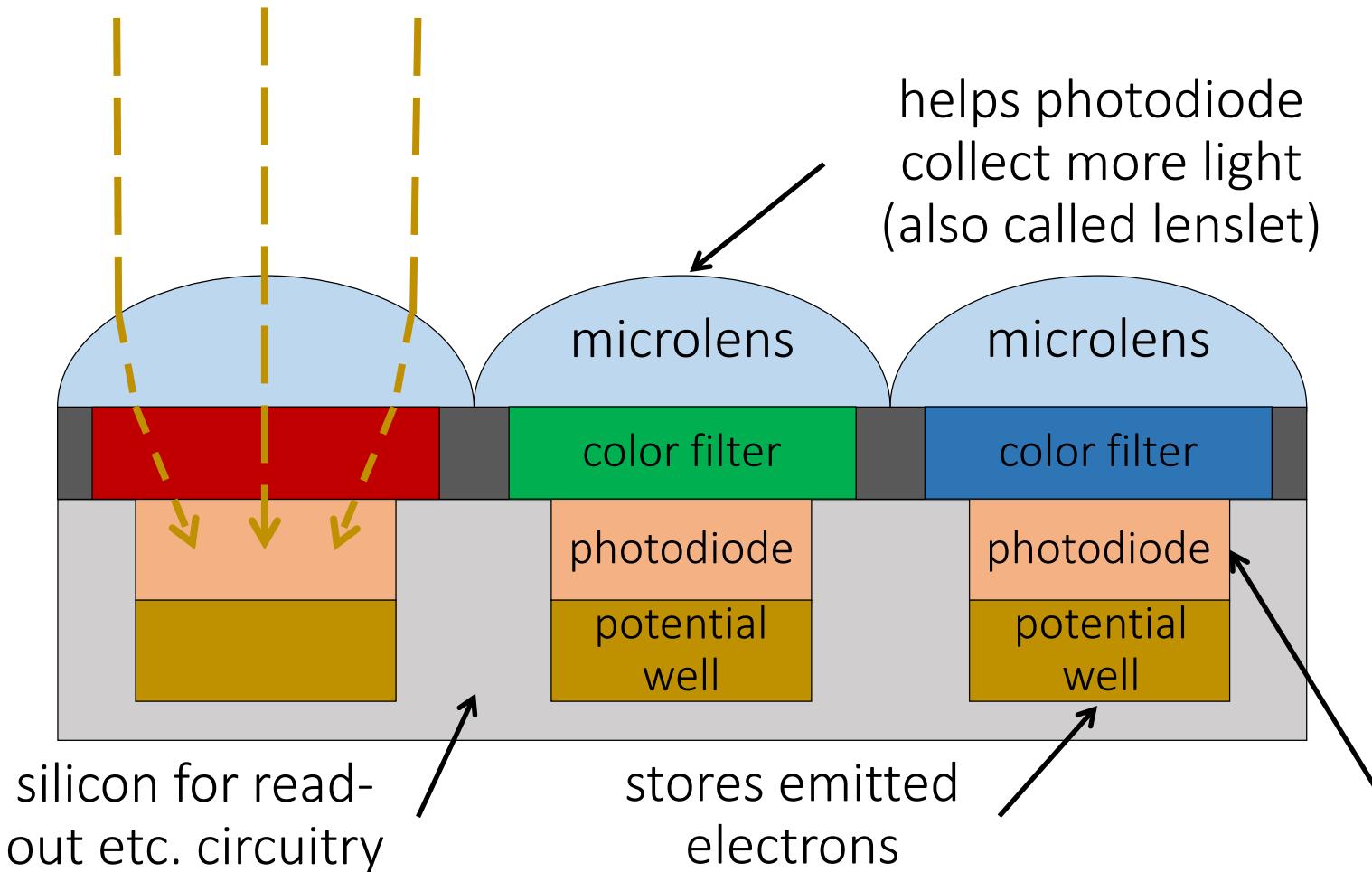
incident  
photons



Albert Einstein

Einstein's Nobel Prize in 1921 "for his services to Theoretical Physics, and especially for his discovery of the law of the photoelectric effect"

# Basic imaging sensor design



- Lenslets also filter the image to avoid resolution artifacts.
- Lenslets are problematic when working with coherent light.
- Many modern cameras do not have lenslet arrays.

We will discuss these issues in more detail at a later lecture.

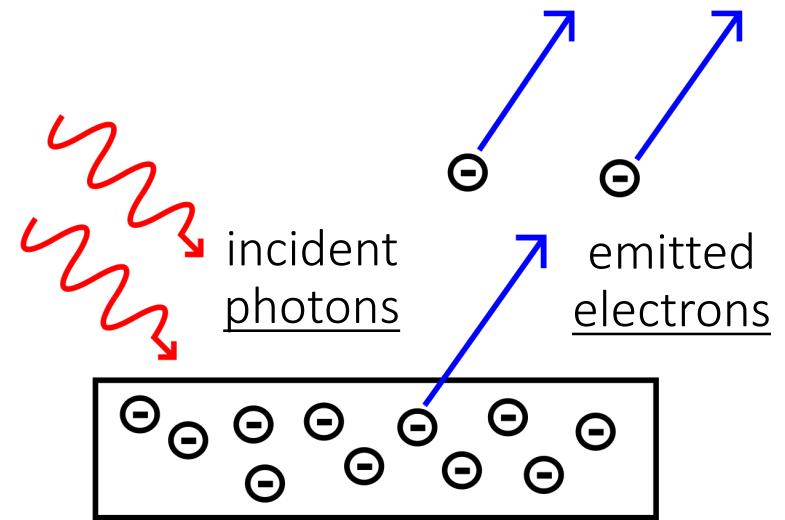
made of silicon, emits electrons from photons

We will see what the color filters are for later in this lecture.

# Photodiode quantum efficiency (QE)

How many of the incident photons will the photodiode convert into electrons?

$$QE = \frac{\# \text{ electrons}}{\# \text{ photons}}$$



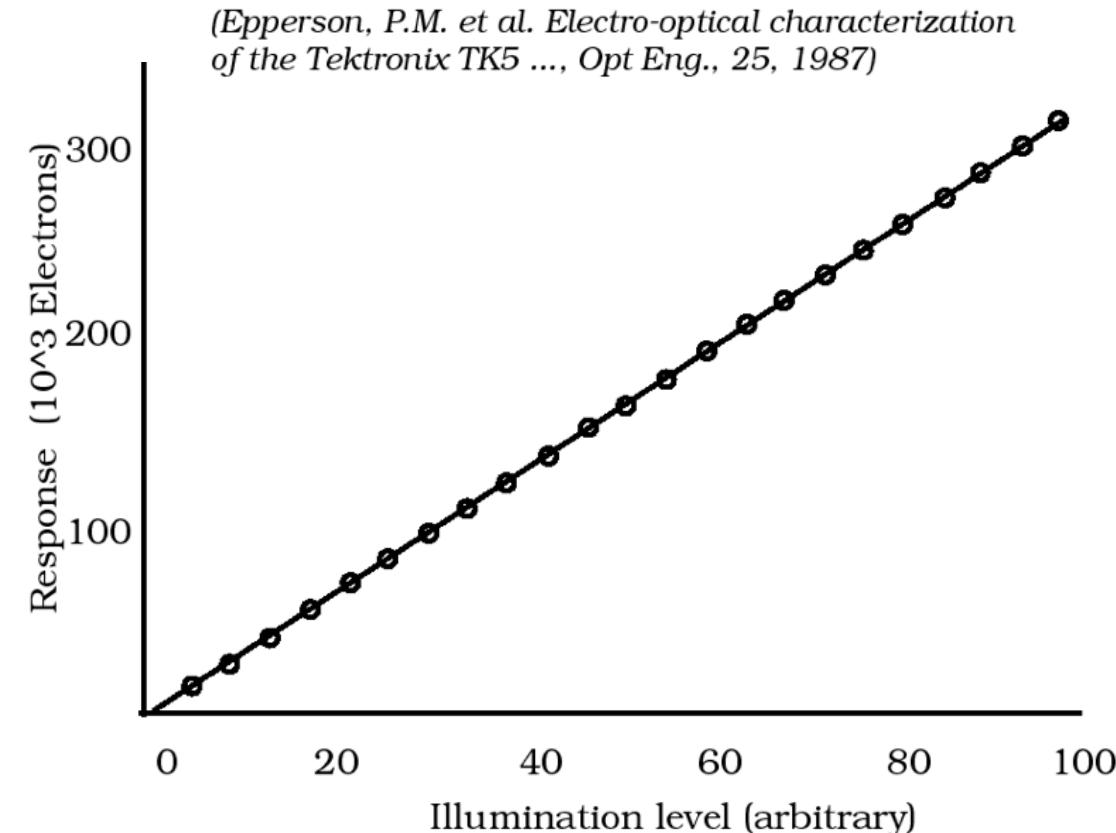
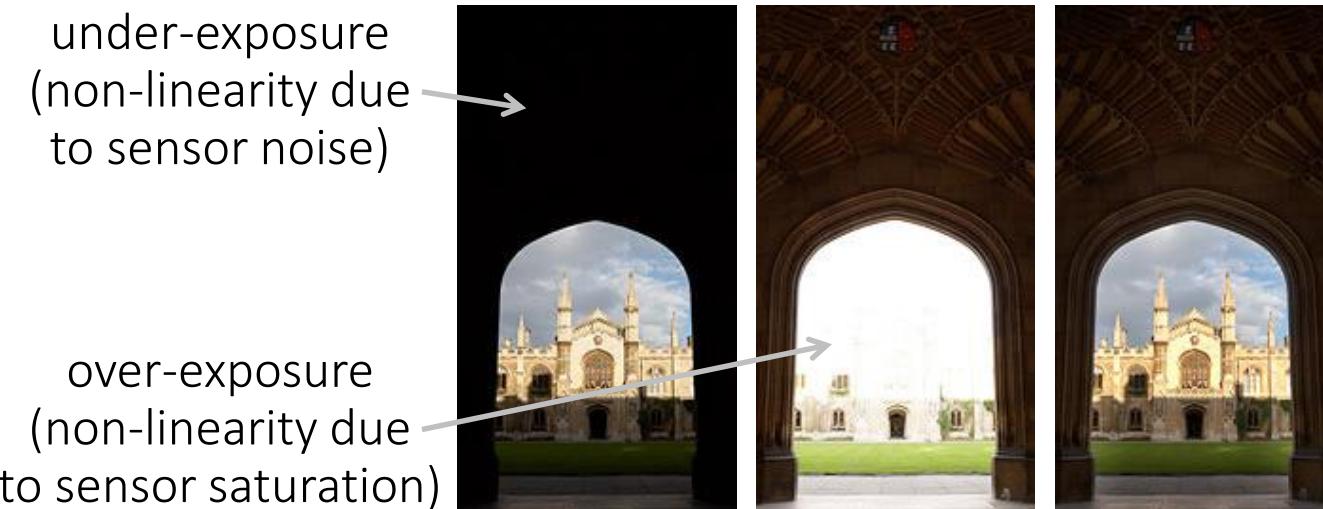
- Fundamental optical performance metric of imaging sensors.
- Not the only important optical performance metric!
- We will see a few more later in the lecture.

# Photodiode response function

For silicon photodiodes, usually linear, but:

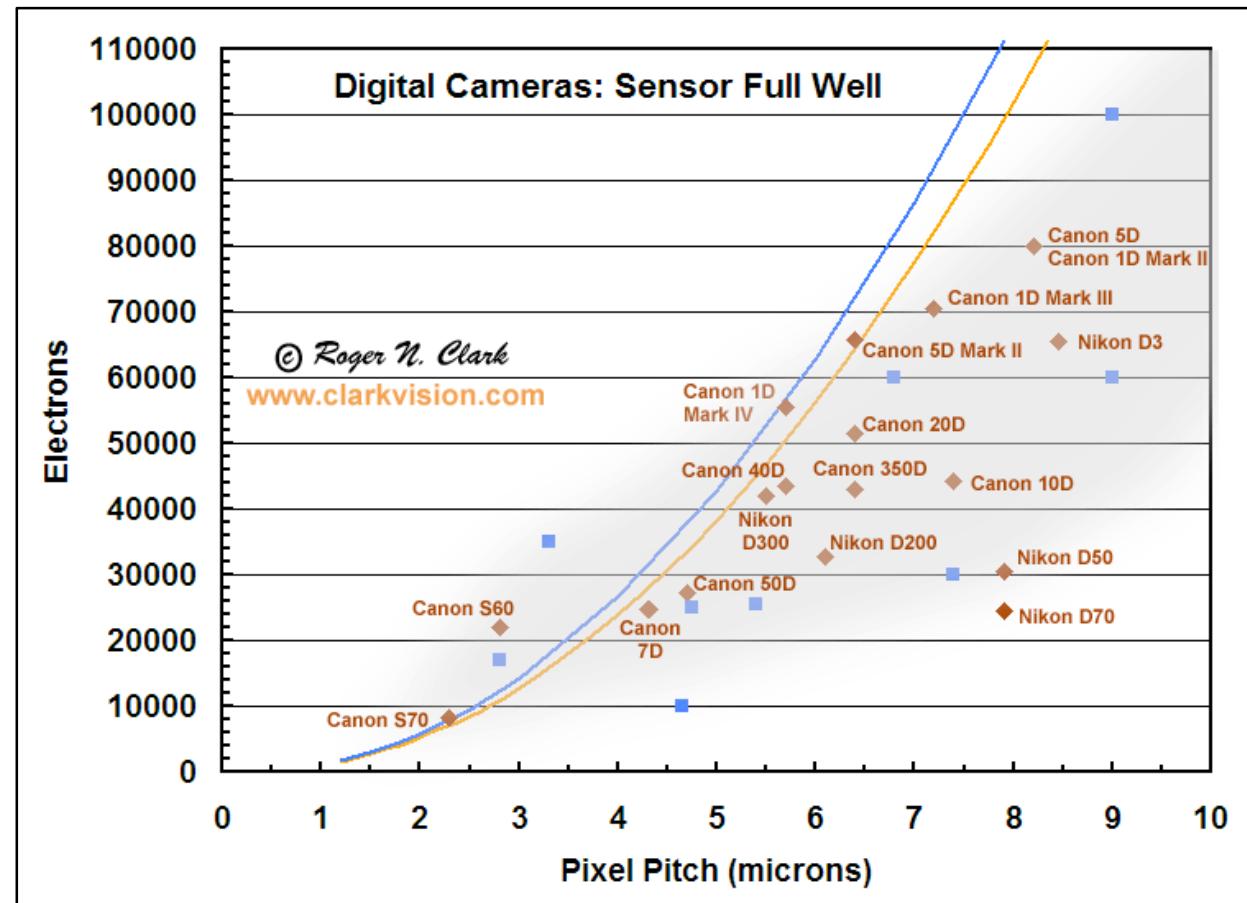
- non-linear when potential well is saturated (over-exposure)
- non-linear near zero (due to noise)

We will see how to deal with these issues in a later lecture (high-dynamic-range imaging).



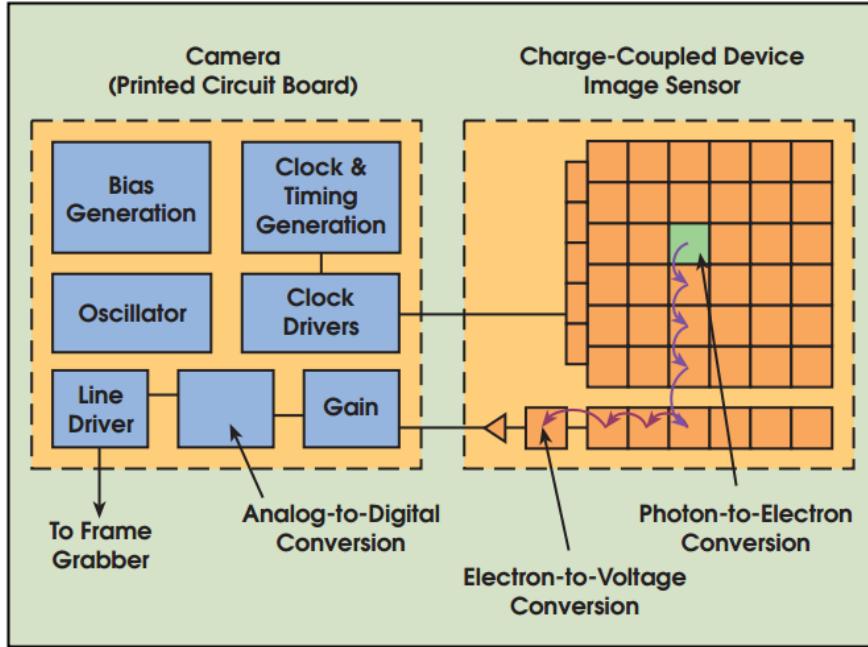
# Photodiode full well capacity

How many electrons can photodiode store before saturation?

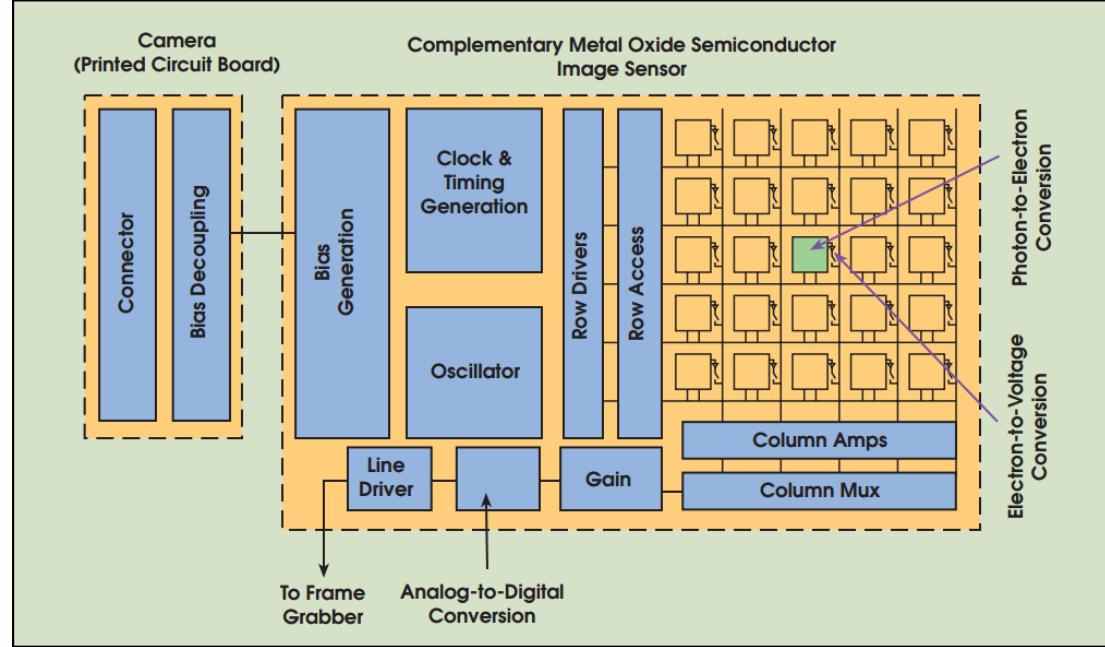


- Another important optical performance metric of imaging sensors.

# Two main types of imaging sensors



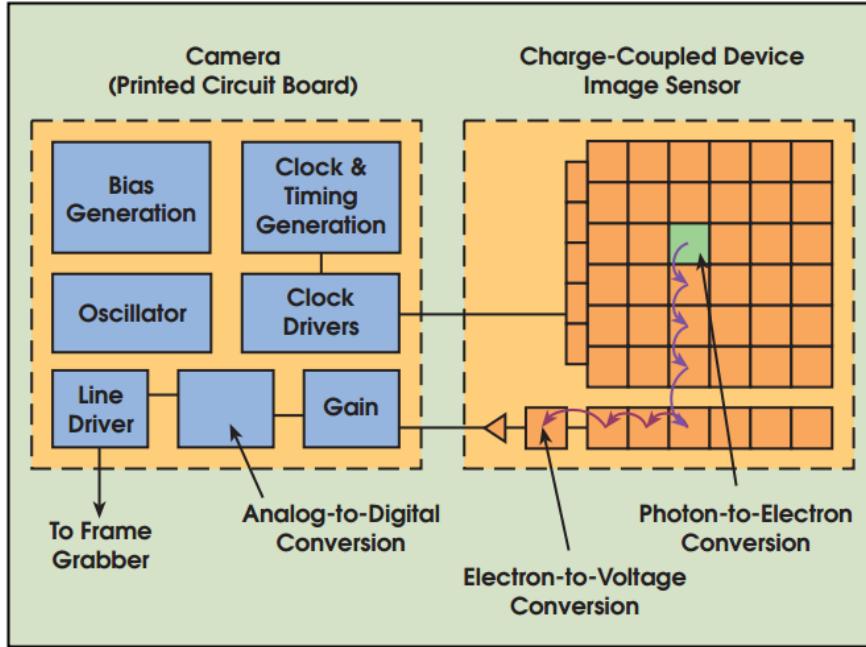
Charged Coupled Device (CCD):  
converts electrons to voltage using  
readout circuitry separate from pixel



Complementary Metal Oxide Semiconductor (CMOS):  
converts electrons to voltage using  
per-pixel readout circuitry

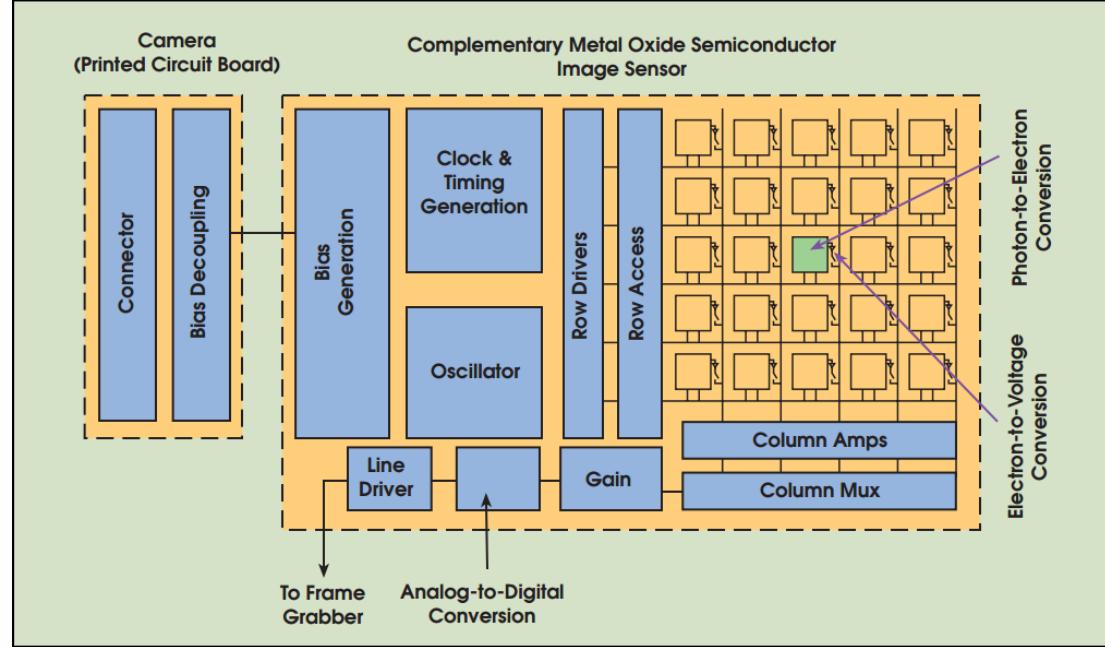
Can you think of advantages and disadvantages of each type?

# Two main types of imaging sensors



Charged Coupled Device (CCD):  
converts electrons to voltage using  
readout circuitry separate from pixel

- ✓ higher sensitivity
- ✓ lower noise



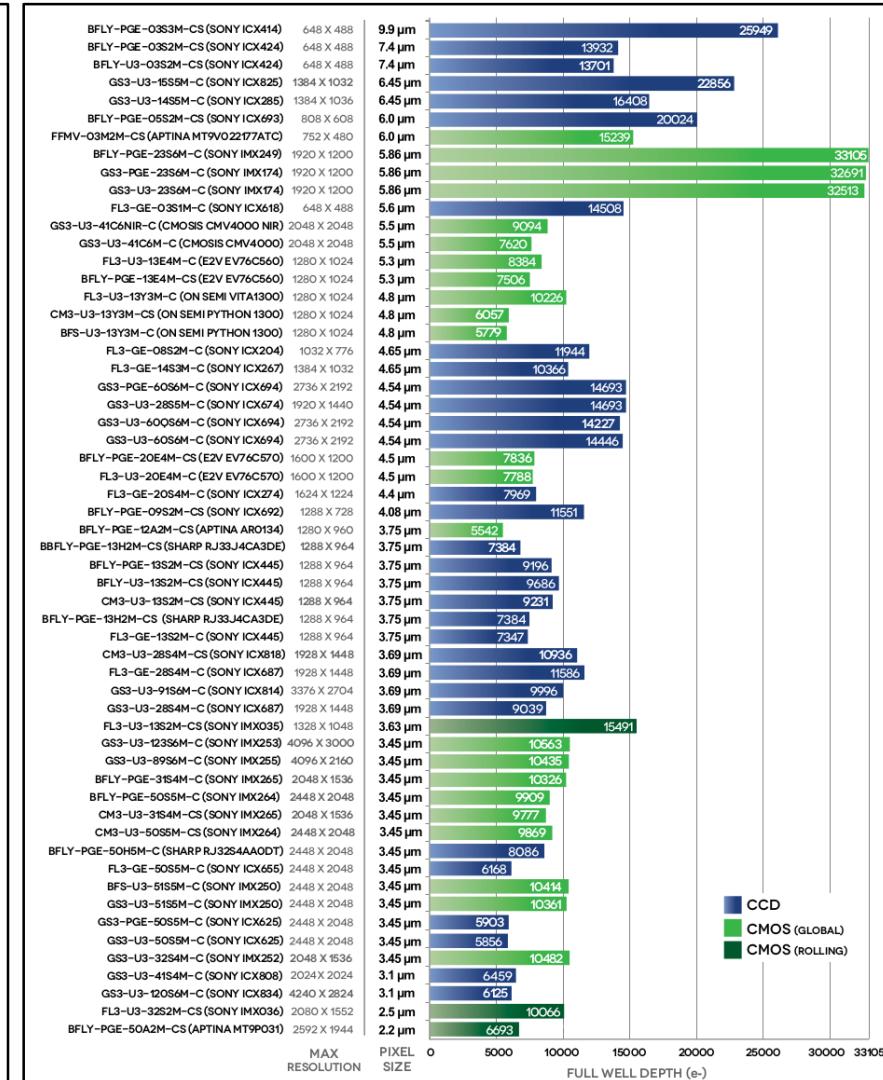
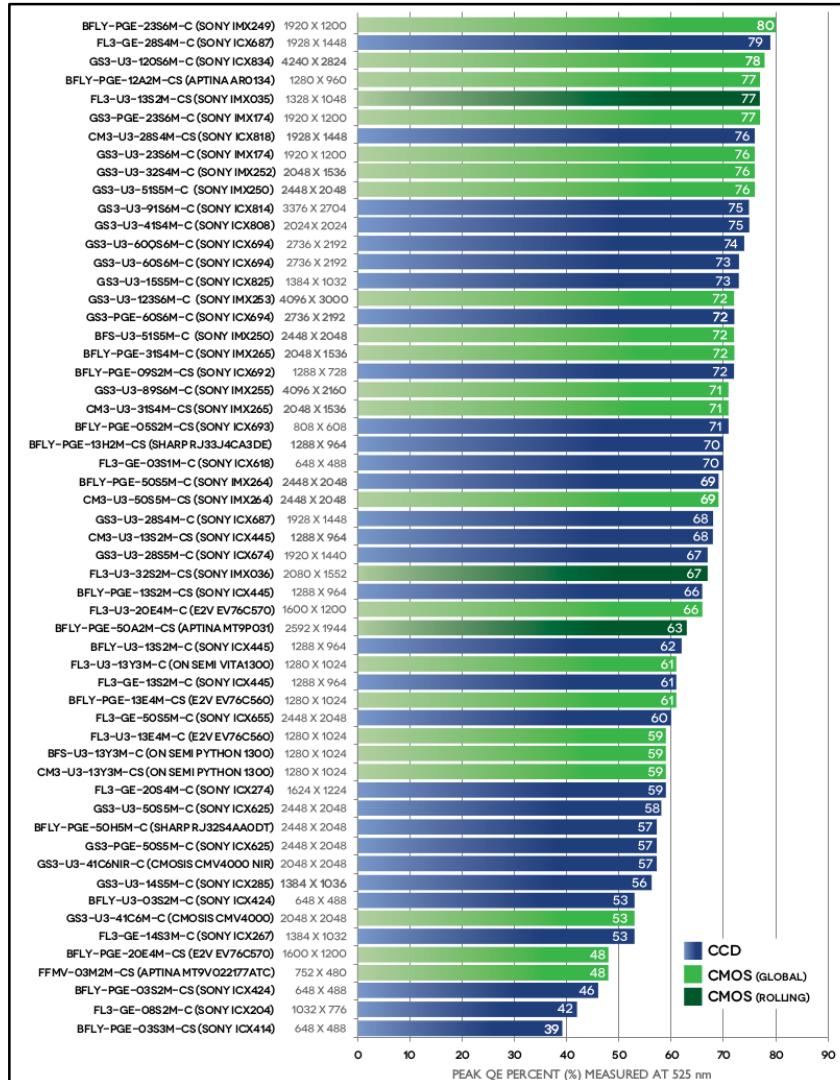
Complementary Metal Oxide Semiconductor (CMOS):  
converts electrons to voltage using  
per-pixel readout circuitry

- ✓ faster read-out
- ✓ lower cost

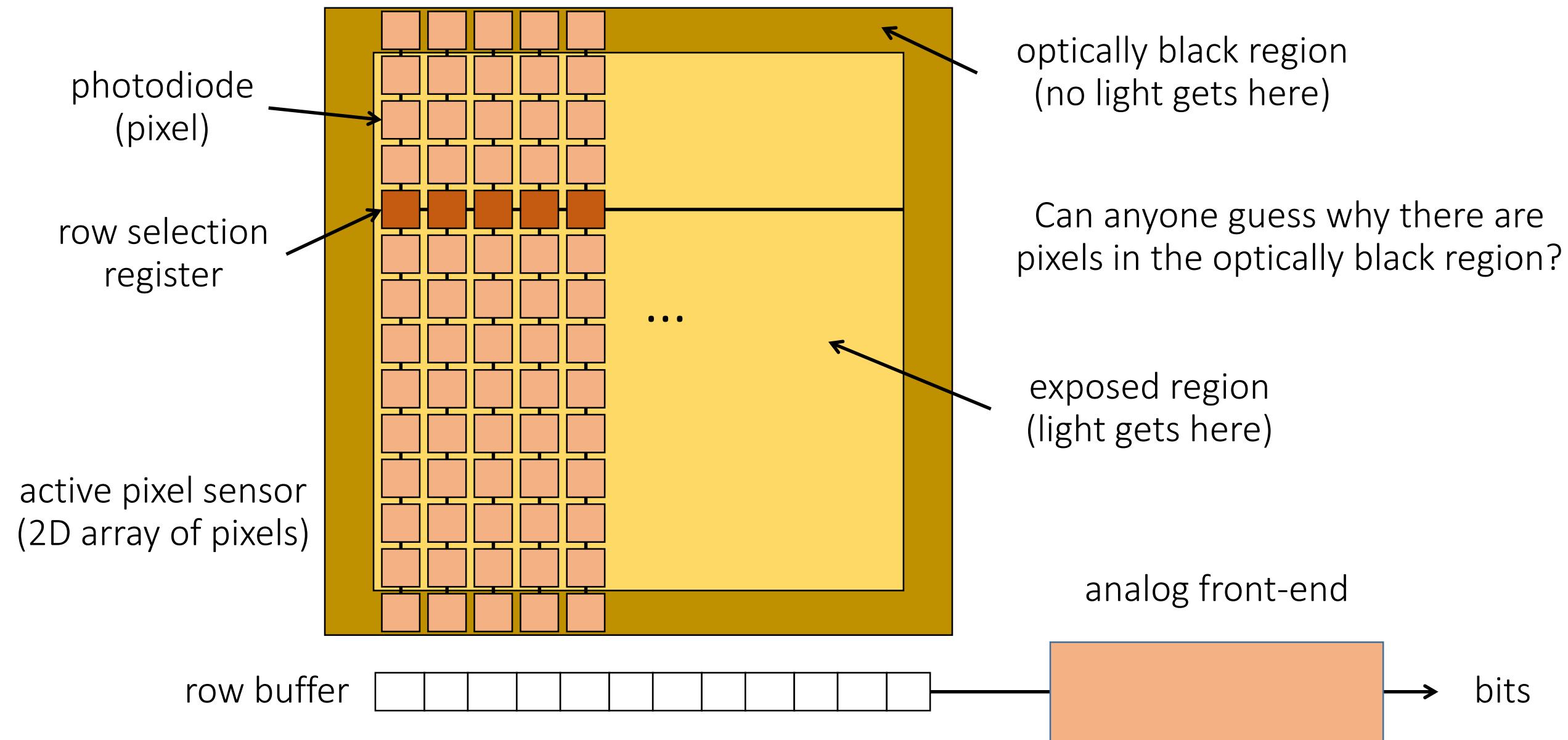
# CCD vs CMOS

- Modern CMOS sensors have optical performance comparable to CCD sensors.
- Most modern commercial and industrial cameras use CMOS sensors.

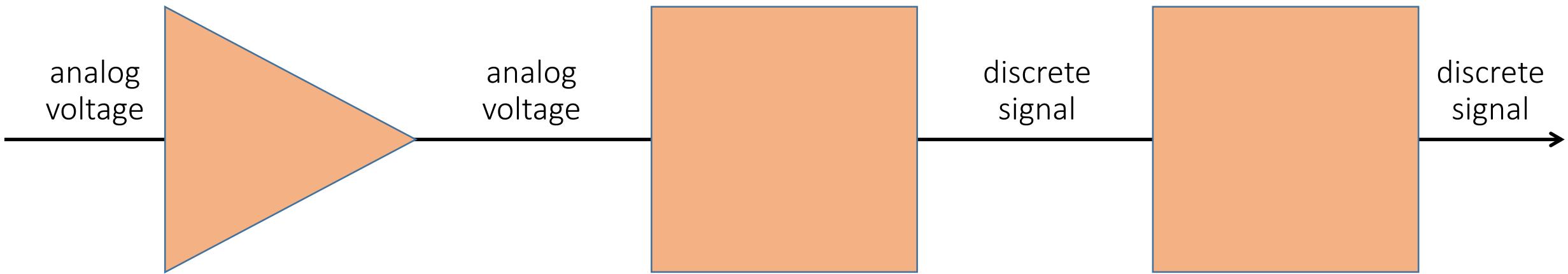
Can you guess what the QE of the human eye is?



# CMOS sensor (very) simplified layout



# Analog front-end



## analog amplifier (gain):

- gets voltage in range needed by A/D converter.
- accommodates ISO settings.
- accounts for vignetting.

## analog-to-digital converter (ADC):

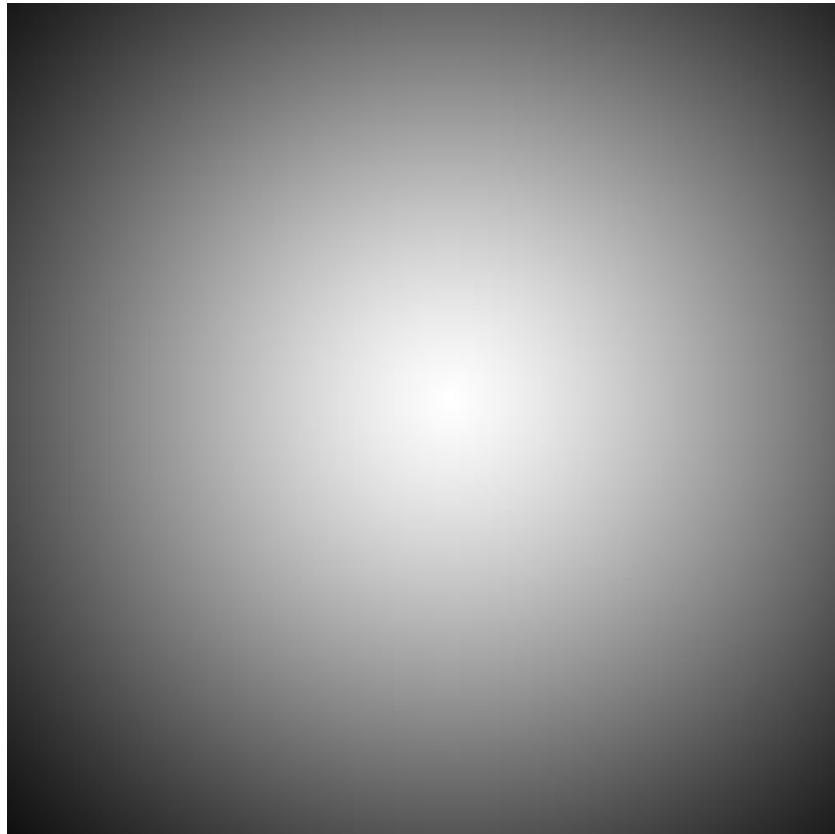
- depending on sensor, output has 10-16 bits.
- most often (?) 12 bits.

## look-up table (LUT):

- corrects non-linearities in sensor's response function (within proper exposure).
- corrects defective pixels.

# Vignetting

Fancy word for: pixels far off the center receive less light



white wall under uniform light

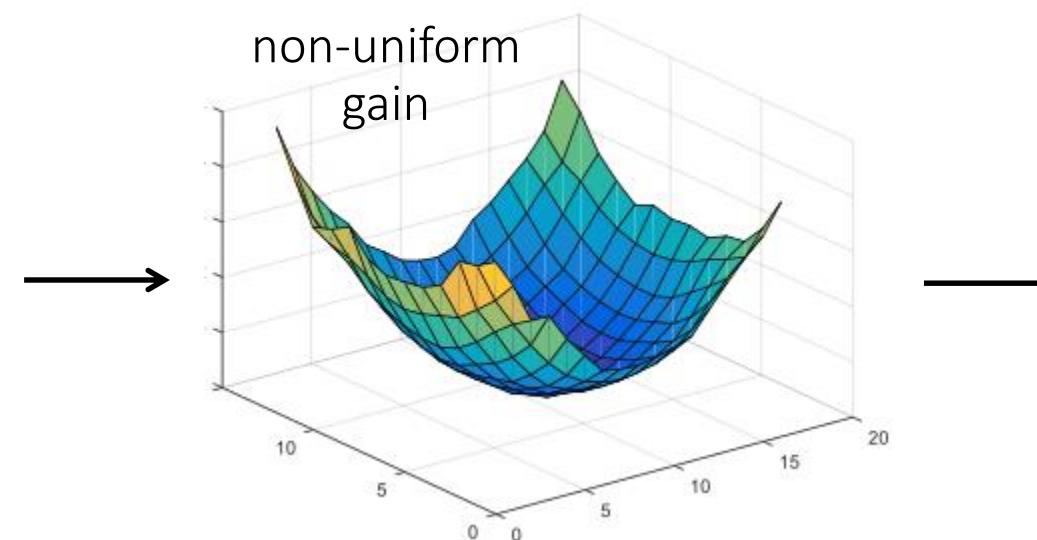
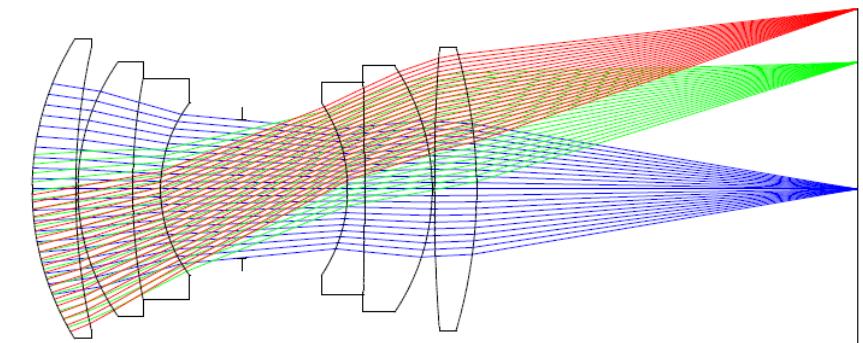


more interesting example of vignetting

# Vignetting

Four types of vignetting:

- Mechanical: light rays blocked by hoods, filters, and other objects.
- Lens: similar, but light rays blocked by lens elements.
- Natural: due to radiometric laws (“cosine fourth falloff”).
- Pixel: angle-dependent sensitivity of photodiodes.



# What does an imaging sensor do?

When the camera shutter opens, the sensor:

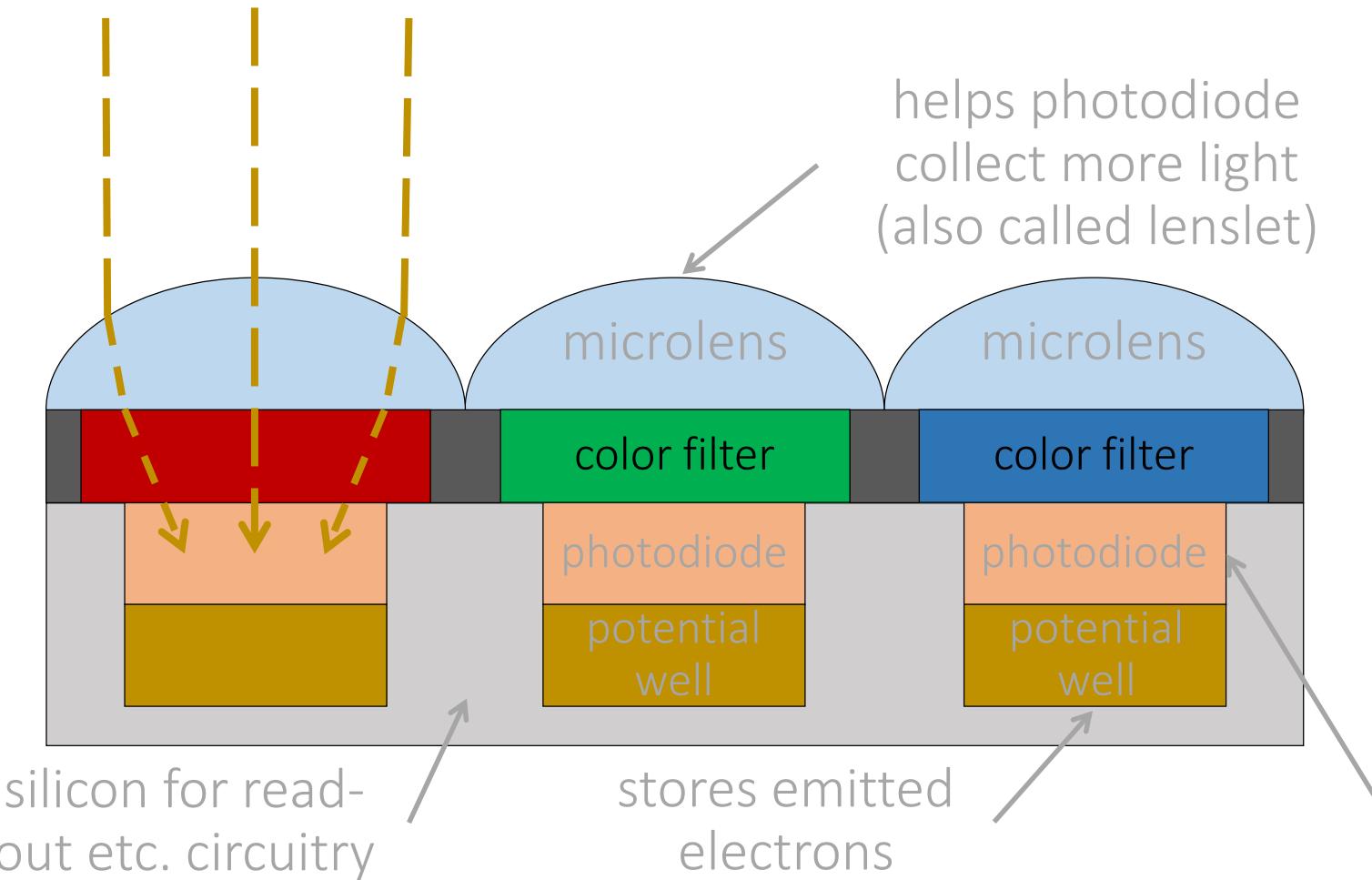
- at every photodiode, converts incident photons into electrons
- stores electrons into the photodiode's potential well until it is full

... until camera shutter closes. Then, the analog front-end:

- reads out photodiodes' wells, row-by-row, and converts them to analog signals
- applies a (possibly non-uniform) gain to these analog signals
- converts them to digital signals
- corrects non-linearities

... and finally returns an image.

# Remember these?



- Lenslets also filter the image to avoid resolution artifacts.
- Lenslets are problematic when working with coherent light.
- Many modern cameras do not have lenslet arrays.

We will discuss these issues in more detail at a later lecture.

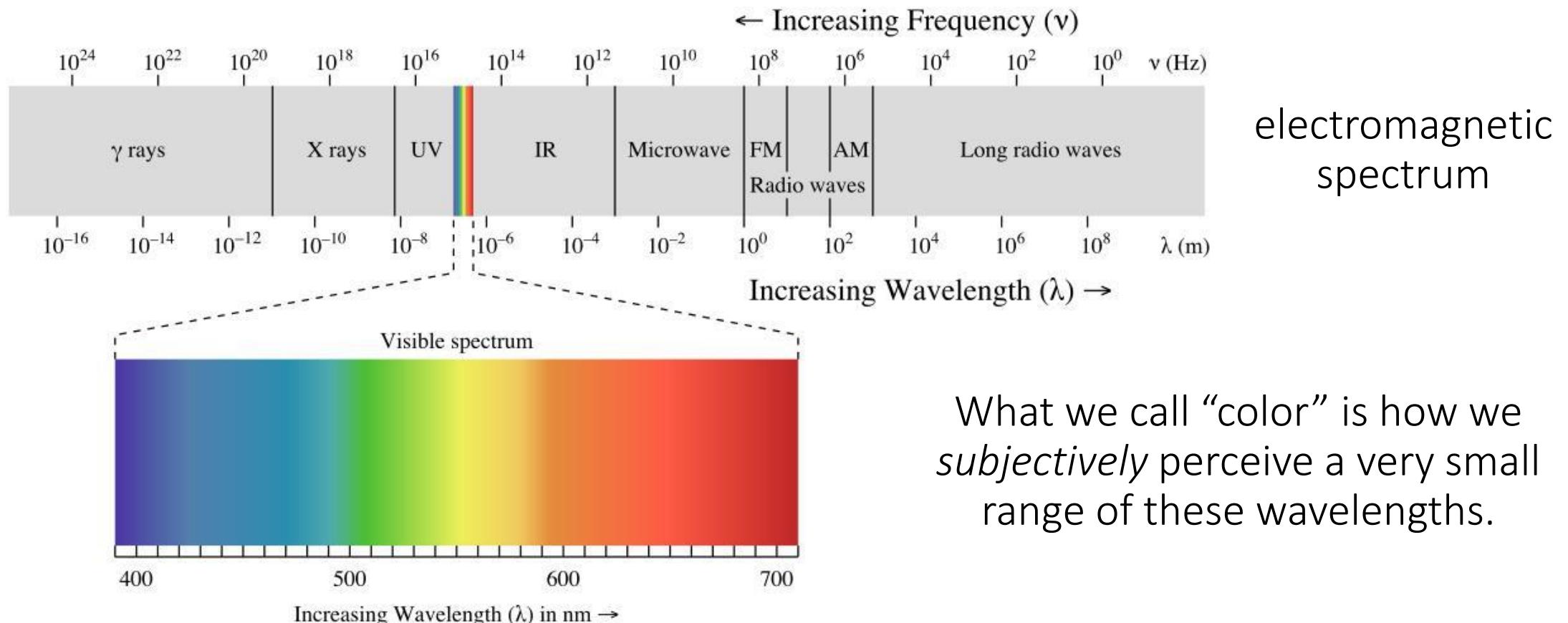
made of silicon, emits electrons from photons

We will see what the color filters are for later in this lecture.

# Color sensing in cameras

# Color is an artifact of human perception

- “Color” is not an *objective* physical property of light (electromagnetic radiation).
- Instead, light is characterized by its wavelength.



What we call “color” is how we *subjectively* perceive a very small range of these wavelengths.

# Spectral Sensitivity Function (SSF)

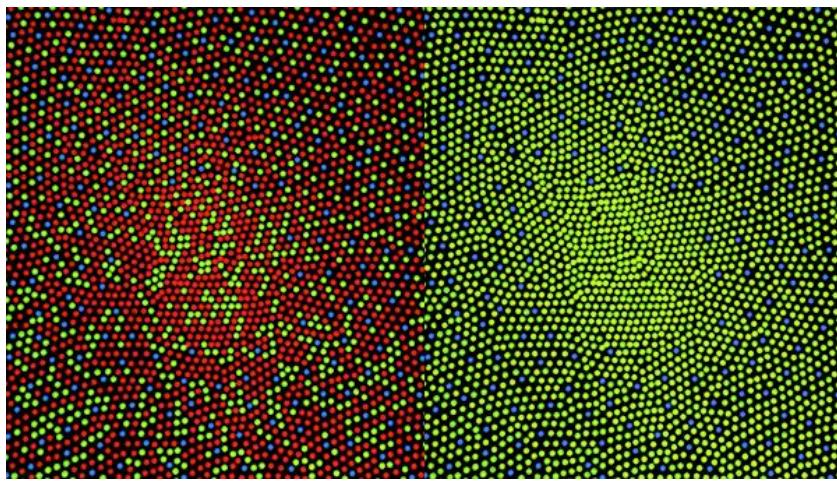
- Any light sensor (digital or not) has different sensitivity to different wavelengths.
- This is described by the sensor's *spectral sensitivity function*  $f(\lambda)$ .
- When measuring light of a some SPD  $\Phi(\lambda)$ , the sensor produces a *scalar* response:

$$\text{sensor response} \longrightarrow R = \int_{\lambda} \text{light SPD} \quad \text{sensor SSF} \quad \Phi(\lambda) f(\lambda) d\lambda$$

Weighted combination of light's SPD: light contributes more at wavelengths where the sensor has higher sensitivity.

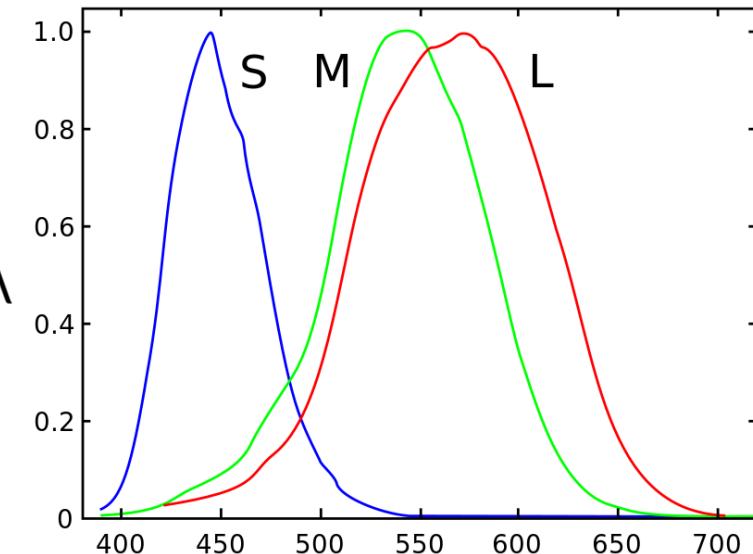
# Spectral Sensitivity Function of Human Eye

- The human eye is a collection of light sensors called cone cells.
- There are three types of cells with different spectral sensitivity functions.
- Human color perception is three-dimensional (*tristimulus color*).



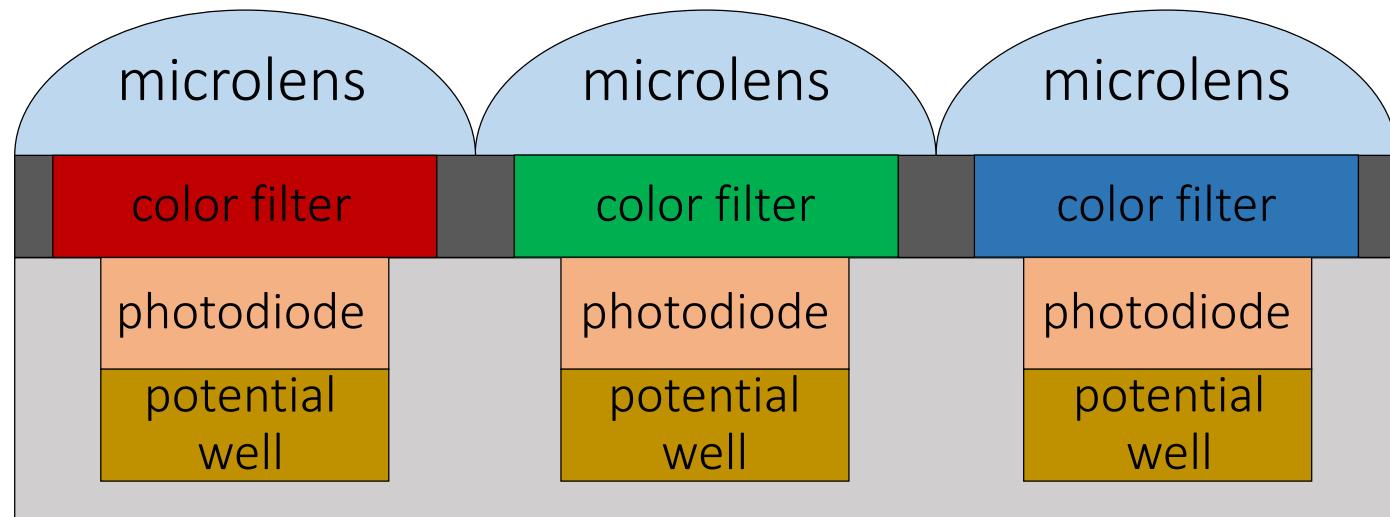
cone distribution  
for normal vision  
(64% L, 32% M)

$$\begin{aligned} \text{"short"} \quad S &= \int_{\lambda} \Phi(\lambda) S(\lambda) d\lambda \\ \text{"medium"} \quad M &= \int_{\lambda} \Phi(\lambda) M(\lambda) d\lambda \\ \text{"long"} \quad L &= \int_{\lambda} \Phi(\lambda) L(\lambda) d\lambda \end{aligned}$$



# Color filter arrays (CFA)

- To measure color with a digital sensor, mimic cone cells of human vision system.
- “Cones” correspond to pixels that are covered by different color filters, each with its own spectral sensitivity function.

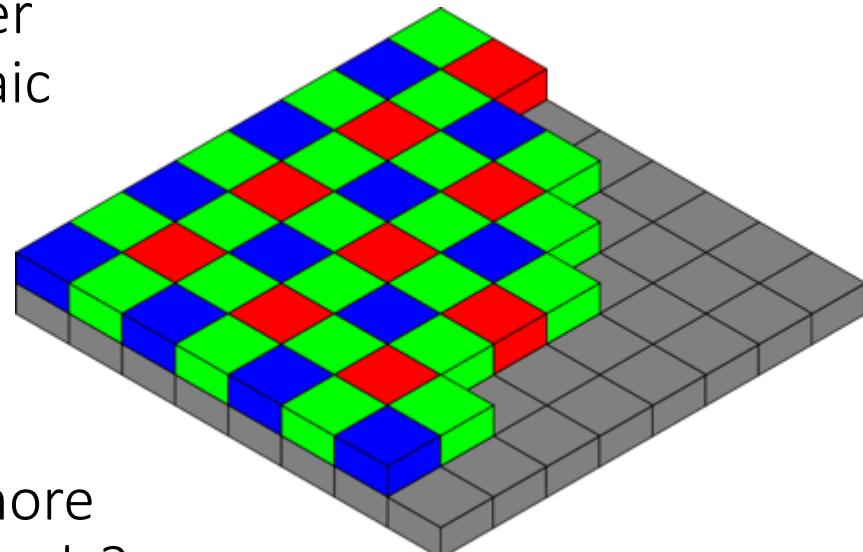


# What color filters to use?

Two design choices:

- What spectral sensitivity functions  $f(\lambda)$  to use for each color filter?
- How to spatially arrange (“mosaic”) different color filters?

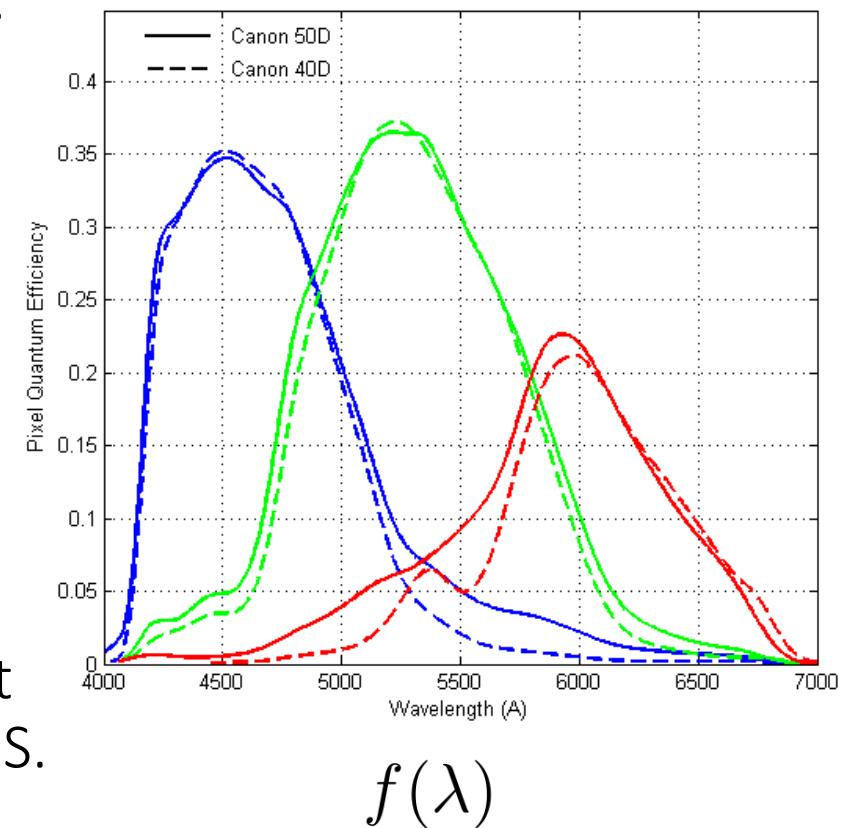
Bayer  
mosaic



Why more  
green pixels?

SSF for  
Canon 50D

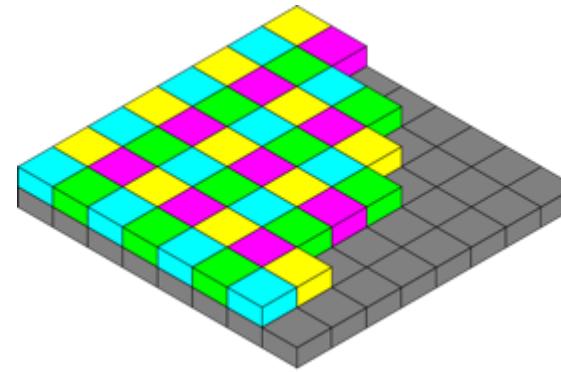
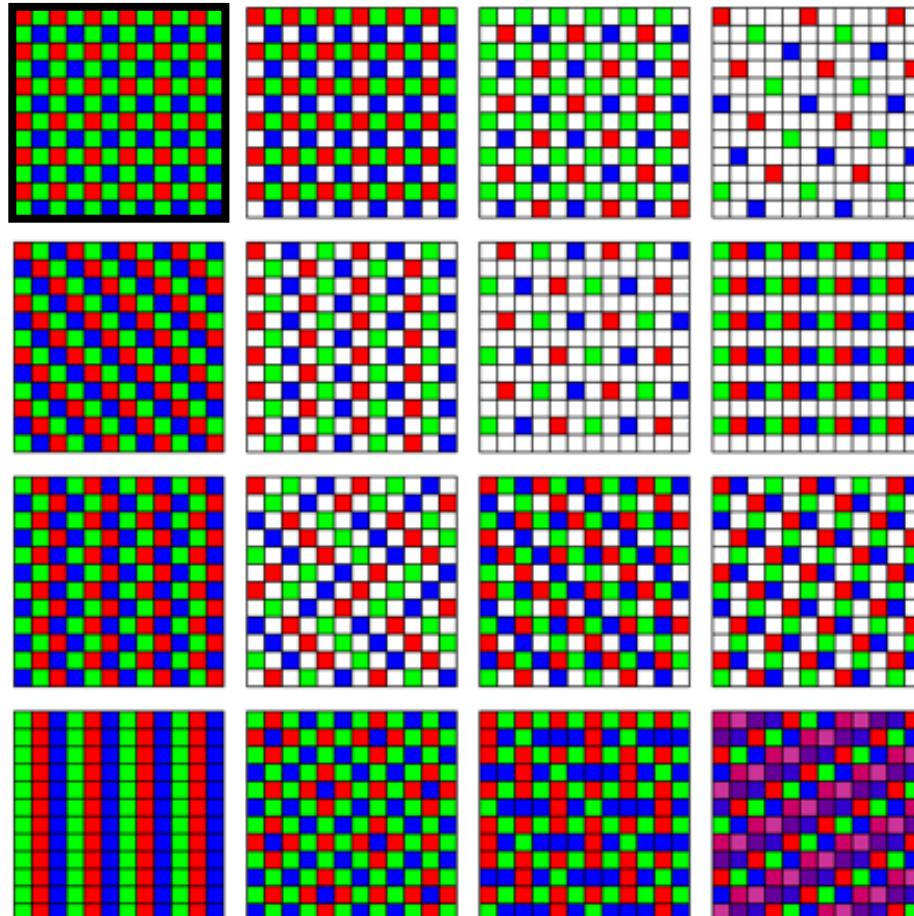
Generally do not  
match human LMS.



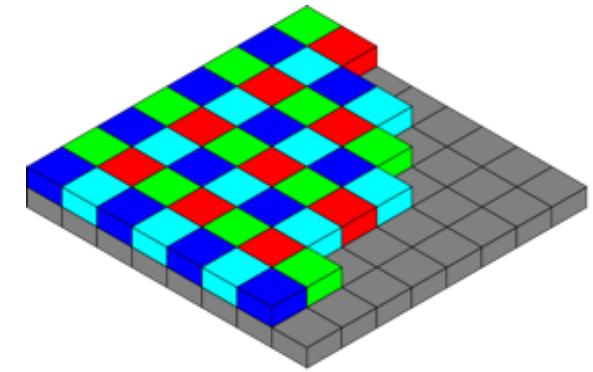
$$f(\lambda)$$

# Many different CFAs

Finding the “best” CFA mosaic is an active research area.



CYGM  
Canon IXUS, Powershot



RGBE  
Sony Cyber-shot

How would you go about designing your own CFA? What criteria would you consider?

# Many different spectral sensitivity functions

Each camera has its more or less unique, and most of the time *secret*, SSF.

- Makes it very difficult to correctly reproduce the color of sensor measurements.



Images of the same scene captured using 3 different cameras with identical sRGB settings.

Aside: can you think of other ways to capture color?

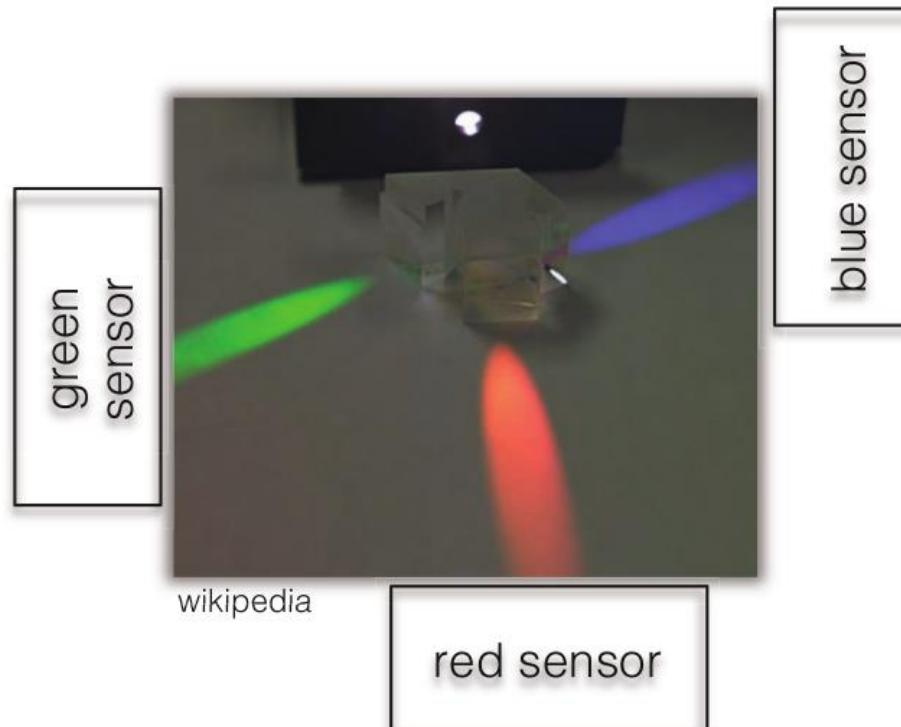
# Aside: can you think of other ways to capture color?

field sequential

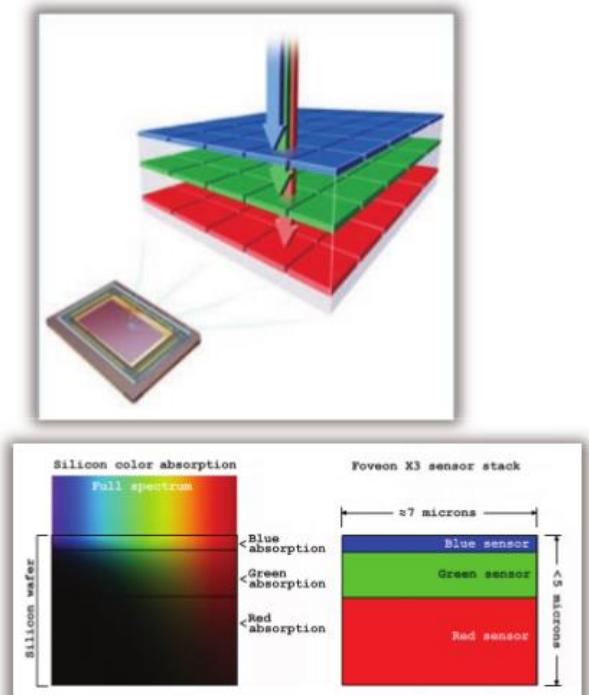


Prokudin-Gorsky

multiple sensors



vertically stacked



[Slide credit: Gordon Wetzstein]

# What does an imaging sensor do?

When the camera shutter opens, the sensor:

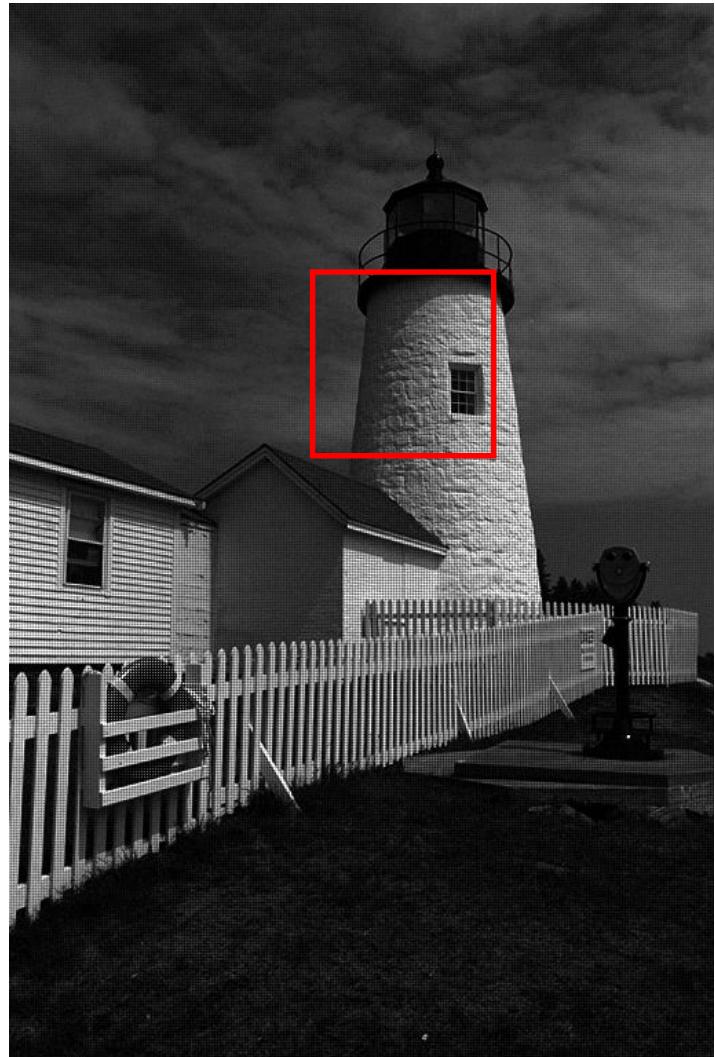
- at every photodiode, converts incident photons into electrons using mosaic's SSF
- stores electrons into the photodiode's potential well until it is full

... until camera shutter closes. Then, the analog front-end:

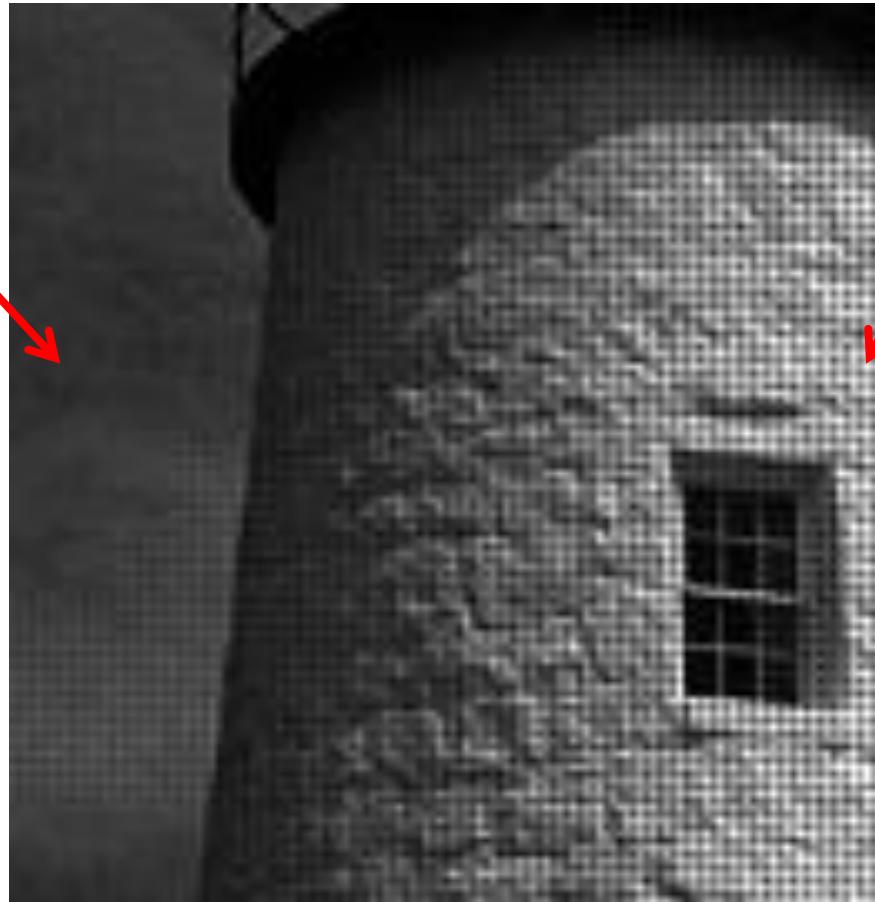
- reads out photodiodes' wells, row-by-row, and converts them to analog signals
- applies a (possibly non-uniform) gain to these analog signals
- converts them to digital signals
- corrects non-linearities

... and finally returns an image.

# After all of this, what does an image look like?



lots of  
noise



mosaicking  
artifacts

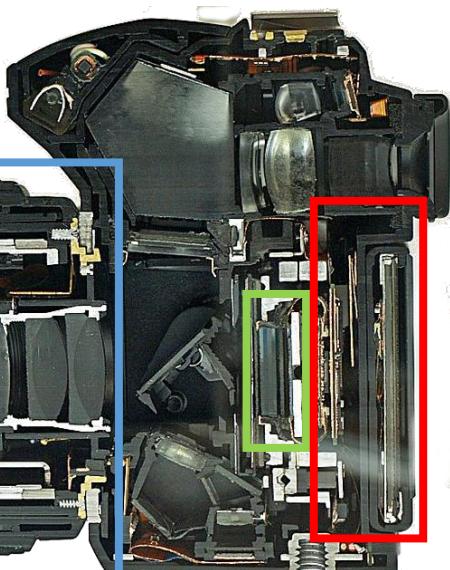
- Kind of disappointing.
- We call this the *RAW* image.



# The modern photography pipeline



post-capture processing  
(lectures 3-12)



optics and  
optical controls

(lectures 13-16)

sensor, analog  
front-end, and  
color filter array

(this lecture)

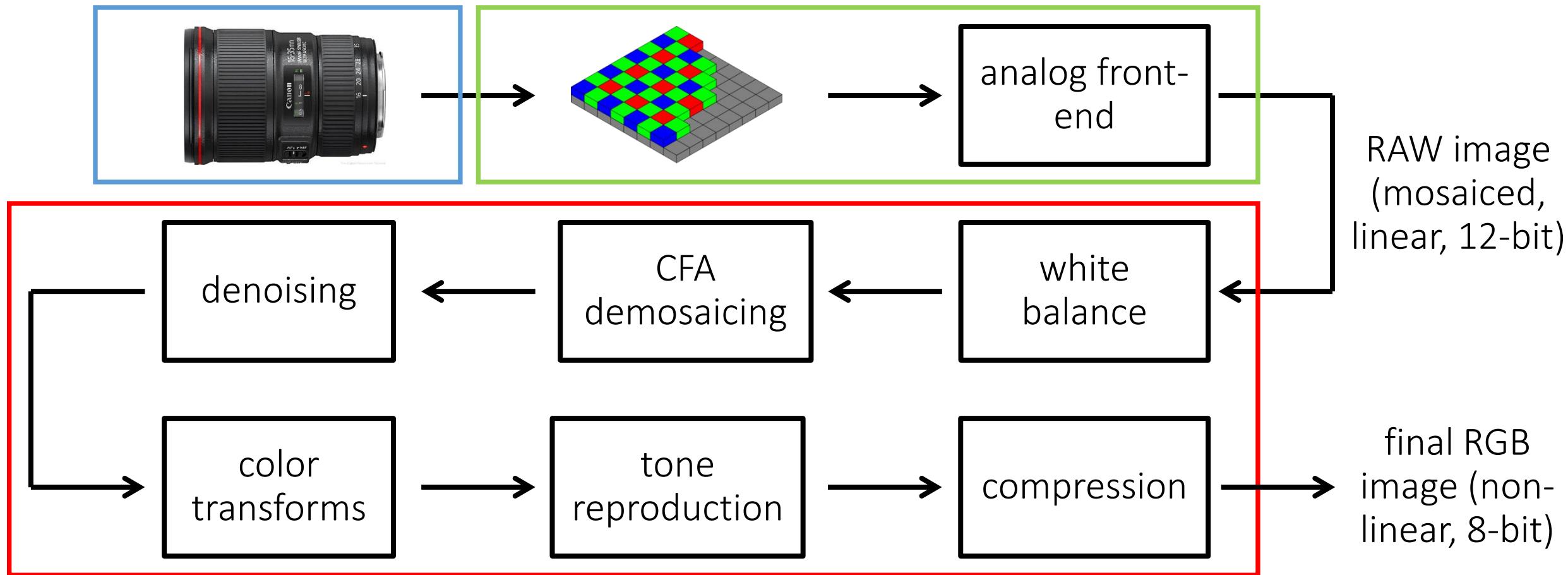
in-camera image  
processing  
pipeline

(this lecture)

The in-camera image processing pipeline

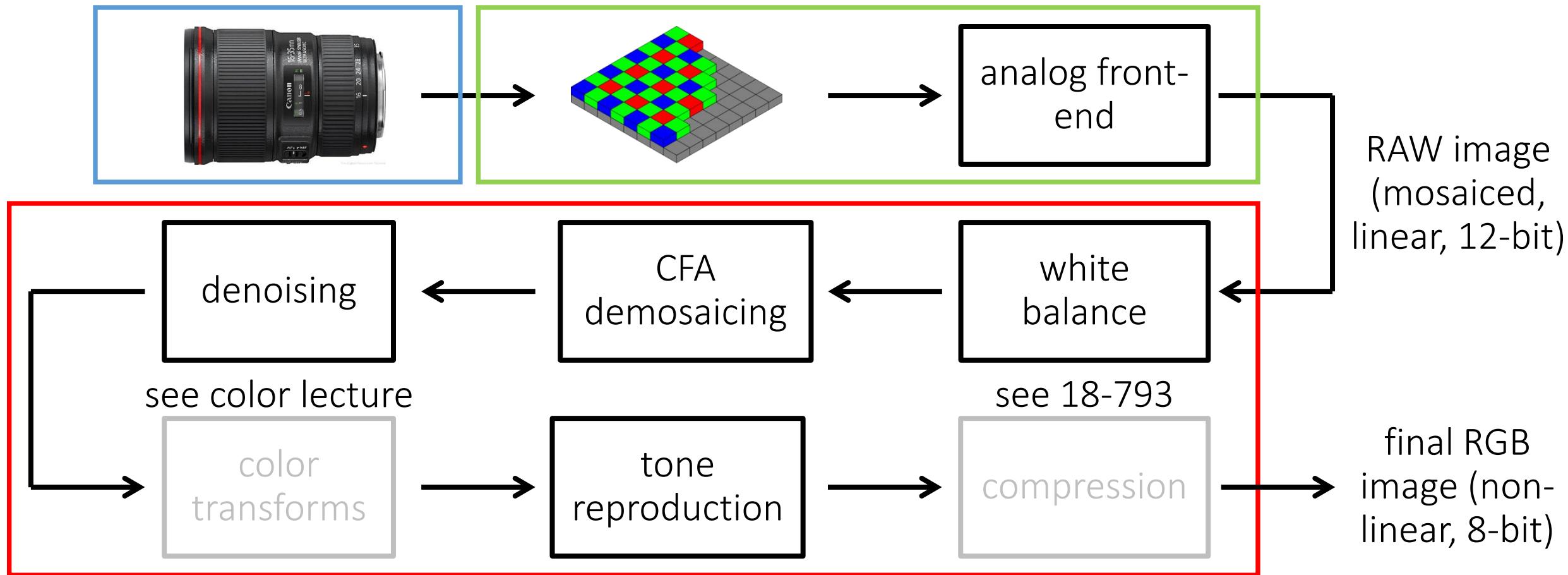
# The (in-camera) image processing pipeline

The sequence of image processing operations applied by the camera's image signal processor (ISP) to convert a RAW image into a "conventional" image.



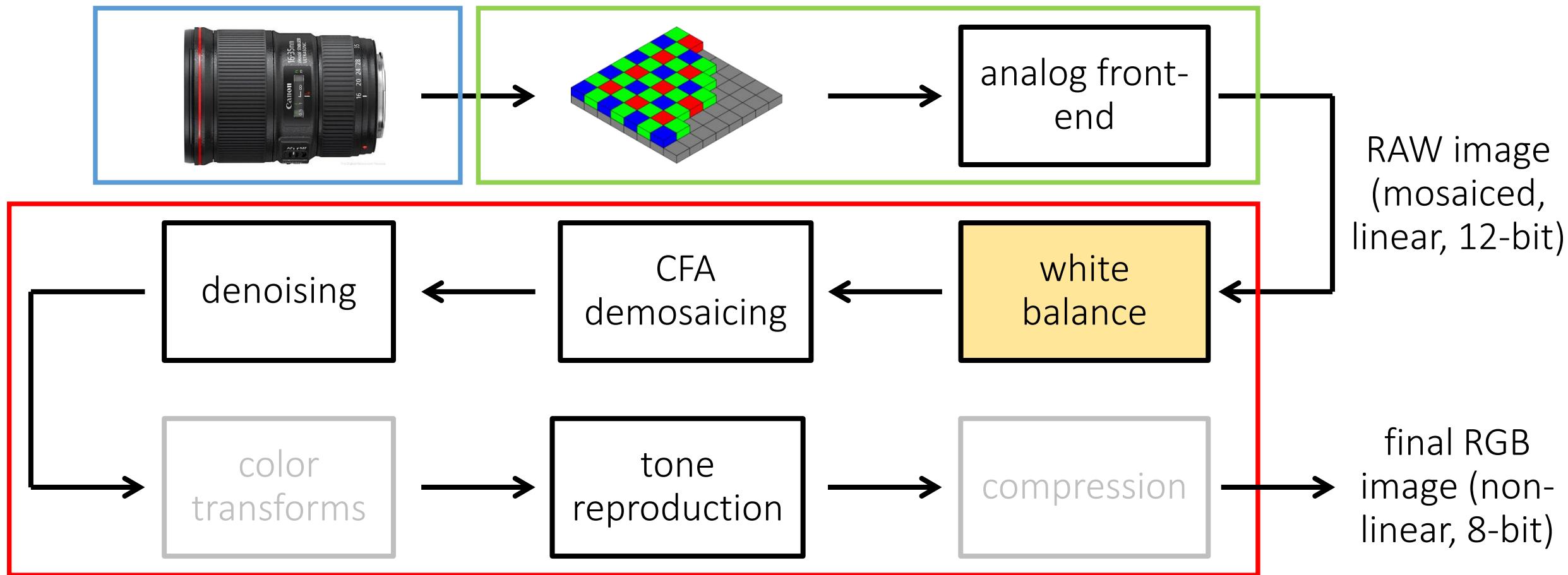
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The sequence of image processing operations applied by the camera's image signal processor (ISP) to convert a RAW image into a "conventional" image.

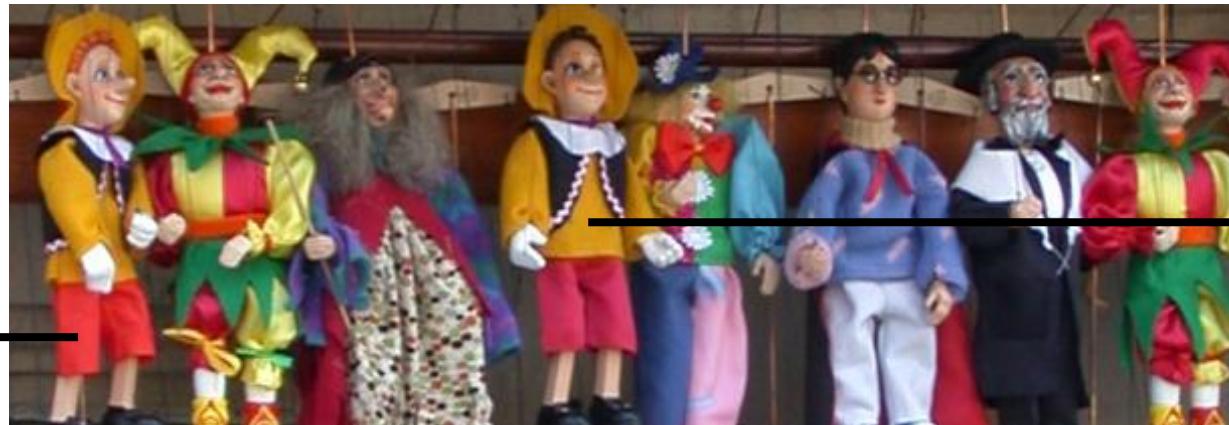


# White balancing

Human visual system has *chromatic adaptation*:

- We can perceive white (and other colors) correctly under different light sources.

Retinal vs  
perceived color.



# White balancing

Human visual system has *chromatic adaptation*:

- We can perceive white (and other colors) correctly under different light sources.
- Cameras cannot do that (there is no “camera perception”).

White balancing: The process of removing color casts so that colors that we would *perceive* as white are *rendered* as white in final image.



different whites



image captured  
under fluorescent



image white-  
balanced to daylight

# White balancing presets

Cameras nowadays come with a large number of presets: You can select which light you are taking images under, and the appropriate white balancing is applied.

WB SETTINGS	COLOR TEMPERATURE	LIGHT SOURCES
	10000 - 15000 K	Clear Blue Sky
	6500 - 8000 K	Cloudy Sky / Shade
	6000 - 7000 K	Noon Sunlight
	5500 - 6500 K	Average Daylight
	5000 - 5500 K	Electronic Flash
	4000 - 5000 K	Fluorescent Light
	3000 - 4000 K	Early AM / Late PM
	2500 - 3000 K	Domestic Lightning
	1000 - 2000 K	Candle Flame

# Manual vs automatic white balancing

Manual white balancing:

- Manually select object in photograph that is color-neutral and use it to normalize.
- Select a camera preset based on lighting.



How can we do automatic white balancing?

# Manual vs automatic white balancing

Manual white balancing:

- Manually select object in photograph that is color-neutral and use it to normalize.
- Select a camera preset based on lighting.



Automatic white balancing:

- Grey world assumption: force average color of scene to be grey.
- White world assumption: force brightest object in scene to be white.
- Sophisticated histogram-based algorithms (what most modern cameras do).

# Automatic white balancing

Grey world assumption:

- Compute per-channel average.
- Normalize each channel by its average.
- Normalize by green channel average.

$$\begin{array}{l} \text{white-balanced} \\ \text{RGB} \end{array} \rightarrow \begin{bmatrix} R' \\ G' \\ B' \end{bmatrix} = \begin{bmatrix} G_{avg}/R_{avg} & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & G_{avg}/B_{avg} \end{bmatrix} \begin{bmatrix} R \\ G \\ B \end{bmatrix} \quad \leftarrow \text{sensor RGB}$$

White world assumption:

- Compute per-channel maximum.
- Normalize each channel by its maximum.
- Normalize by green channel maximum.

$$\begin{array}{l} \text{white-balanced} \\ \text{RGB} \end{array} \rightarrow \begin{bmatrix} R' \\ G' \\ B' \end{bmatrix} = \begin{bmatrix} G_{max}/R_{max} & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & G_{max}/B_{max} \end{bmatrix} \begin{bmatrix} R \\ G \\ B \end{bmatrix} \quad \leftarrow \text{sensor RGB}$$

# Automatic white balancing example



input image



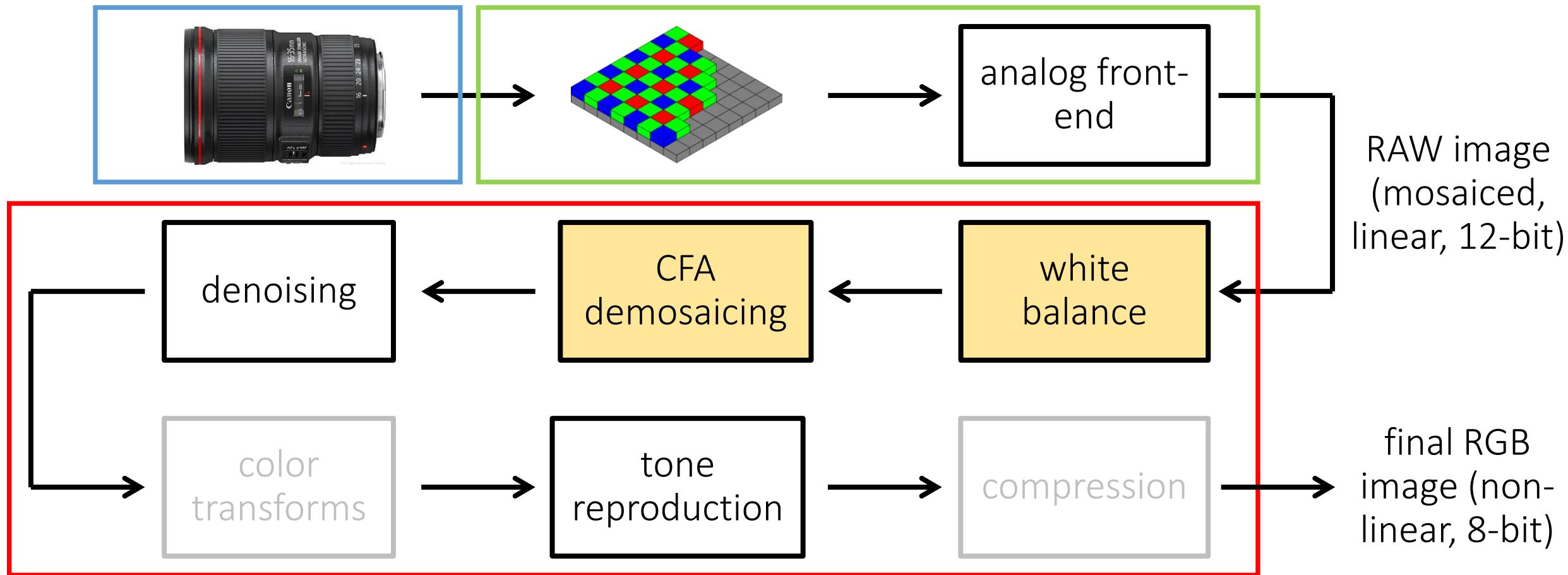
grey world



white world

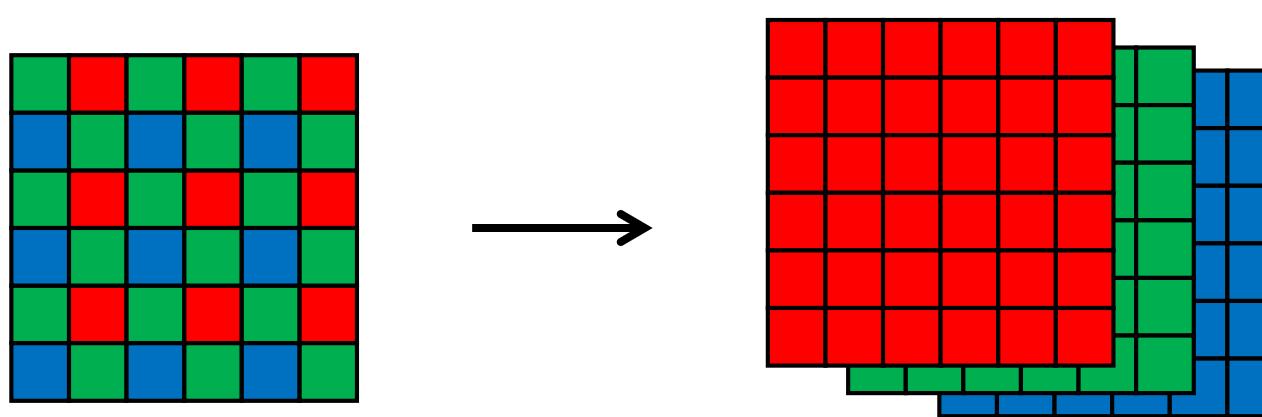
# The (in-camera) image processing pipeline

The sequence of image processing operations applied by the camera's image signal processor (ISP) to convert a RAW image into a "conventional" image.



# CFA demosaicing

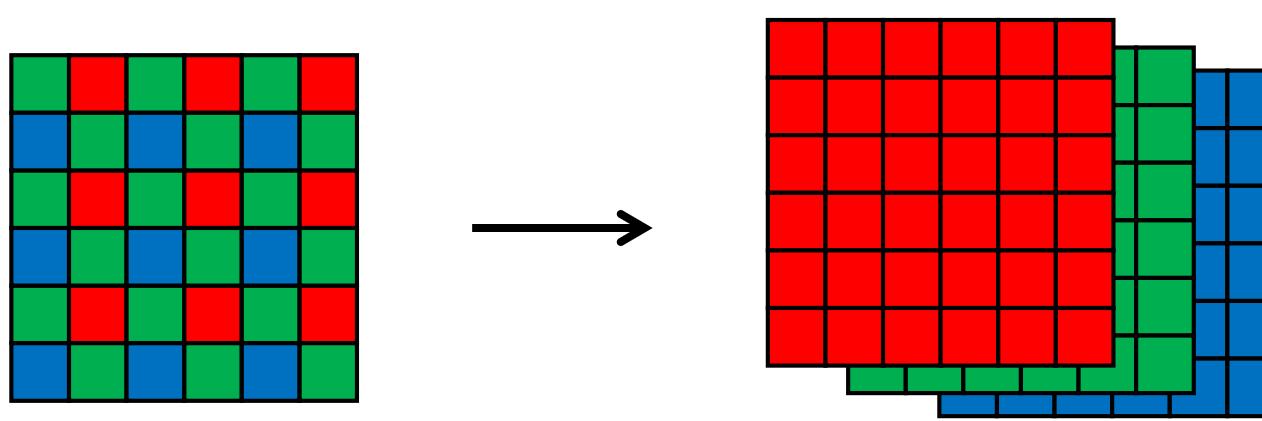
Produce full RGB image from mosaiced sensor output.



Any ideas on how to do this?

# CFA demosaicing

Produce full RGB image from mosaiced sensor output.



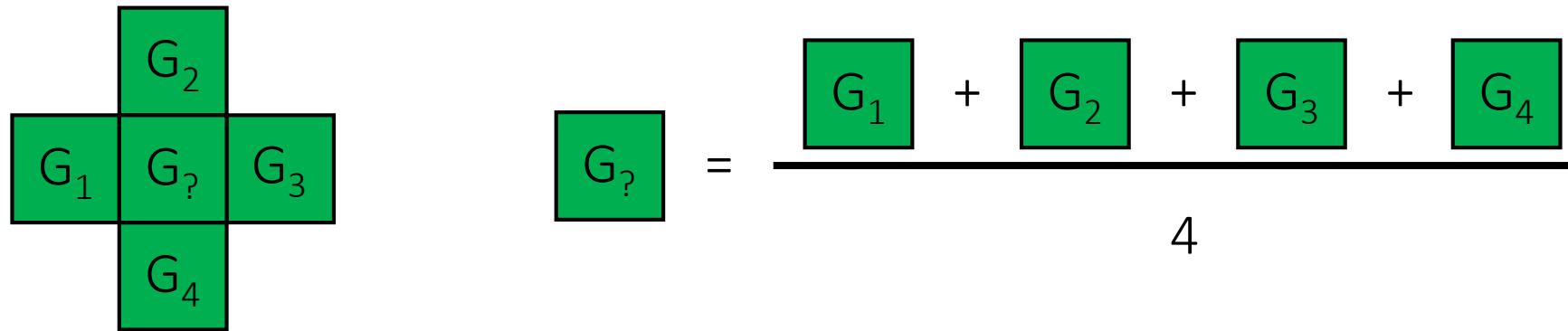
Interpolate from neighbors:

- Bilinear interpolation (needs 4 neighbors).
- Bicubic interpolation (needs more neighbors, may overblur).
- Edge-aware interpolation.

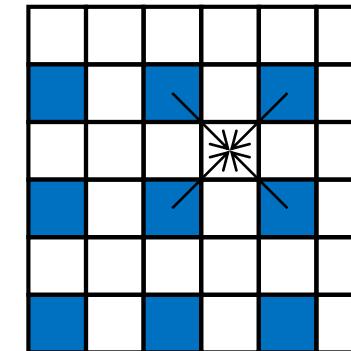
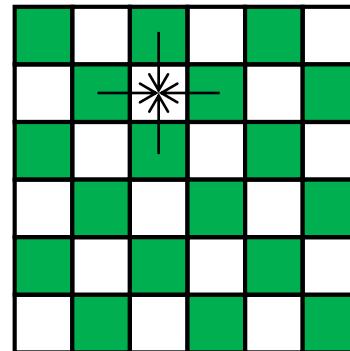
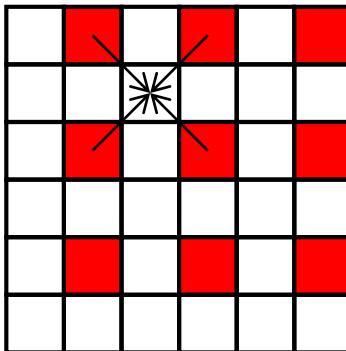
Large area of research.

# Demosaicing by bilinear interpolation

Bilinear interpolation: Simply average your 4 neighbors.

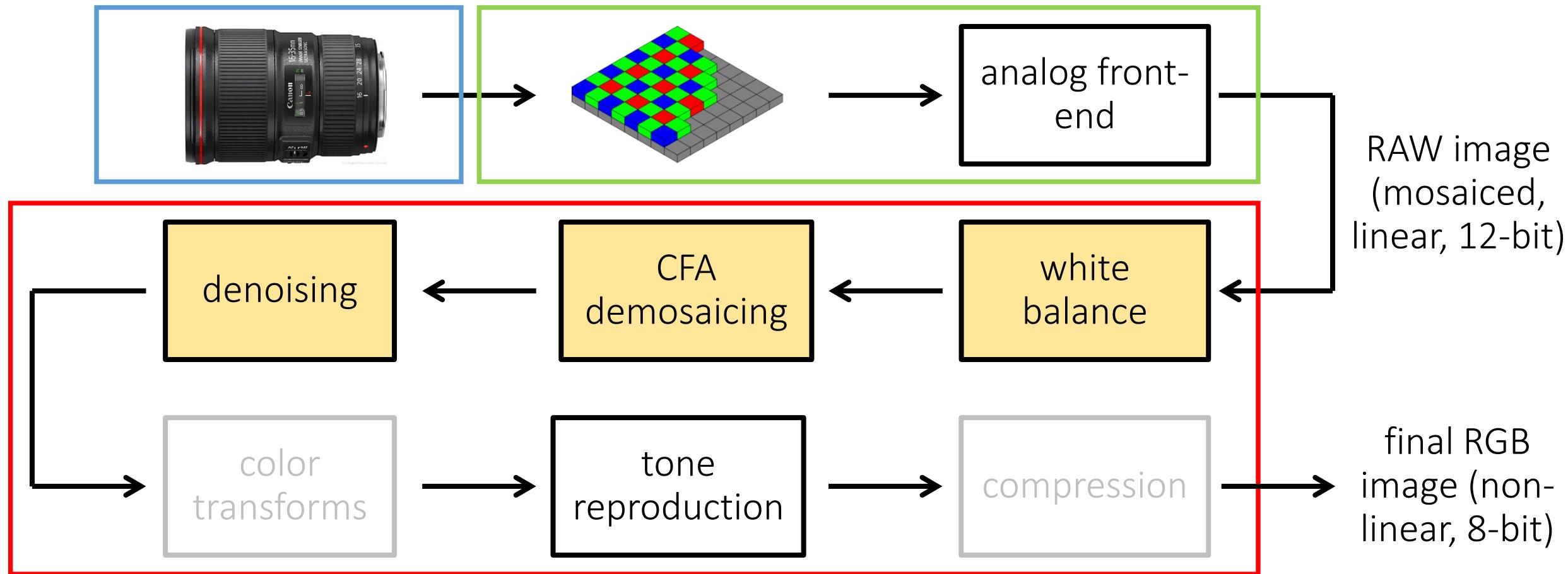


Neighborhood changes for different channels:



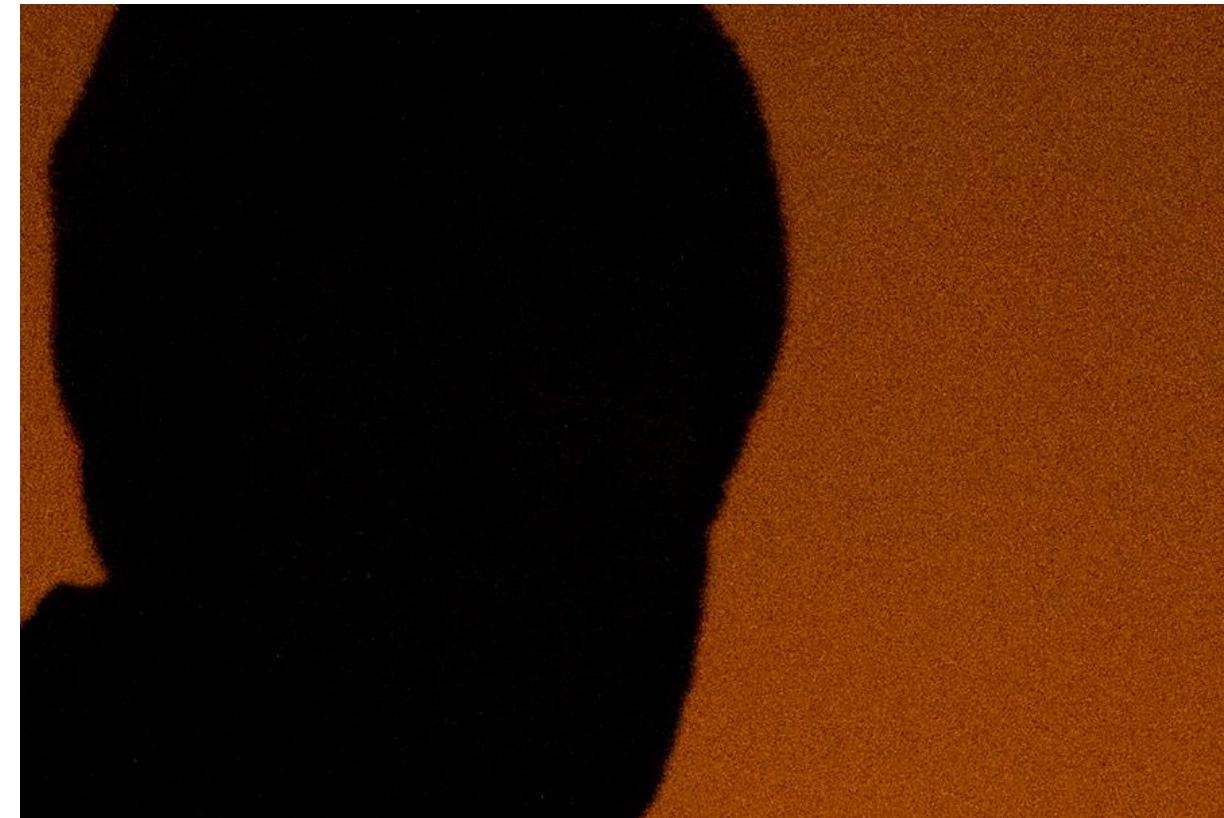
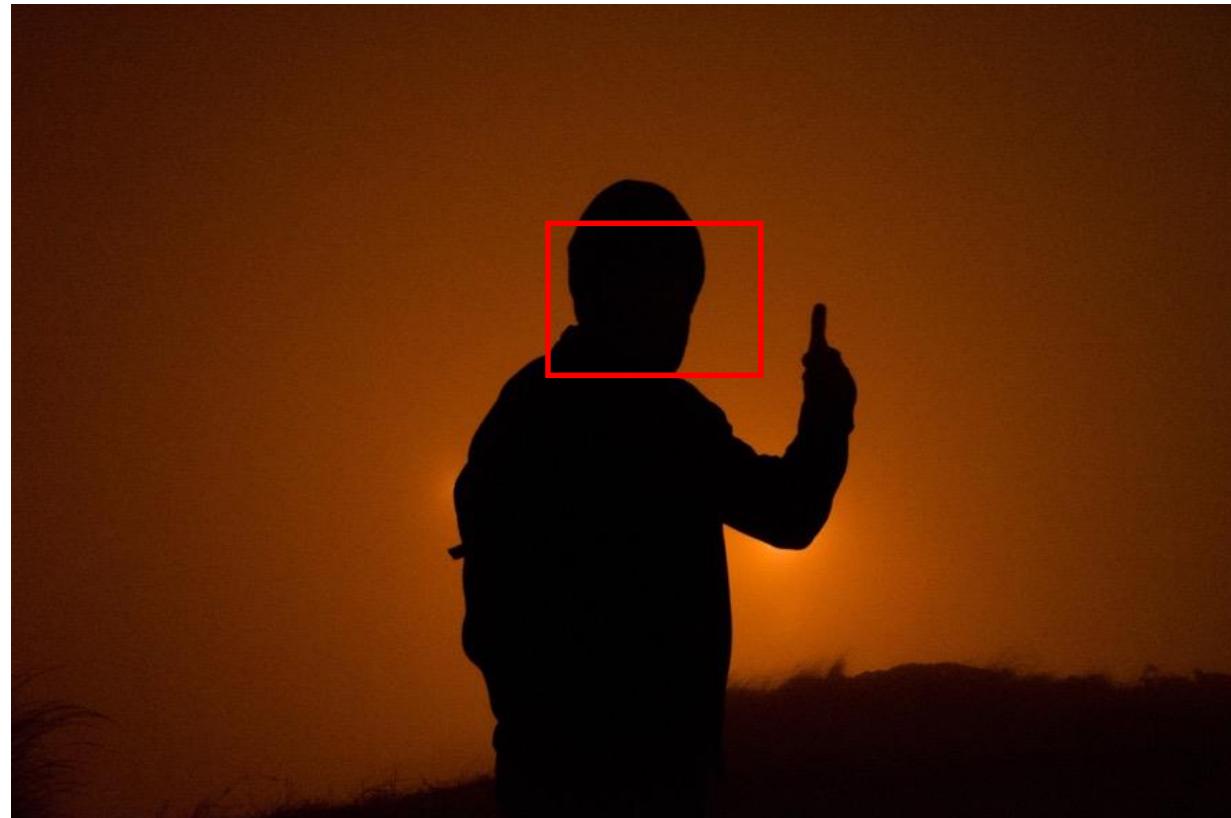
# The (in-camera) image processing pipeline

The sequence of image processing operations applied by the camera's image signal processor (ISP) to convert a RAW image into a "conventional" image.



# Noise in images

Can be very pronounced in low-light images.



# Three types of sensor noise

## 1) (Photon) shot noise:

- Photon arrival rates are a random process (Poisson distribution).
- The brighter the scene, the smaller the variance of the distribution.

## 2) Dark-shot noise:

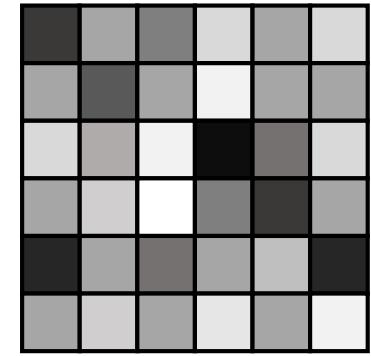
- Emitted electrons due to thermal activity (becomes worse as sensor gets hotter.)

## 3) Read noise:

- Caused by read-out and AFE electronics (e.g., gain, A/D converter).

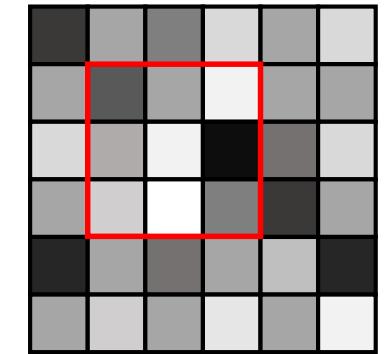
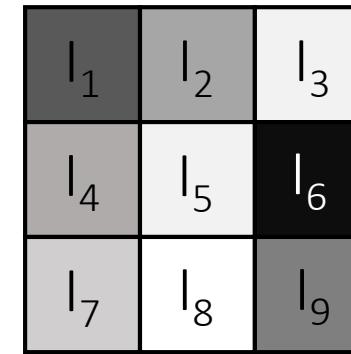
Bright scene and large pixels: photon shot noise is the main noise source.

# How to denoise?



# How to denoise?

Simple denoising: look at the neighborhood around you.



- Mean filtering (take average):

$$I'_5 = \frac{I_1 + I_2 + I_3 + I_4 + I_5 + I_6 + I_7 + I_8 + I_9}{9}$$

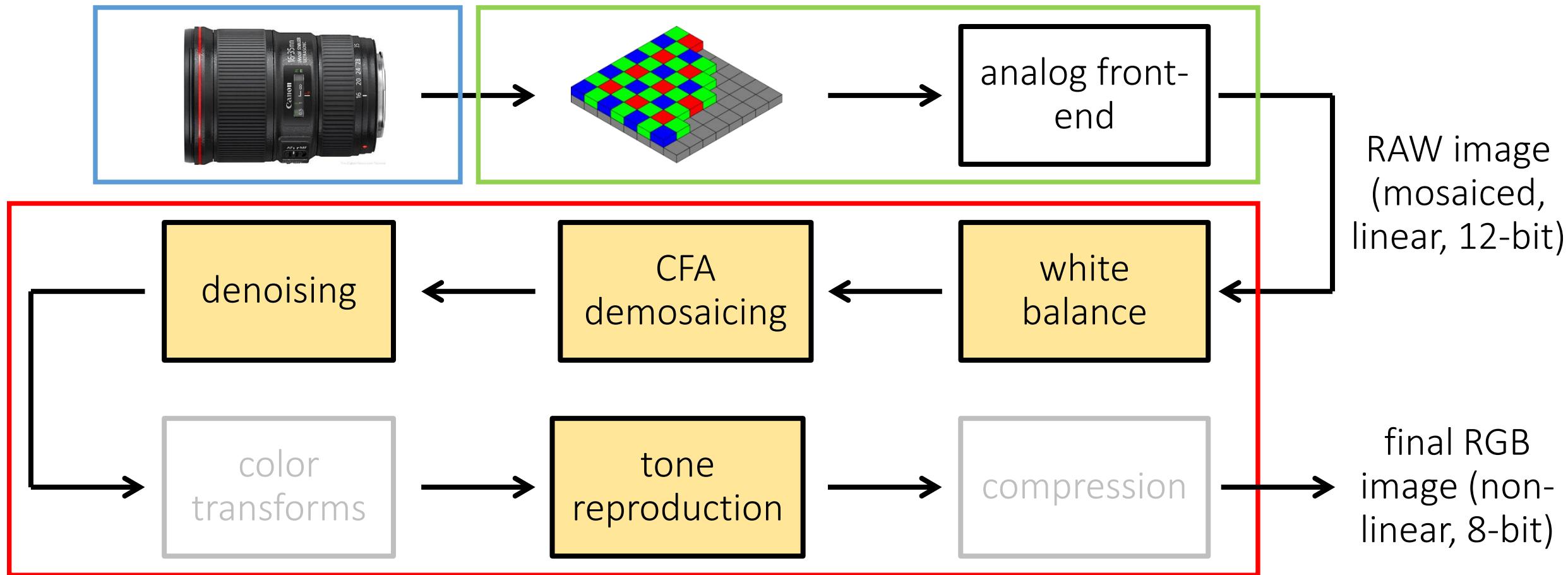
- Median filtering (take median):

$$I'_5 = \text{median}(I_1, I_2, I_3, I_4, I_5, I_6, I_7, I_8, I_9)$$

Large area of research.

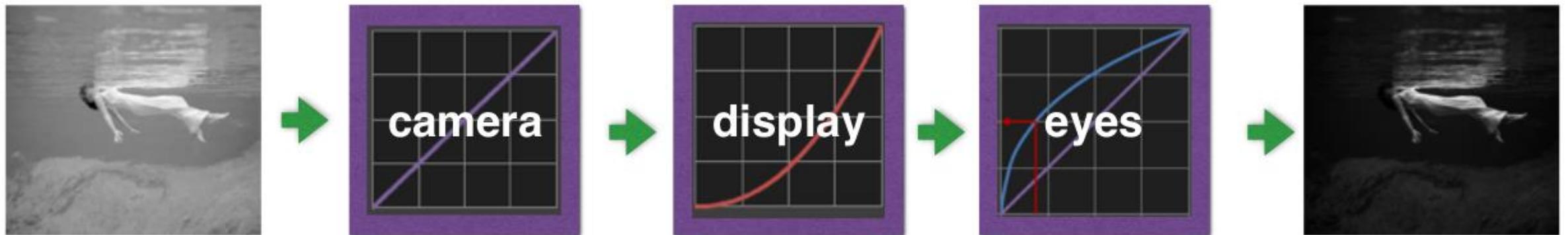
# The (in-camera) image processing pipeline

The sequence of image processing operations applied by the camera's image signal processor (ISP) to convert a RAW image into a "conventional" image.



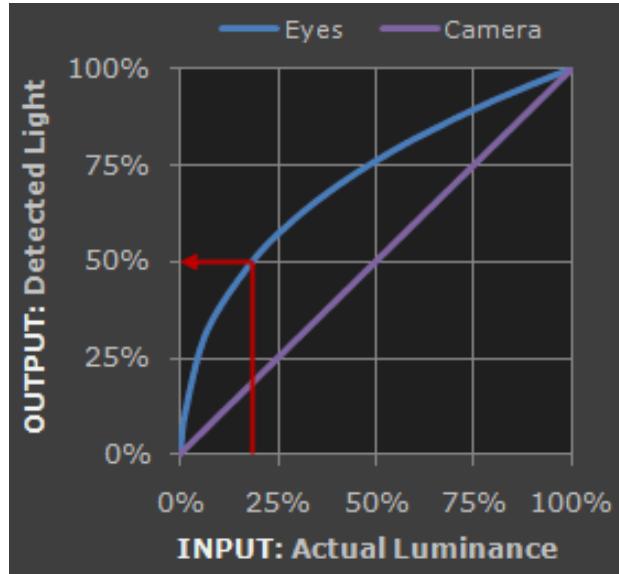
# Tone reproduction

- Also known as gamma correction.
- Without tone reproduction, images look very dark.



Why does this happen?

# Perceived vs measured brightness by human eye



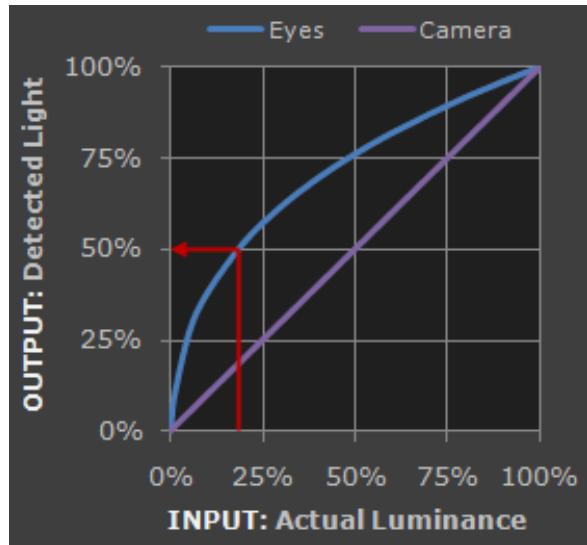
We have already seen that sensor response is linear.

Human-eye *response* (measured brightness) is also linear.

However, human-eye *perception* (perceived brightness) is *non-linear*:

- More sensitive to dark tones.
- Approximately a Gamma function.

# What about displays?

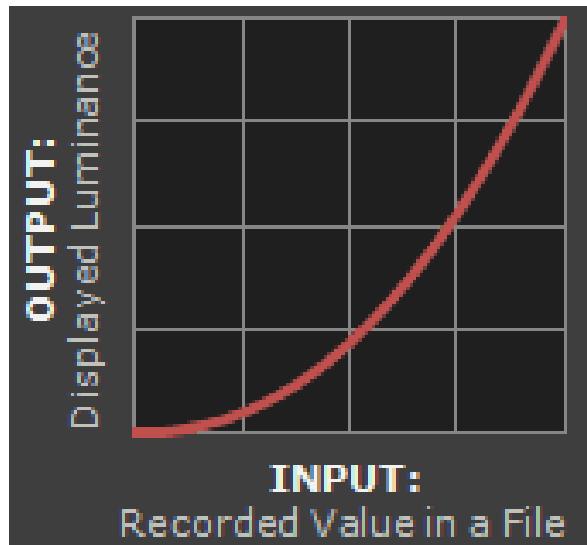


We have already seen that sensor response is linear.

Human-eye *response* (measured brightness) is also linear.

However, human-eye *perception* (perceived brightness) is *non-linear*:

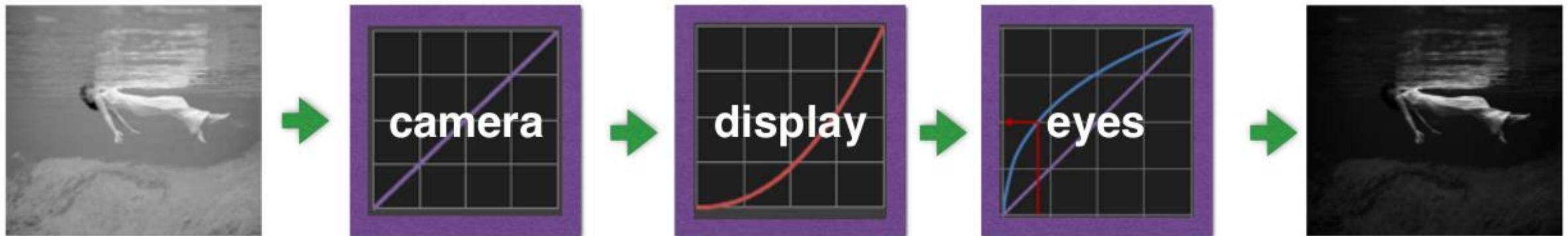
- More sensitive to dark tones.
- Approximately a Gamma function.



Displays have a response opposite to that of human perception.

# Tone reproduction

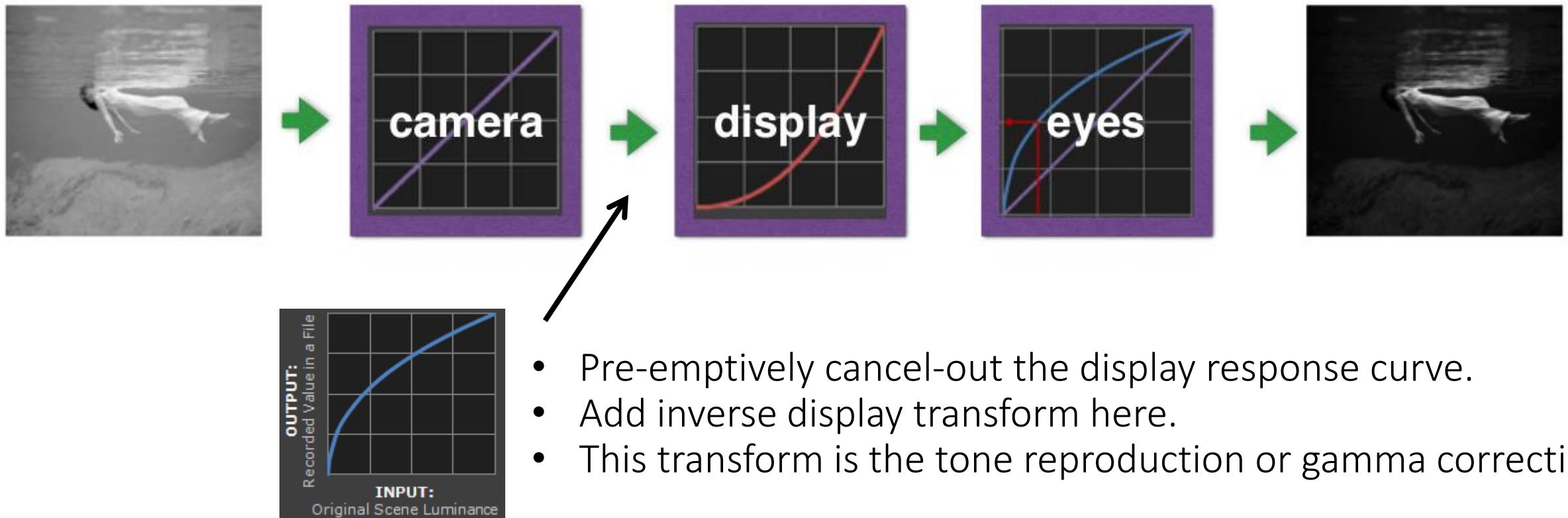
- Because of mismatch in displays and human eye perception, images look very dark.



How do we fix this?

# Tone reproduction

- Because of mismatch in displays and human eye perception, images look very dark.



# Tone reproduction curves

The exact tone reproduction curve depends on the camera.

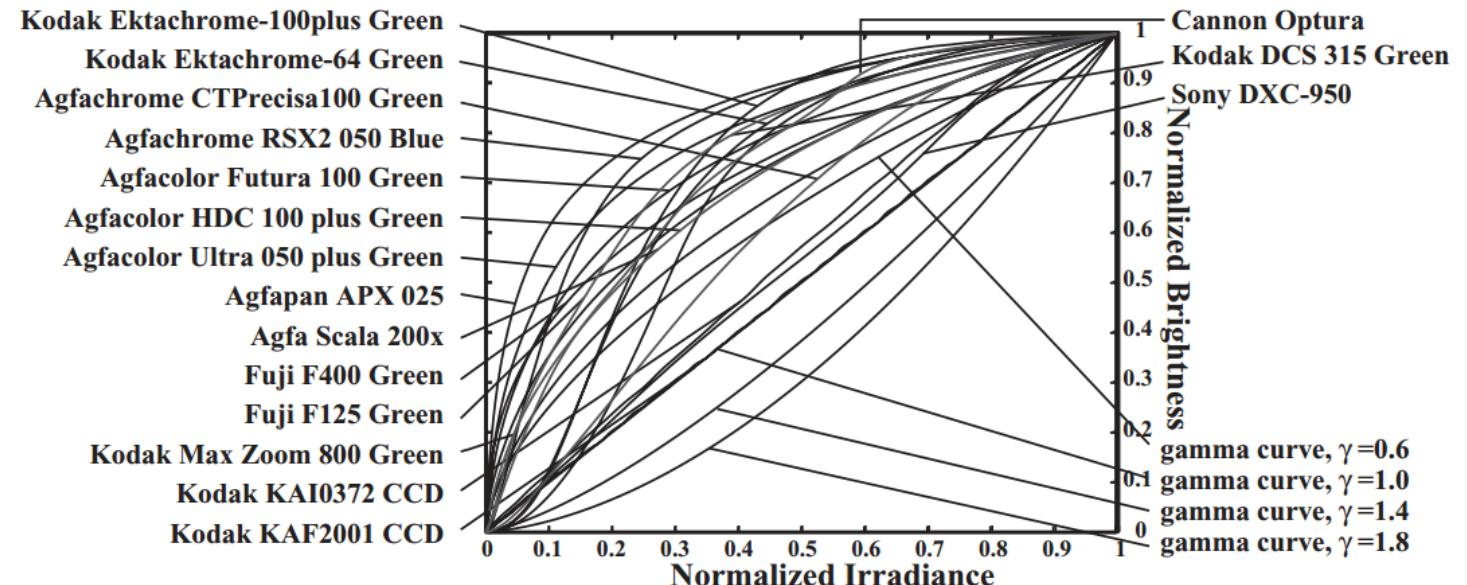
- Often well approximated as  $L^\gamma$ , for different values of the power  $\gamma$  (“gamma”).
- A good default is  $\gamma = 1 / 2.2$ .



before gamma



after gamma



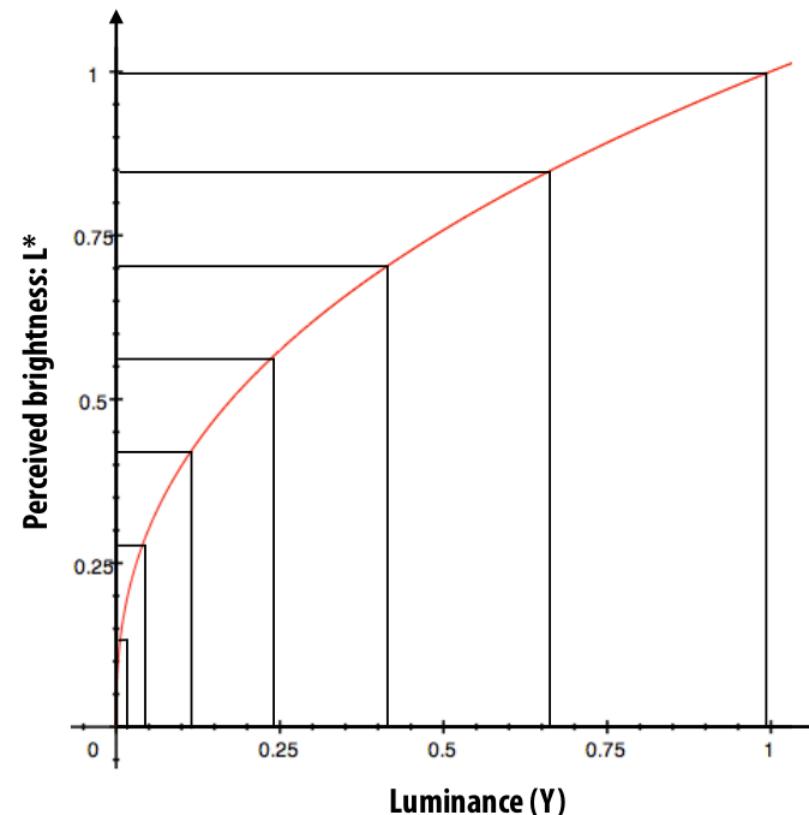
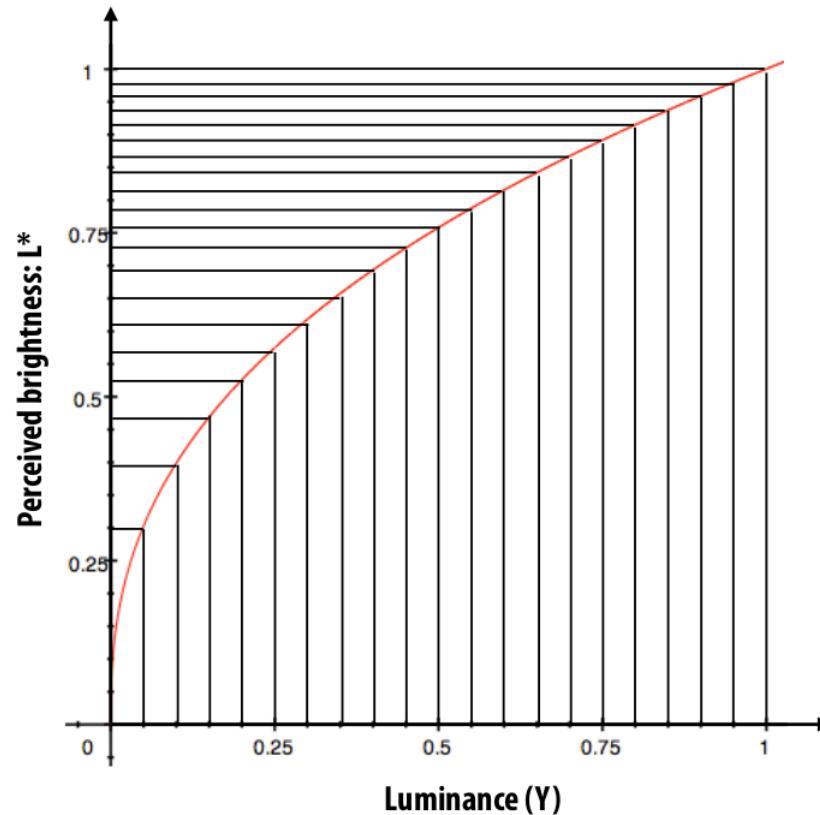
Warning: Our values are no longer linear relative to scene radiance!

# Tone reproduction

Question: Why not just keep measurements linear and do gamma correction right before we display the image?

# Tone reproduction

Question: Why not just keep measurements linear and do gamma correction right before we display the image?

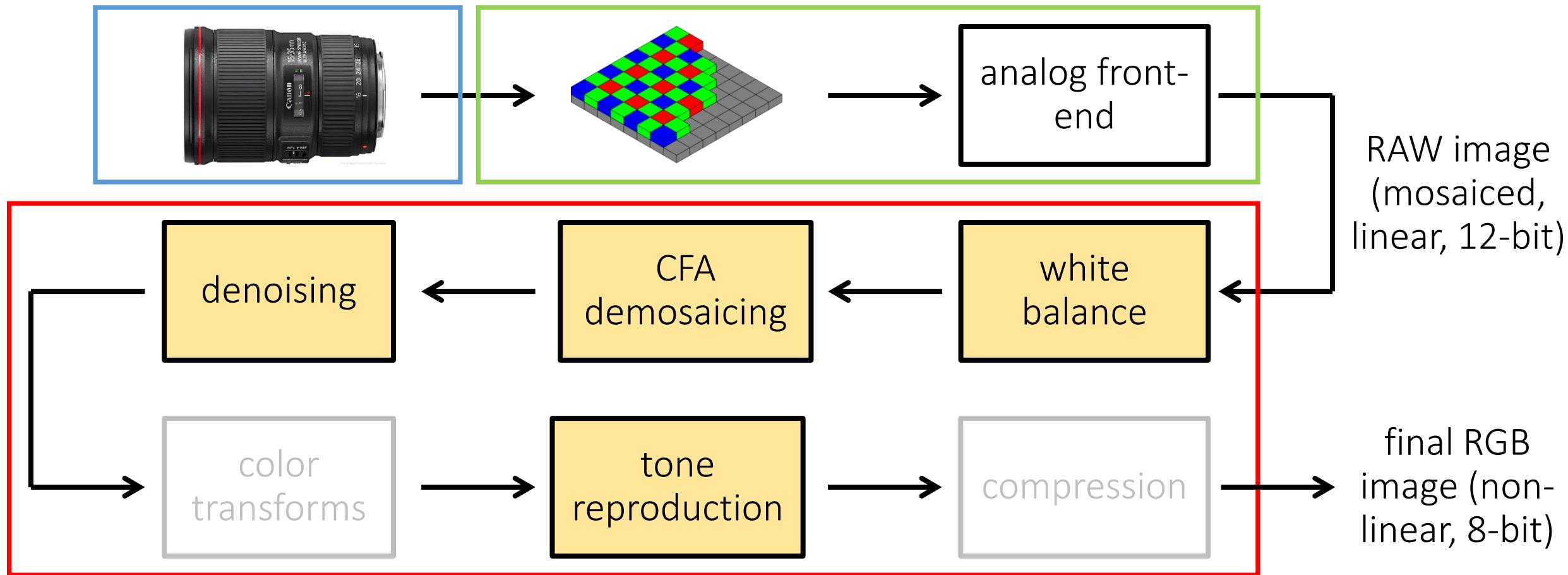


Answer: After this stage, we perform compression, which includes change from 12 to 8 bits.

- Better to use our available bits to encode the information we are going to need.

# The (in-camera) image processing pipeline

The sequence of image processing operations applied by the camera's image signal processor (ISP) to convert a RAW image into a "conventional" image.



Some general thoughts on the image processing  
pipeline

# Do I ever need to use RAW?

# Do I ever need to use RAW?

Emphatic yes!

- Every time you use a physics-based computer vision algorithm, you need *linear measurements of radiance*.
- Examples: photometric stereo, shape from shading, image-based relighting, illumination estimation, anything to do with light transport and inverse rendering, etc.
- Applying the algorithms on non-linear (i.e., not RAW) images will produce completely invalid results.

What if I don't care about physics-based vision?

# What if I don't care about physics-based vision?

You often still *want* (rather than need) to use RAW!

- If you like re-finishing your photos (e.g., on Photoshop), RAW makes your life much easier and your edits much more flexible.

Is it even possible to get access to RAW images?

# Is it even possible to get access to RAW images?

Quite often yes!

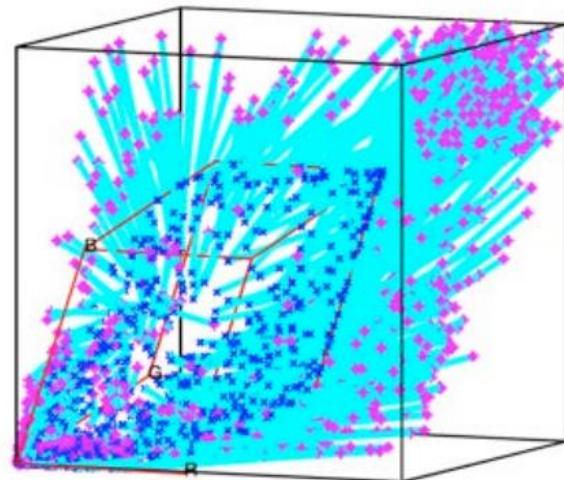
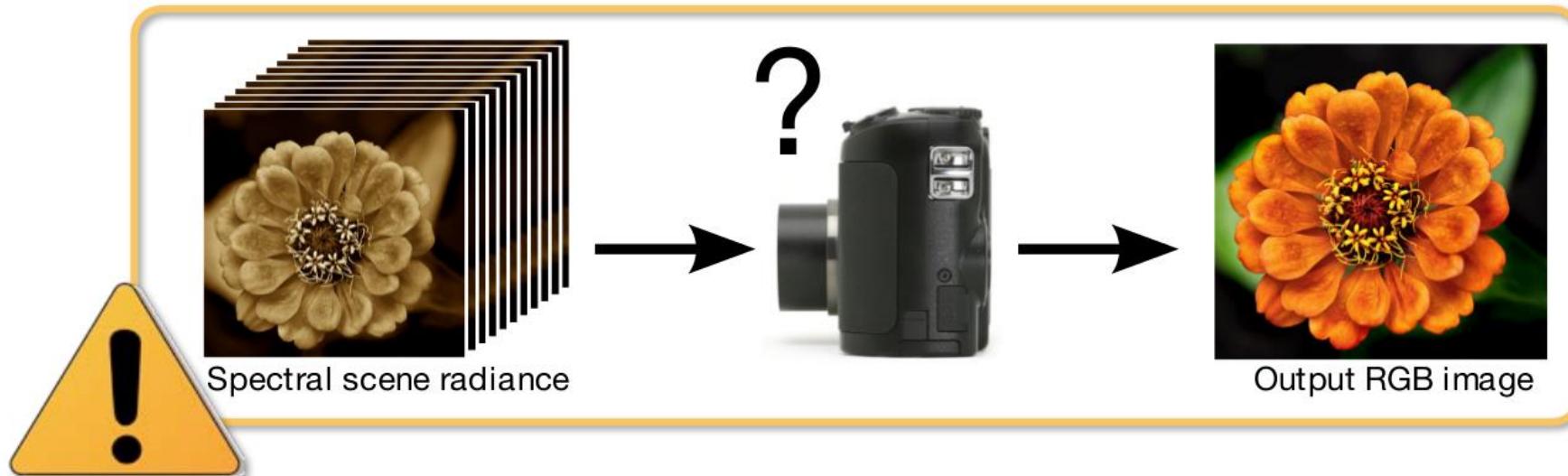
- Most DSLR cameras provide an option to store RAW image files.
- Certain phone cameras allow, directly or indirectly, access to RAW.
- Sometimes, it may not be “fully” RAW. The Lightroom app provides images after demosaicking but before tone reproduction.

# I forgot to set my camera to RAW, can I still get the RAW file?

Nope, tough luck.

- The image processing pipeline is lossy: After all the steps, information about the original image is lost.
- Sometimes we may be able to reverse a camera's image processing pipeline *if we know exactly what it does* (e.g., by using information from other similar RAW images).
- The conversion of PNG/JPG back to RAW is known as “de-rendering” and is an active research area.

# Derendering



RAW ←  
JPEG/sRGB ←

Panasonic DMC-LX3

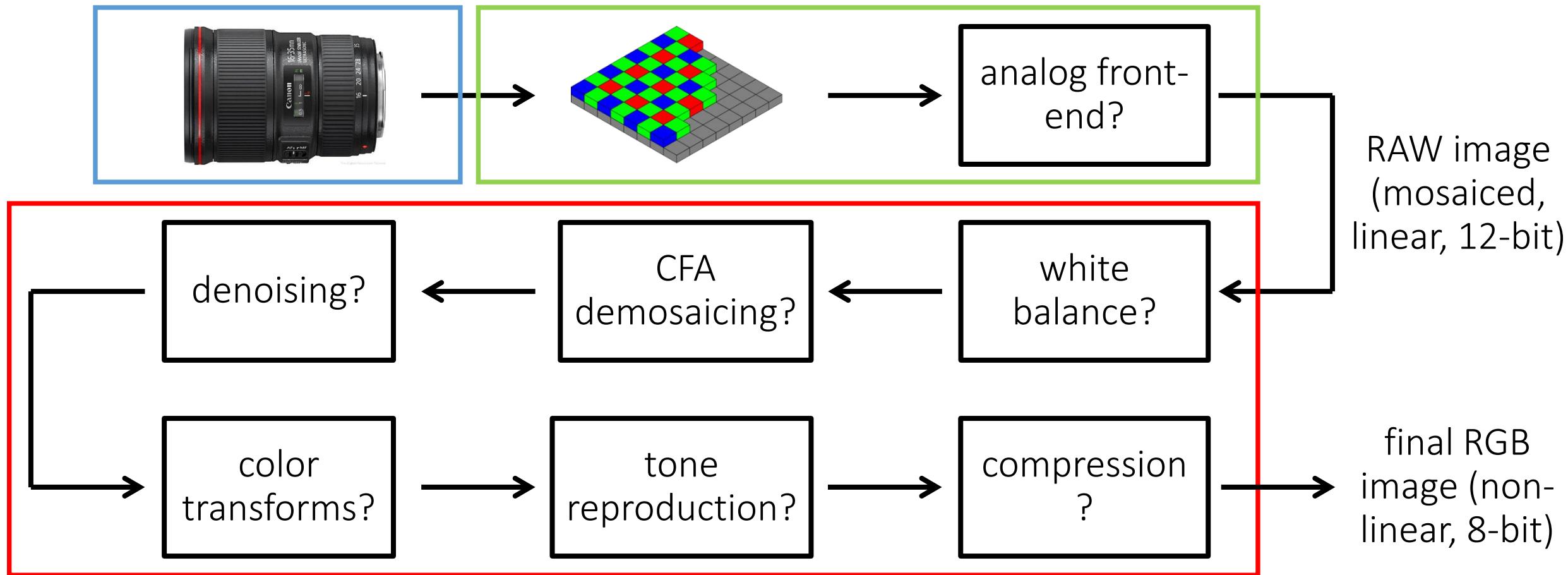
# Why did you use italics in the previous slide?

What I described today is an “idealized” version of what we *think* commercial cameras do.

- Almost all of the steps in both the sensor and image processing pipeline I described earlier are camera-dependent.
- Even if we know the basic steps, the implementation details are proprietary information that companies actively try to keep secret.
- I will go back to a few of my slides to show you examples of the above.

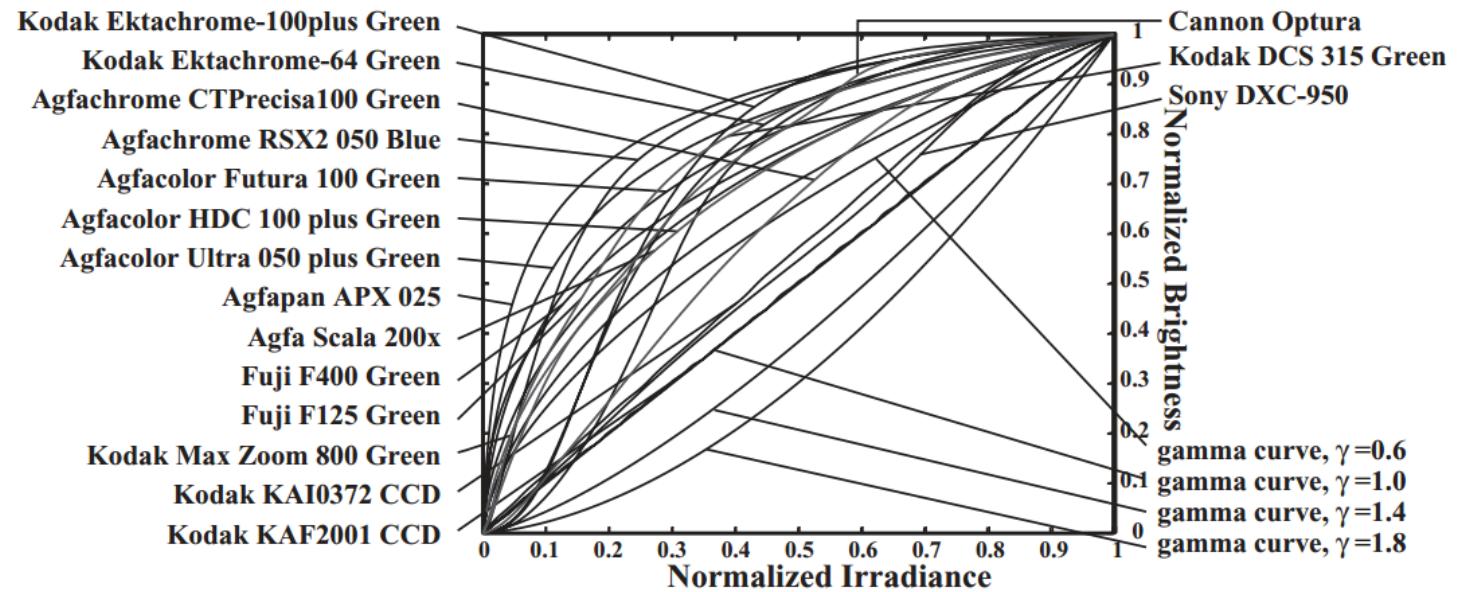
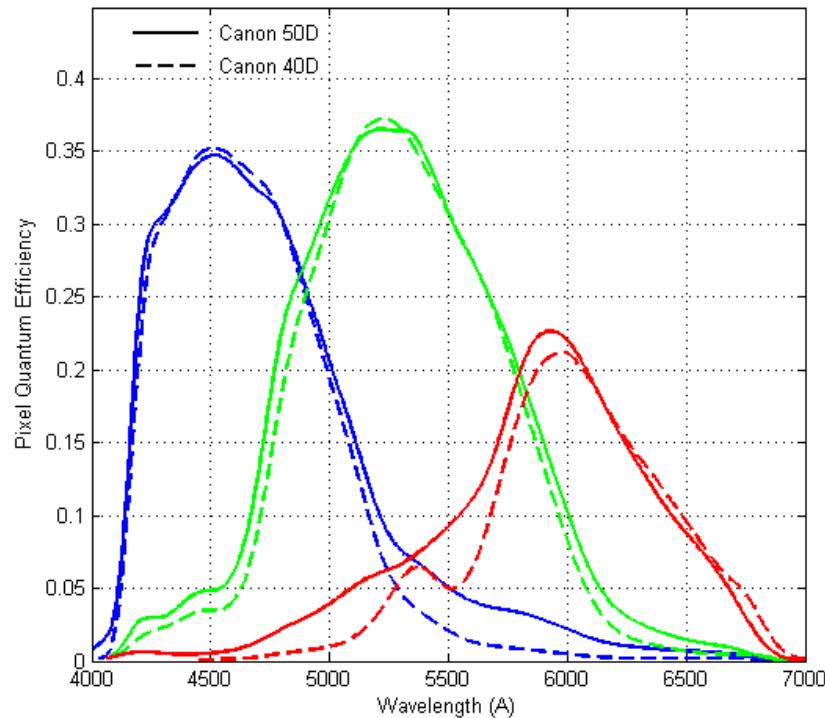
# The hypothetical image processing pipeline

The sequence of image processing operations applied by the camera's image signal processor (ISP) to convert a RAW image into a "conventional" image.



# Various curves

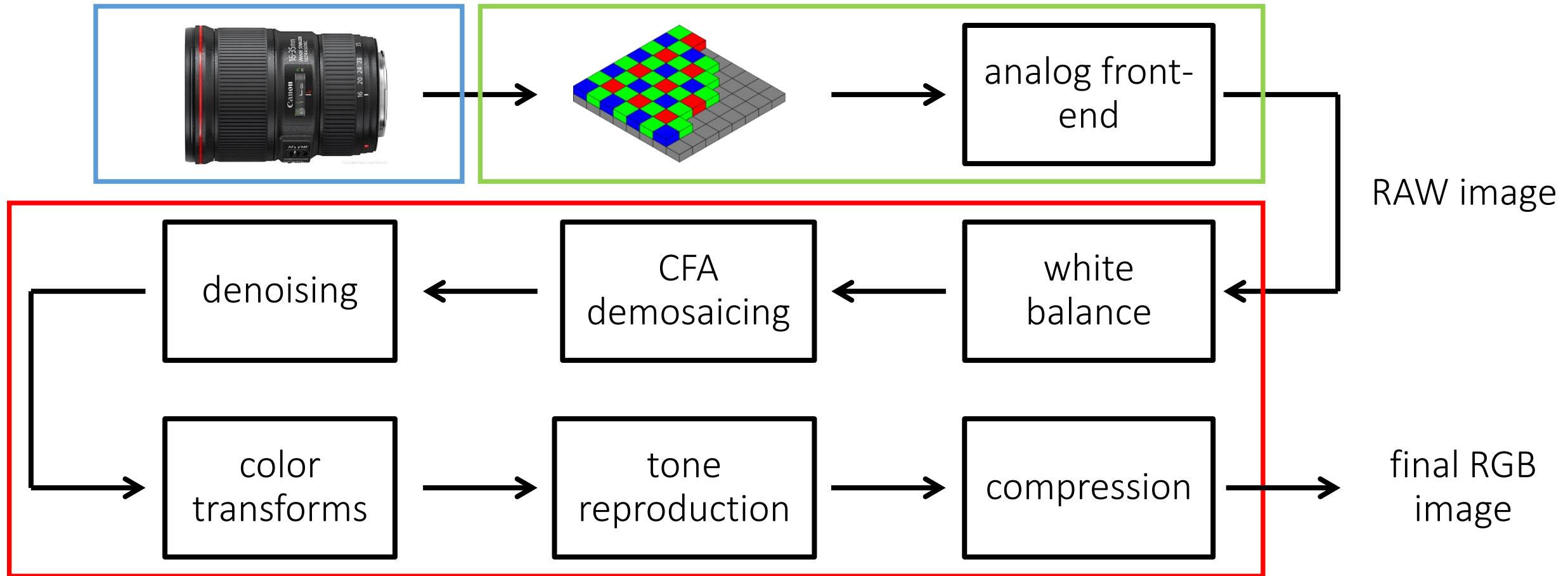
All of these sensitivity curves are different from camera to camera and kept secret.



Radiometric calibration  
(a.k.a. high dynamic range imaging)  
(a.k.a. capturing linear images)

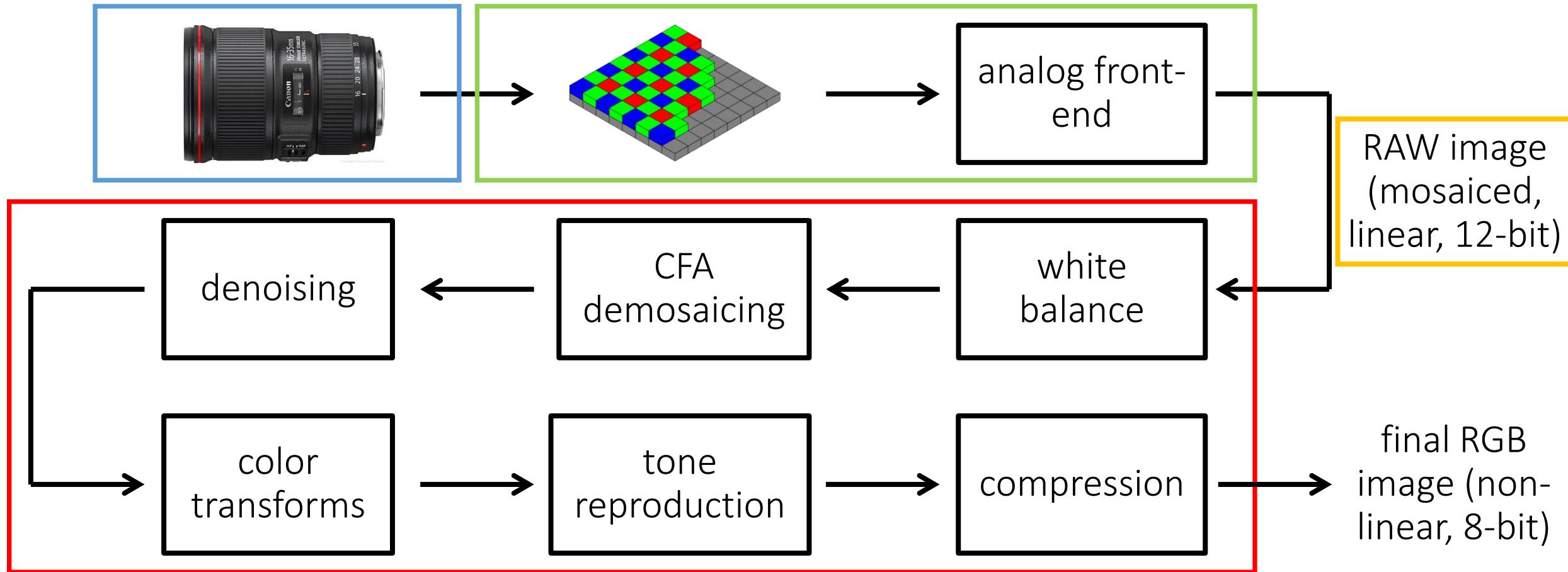
# The image processing pipeline

Which parts of the image processing pipeline introduce non-linearities?



# The image processing pipeline

Is using RAW images sufficient to get linear images?

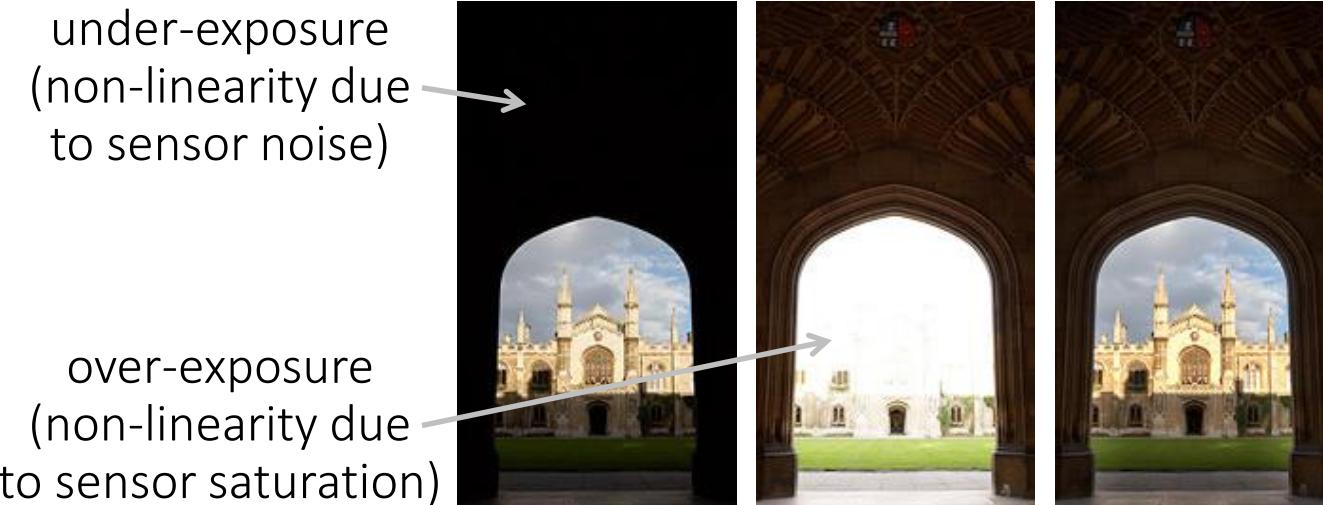


# Photodiode response function

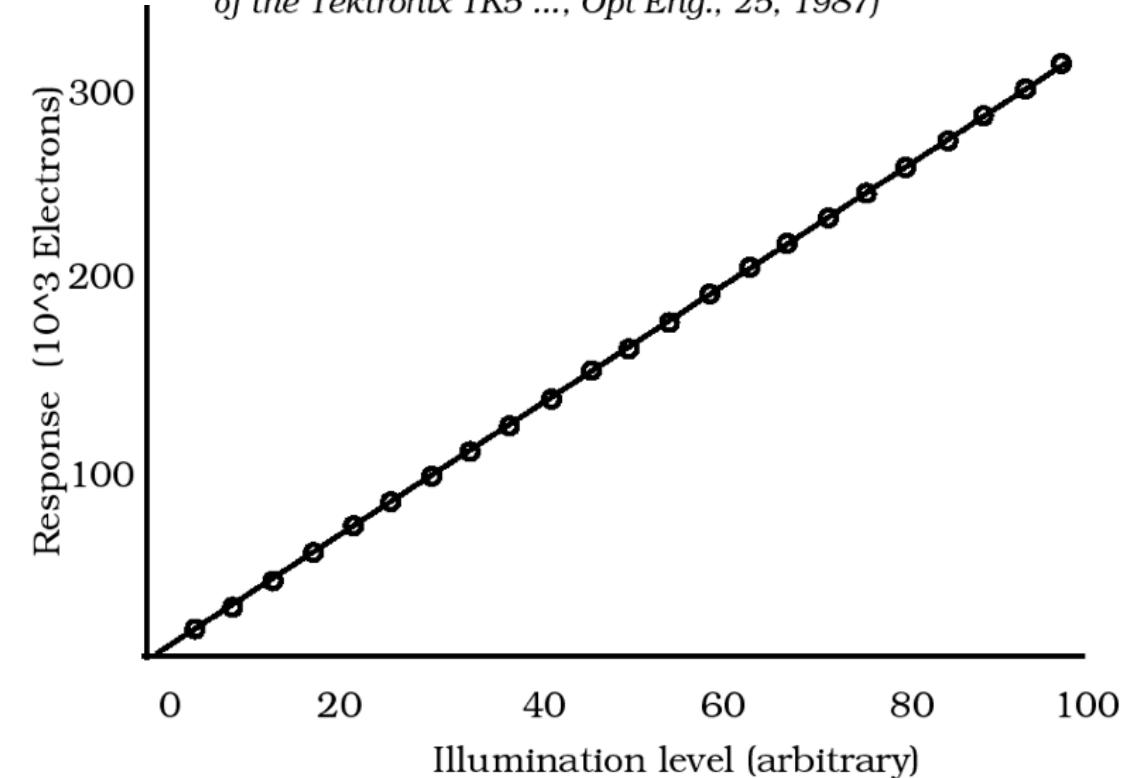
For silicon photodiodes, usually linear, but:

- non-linear when potential well is saturated (over-exposure)
- non-linear near zero (due to noise)

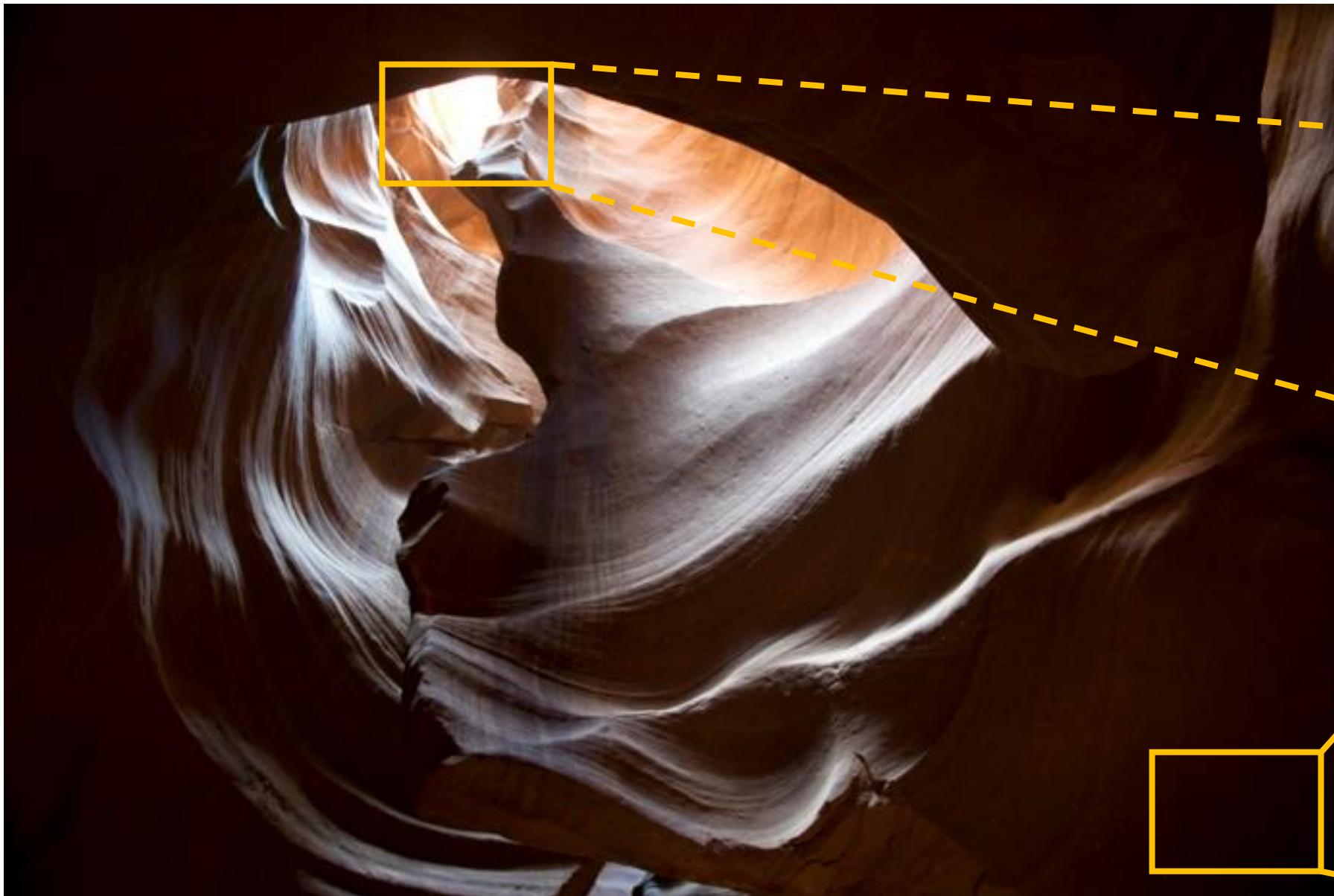
We will see how to deal with these issues in a later lecture (high-dynamic-range imaging).



(Epperson, P.M. et al. *Electro-optical characterization of the Tektronix TK5 ...*, *Opt Eng.*, 25, 1987)



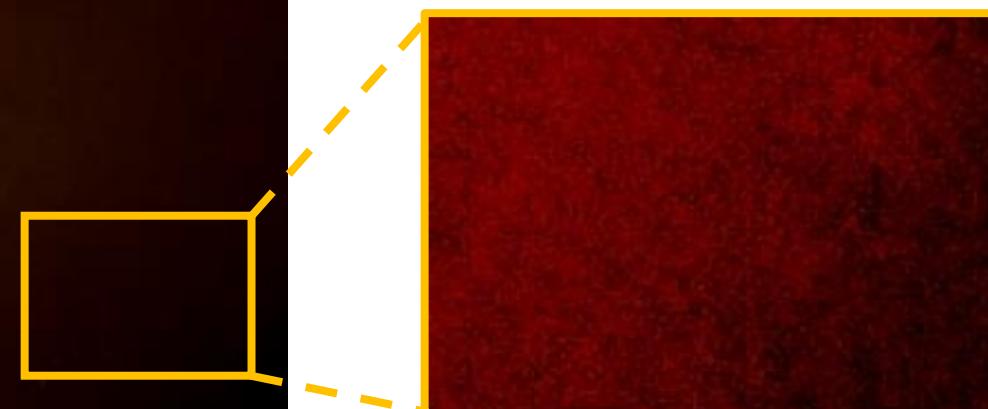
# Over/under exposure



in highlights we are limited by clipping



in shadows we are limited by noise













Our devices do not match the world

# The world has a high dynamic range



1



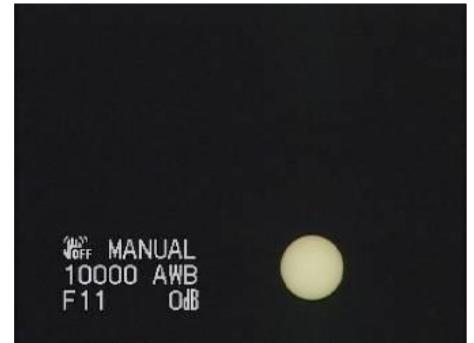
1500



25,000



400,000

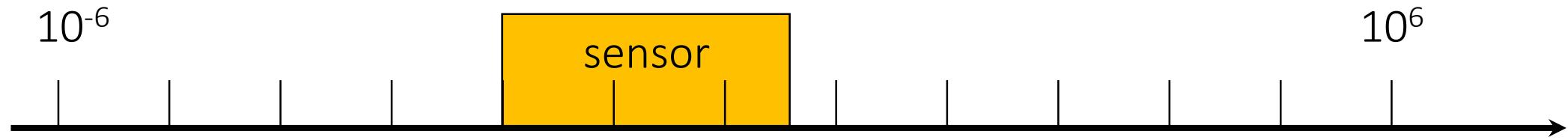


100,000,000

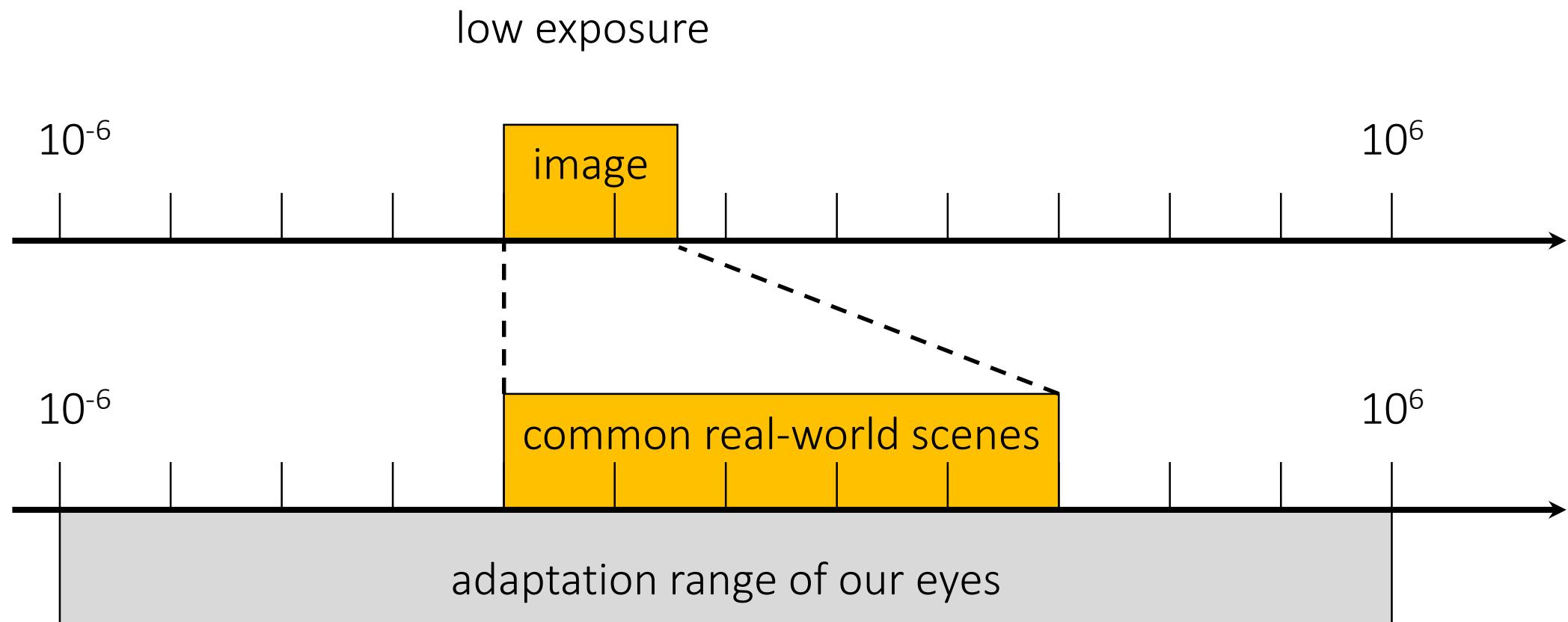
# The world has a high dynamic range



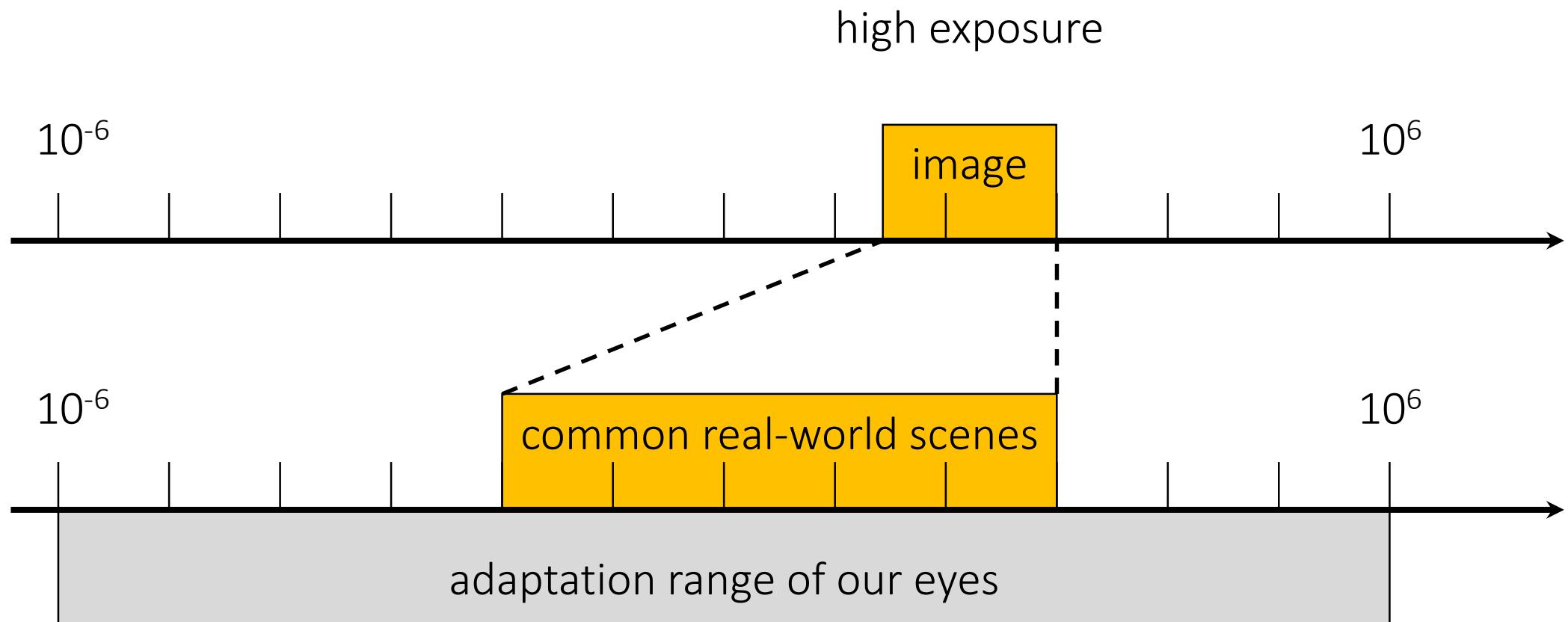
(Digital) sensors also have a low dynamic range



# (Digital) images have an even lower dynamic range



# (Digital) images have an even lower dynamic range



# Our devices do not match the real world

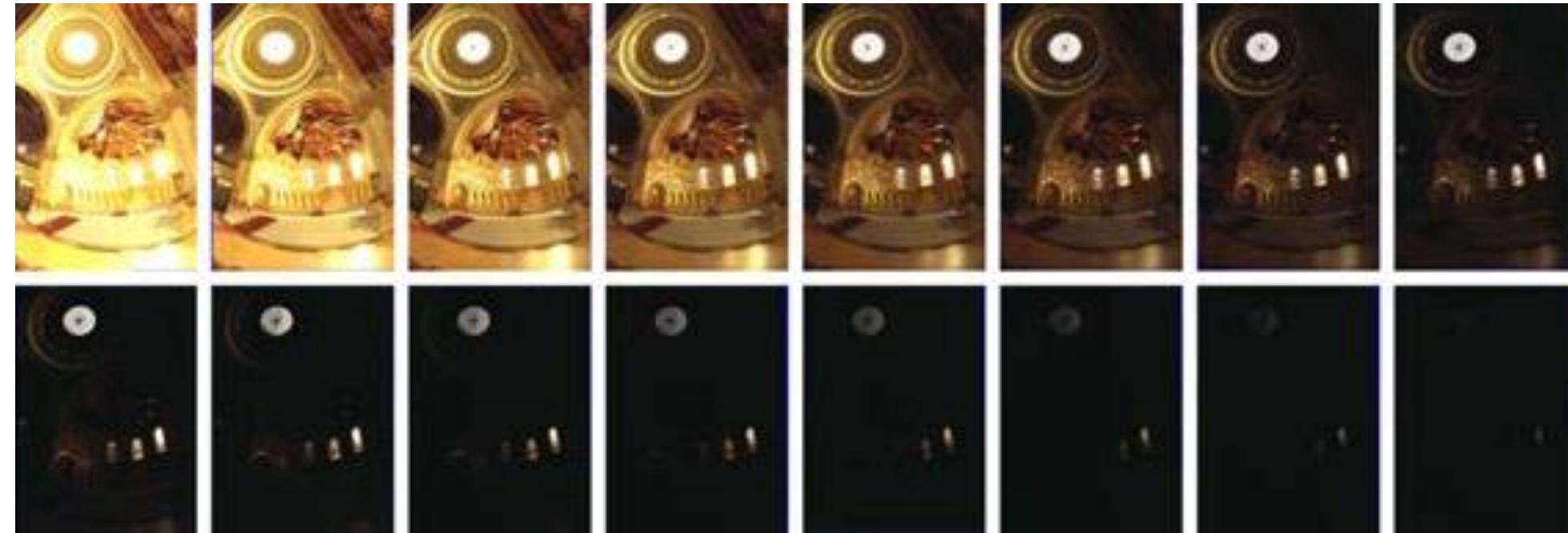
- 10:1 photographic print (higher for glossy paper)
- 20:1 artist's paints
- 200:1 slide film
- 500:1 negative film
- 1000:1 LCD display
- 2000:1 digital SLR (at 12 bits)
- 100000:1 real world

Two challenges:

1. HDR imaging – which parts of the world to include to the 8-12 bits available to our device?
2. Tonemapping – which parts of the world to display in the 4-10 bits available to our device?

# Key idea

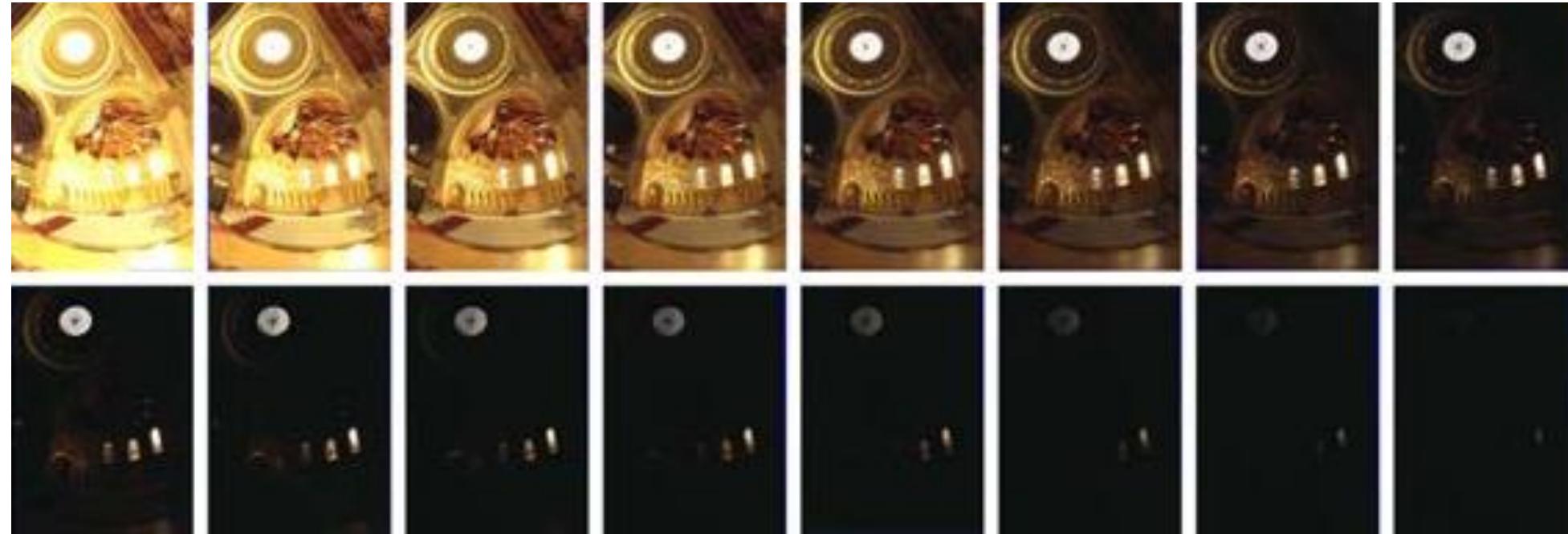
1. Capture multiple LDR images at different exposures



2. Merge them into a single HDR image

# Key idea

1. Capture multiple LDR images at different exposures

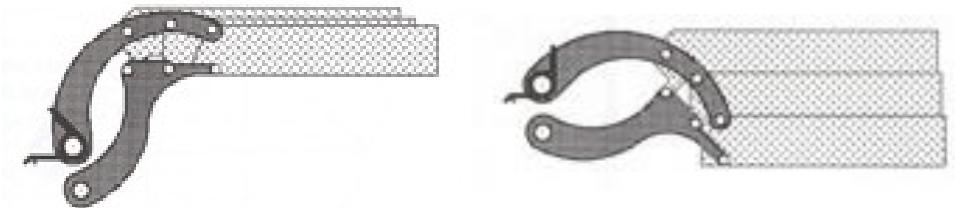


2. Merge them into a single HDR image

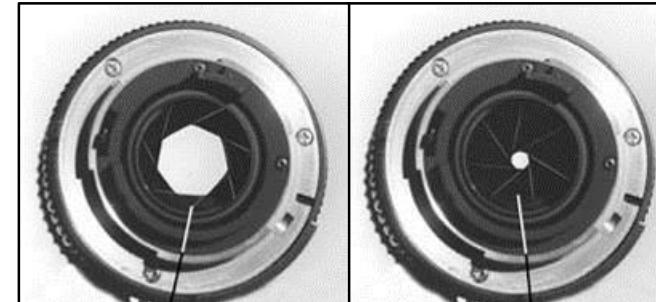


# Ways to vary exposure

1. Shutter speed



2. F-stop (aperture, iris)



3. ISO



4. Neutral density (ND) filters

Pros and cons of each?

# Ways to vary exposure

1. Shutter speed
  - Range: about 30 sec to 1/4000 sec (6 orders of magnitude)
  - Pros: repeatable, linear
  - Cons: noise and motion blur for long exposure
2. F-stop (aperture, iris)
  - Range: about f/0.98 to f/22 (3 orders of magnitude)
  - Pros: fully optical, no noise
  - Cons: changes depth of field
3. ISO
  - Range: about 100 to 1600 (1.5 orders of magnitude)
  - Pros: no movement at all
  - Cons: noise
3. Neutral density (ND) filters
  - Range: up to 6 densities (6 orders of magnitude)
  - Pros: works with strobe/flash
  - Cons: not perfectly neutral (color shift), extra glass (interreflections, aberrations), need to touch camera (shake)

# Shutter speed

Note: shutter times usually obey a power series – each “stop” is a factor of 2

$1/4, 1/8, 1/15, 1/30, 1/60, 1/125, 1/250, 1/500, 1/1000$  sec

usually really is

$1/4, 1/8, 1/16, 1/32, 1/64, 1/128, 1/256, 1/512, 1/1024$  sec

Questions:

1. How many exposures?
2. What exposures?

# Shutter speed

Note: shutter times usually obey a power series – each “stop” is a factor of 2

$1/4, 1/8, 1/15, 1/30, 1/60, 1/125, 1/250, 1/500, 1/1000$  sec

usually really is

$1/4, 1/8, 1/16, 1/32, 1/64, 1/128, 1/256, 1/512, 1/1024$  sec

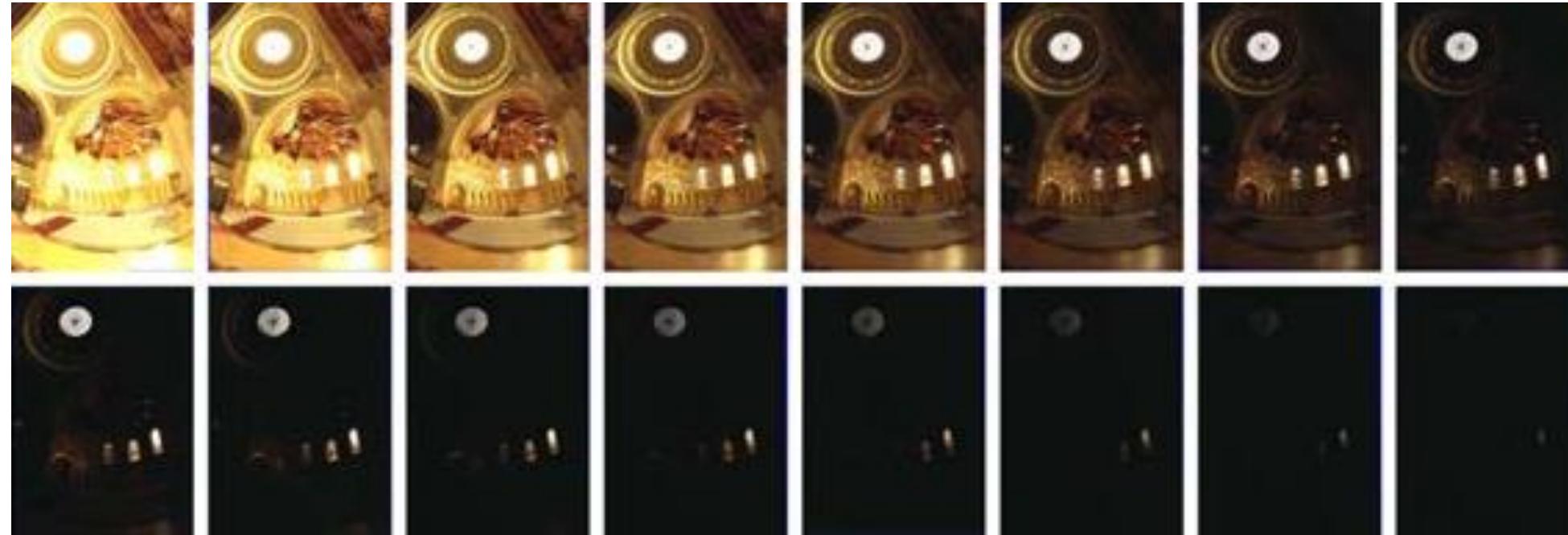
Questions:

1. How many exposures?
2. What exposures?

Answer: Depends on the scene, but a good default is 5 exposures, metered exposure and +- 2 stops around that

# Key idea

1. Capture multiple LDR images at different exposures

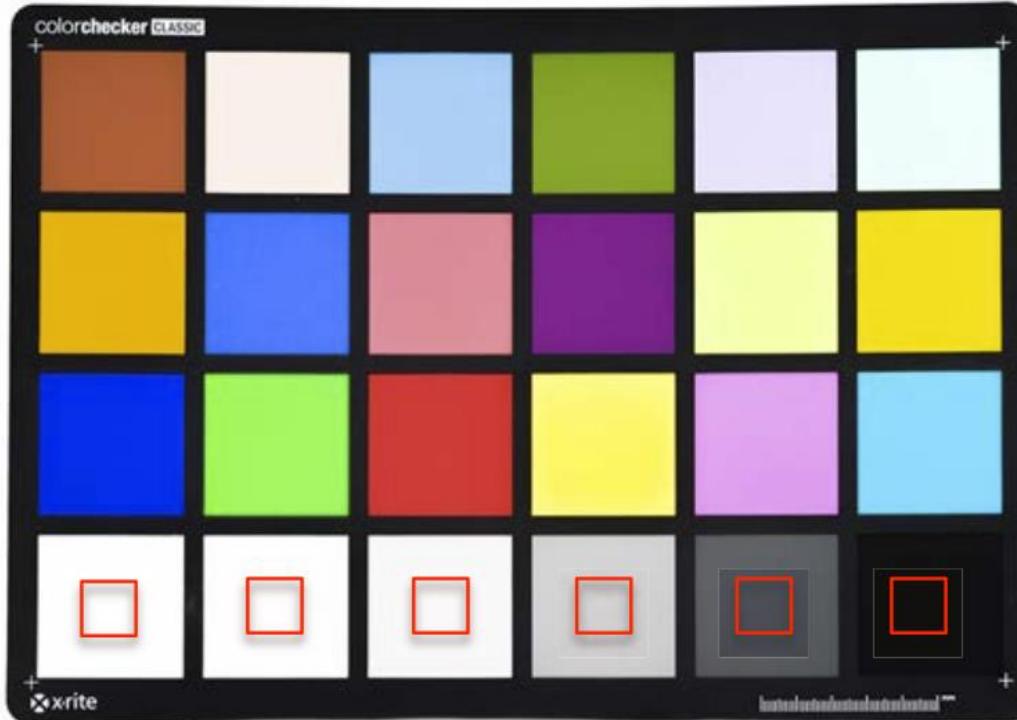


2. Merge them into a single HDR image



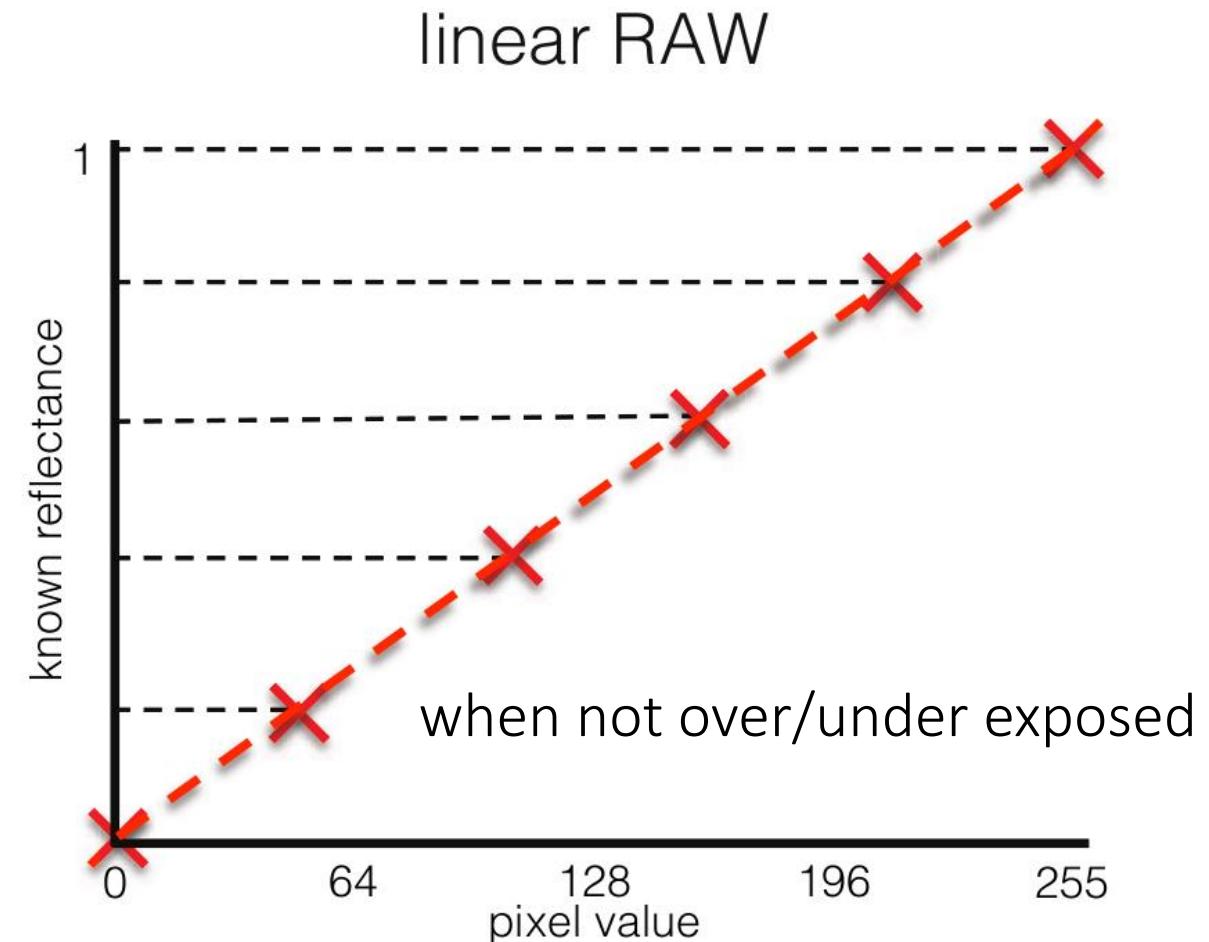
# RAW images have a linear response curve

No need for calibration in this case

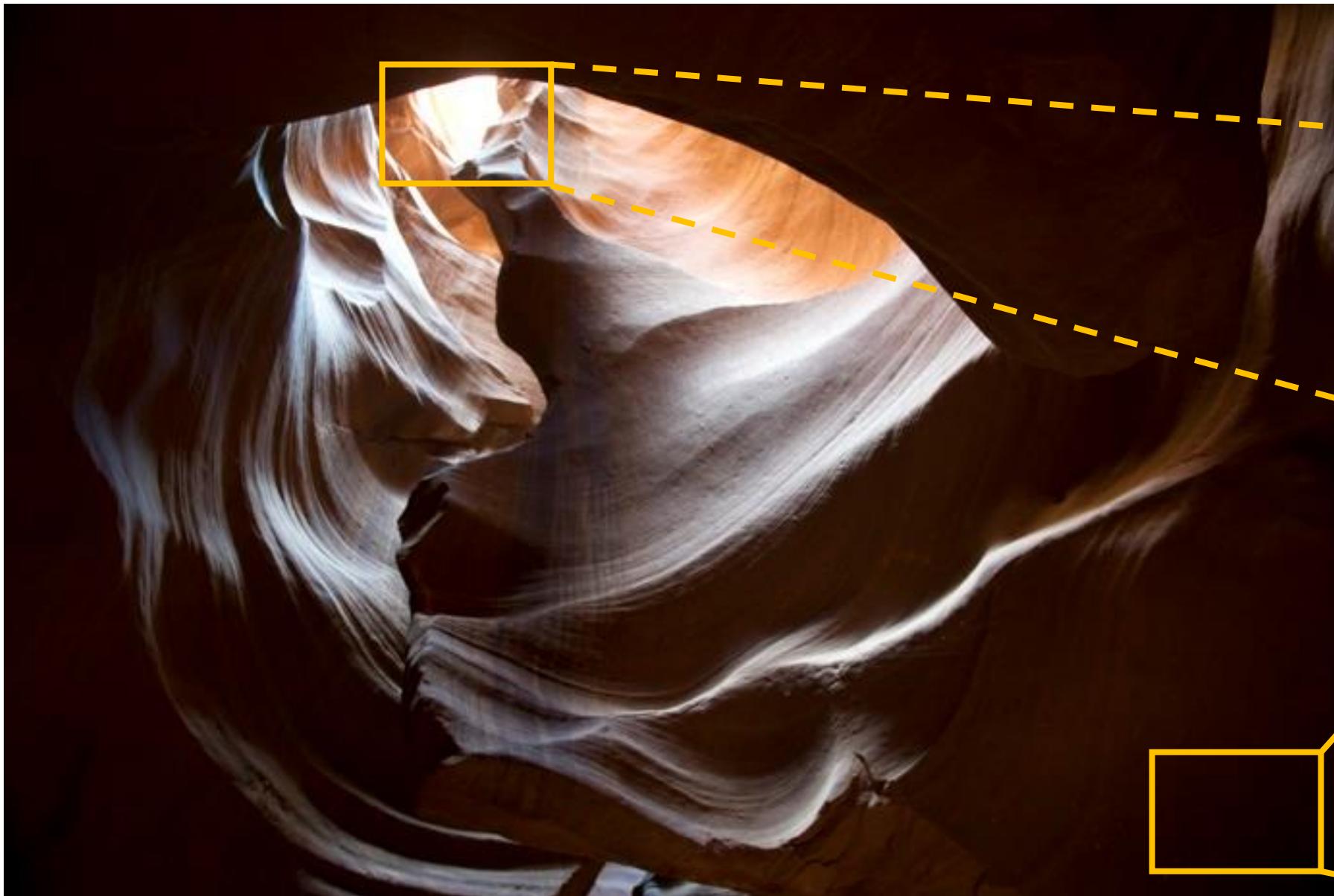


Calibration chart can be used for:

1. color calibration
2. radiometric calibration (i.e., response curve) using the bottom row



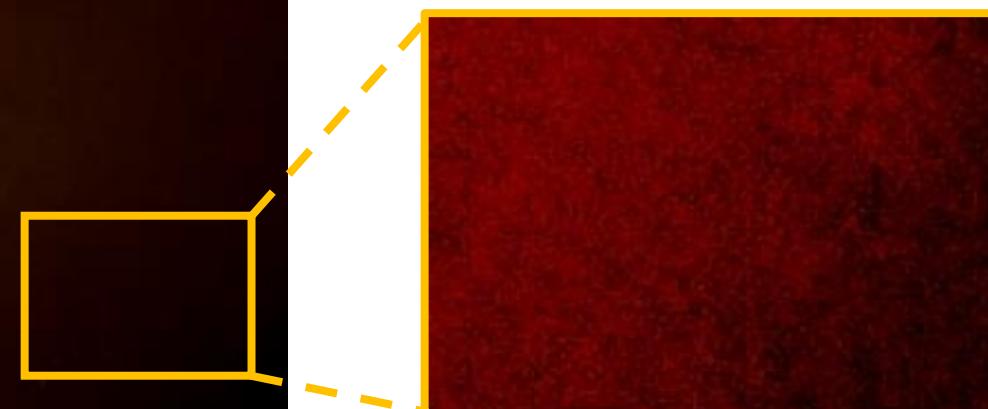
# Over/under exposure



in highlights we are limited by clipping



in shadows we are limited by noise



# RAW (linear) image formation model

Real scene radiance for image pixel  $(x, y)$ :  $L(x, y)$

Exposure time:

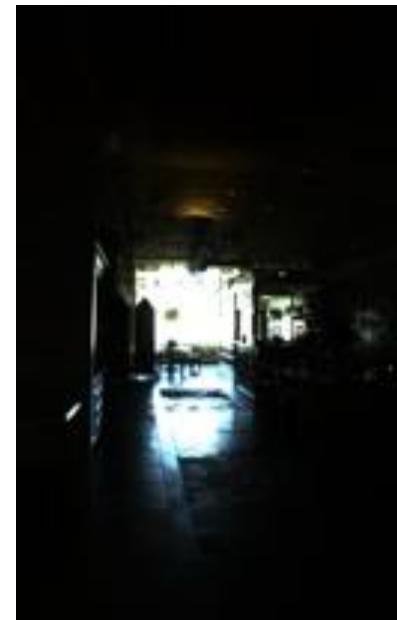
$t_5$



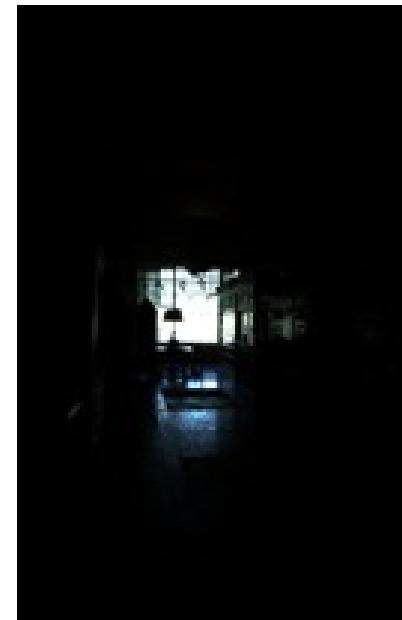
$t_4$



$t_3$



$t_2$



$t_1$



What is an expression for the image  $I(x, y)$  as a function of  $L(x, y)$ ?

# RAW (linear) image formation model

Real scene radiance for image pixel  $(x, y)$ :  $L(x, y)$

Exposure time:



What is an expression for the image  $I_{\text{linear}}(x, y)$  as a function of  $L(x, y)$ ?

$$I_{\text{linear}}(x, y) = \text{clip}[ t_i \cdot L(x, y) + \text{noise} ]$$

How would you merge these images into an HDR one?

# Merging RAW (linear) exposure stacks

For each pixel:

1. Find “valid” images
2. Weight valid pixel values appropriately
3. Form a new pixel value as the weighted average of valid pixel values

How would you  
implement steps 1-2?

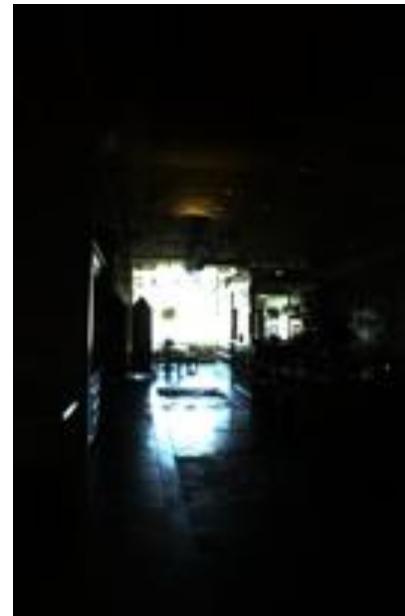
$t_5$



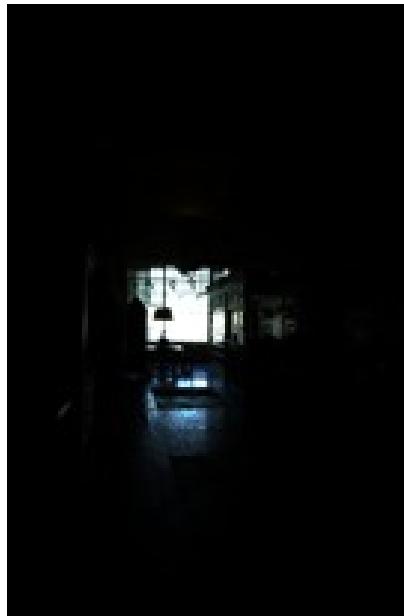
$t_4$



$t_3$



$t_2$



$t_1$



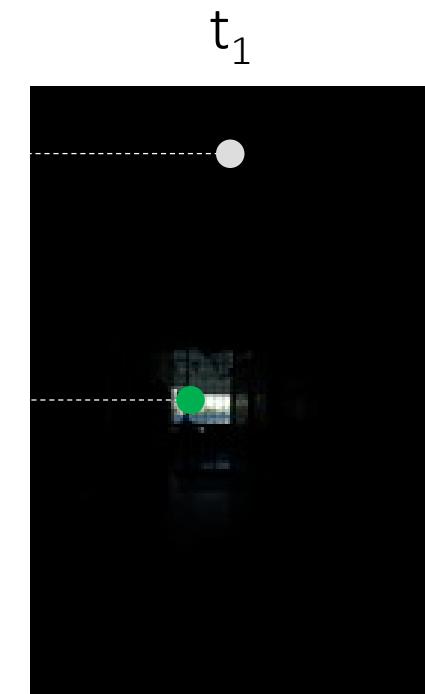
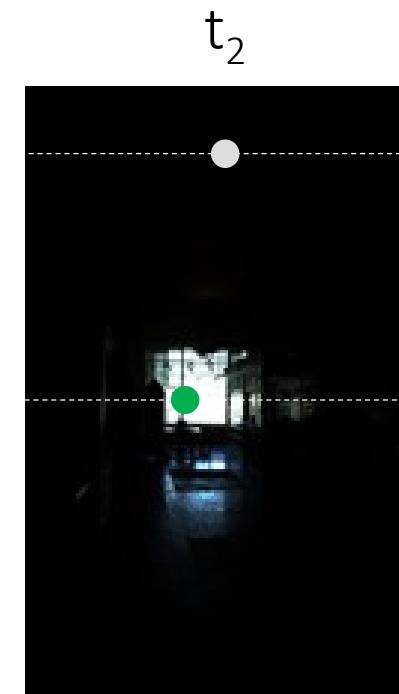
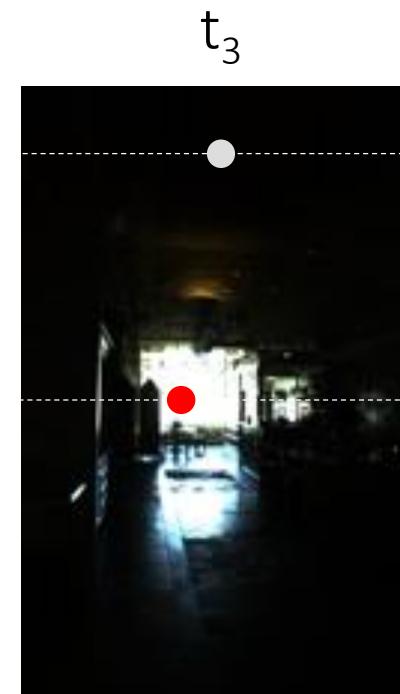
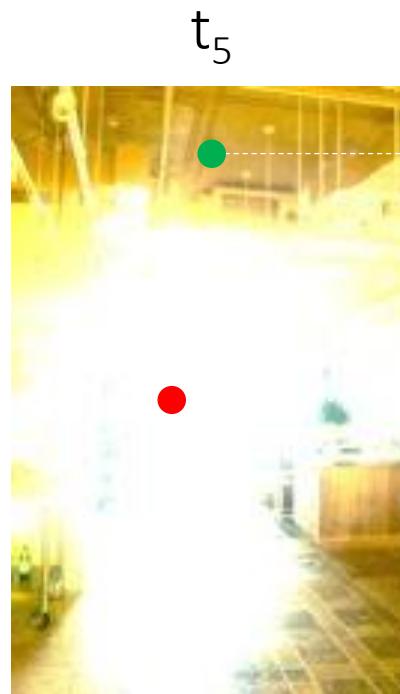
# Merging RAW (linear) exposure stacks

For each pixel:

1. Find “valid” images
2. Weight valid pixel values appropriately
3. Form a new pixel value as the weighted average of valid pixel values

← (noise)  $0.05 < \text{pixel} < 0.95$  (clipping)

- noise
- valid
- clipped



# Merging RAW (linear) exposure stacks

For each pixel:

1. Find “valid” images  $\leftarrow$  (noise)  $0.05 < \text{pixel} < 0.95$  (clipping)
2. Weight valid pixel values appropriately  $\leftarrow$  (pixel value) /  $t_i$
3. Form a new pixel value as the weighted average of valid pixel values

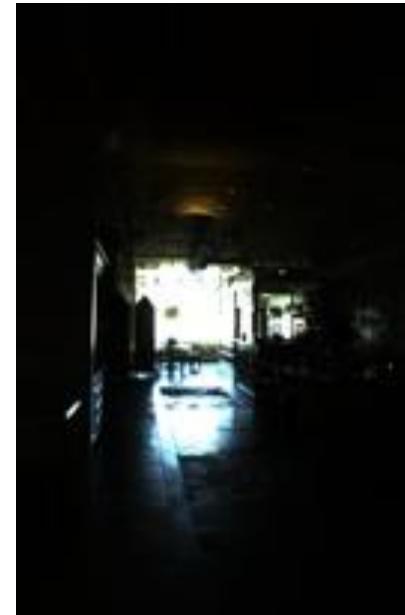
$t_5$



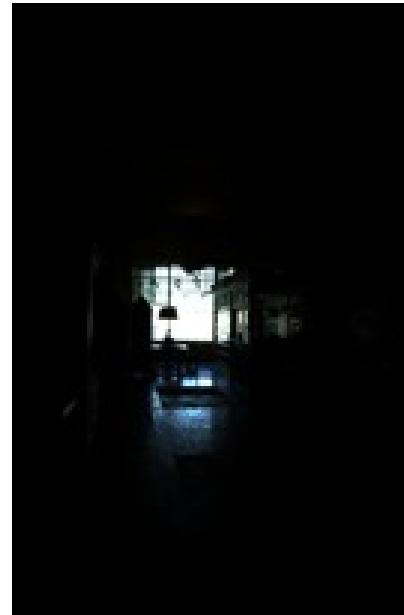
$t_4$



$t_3$



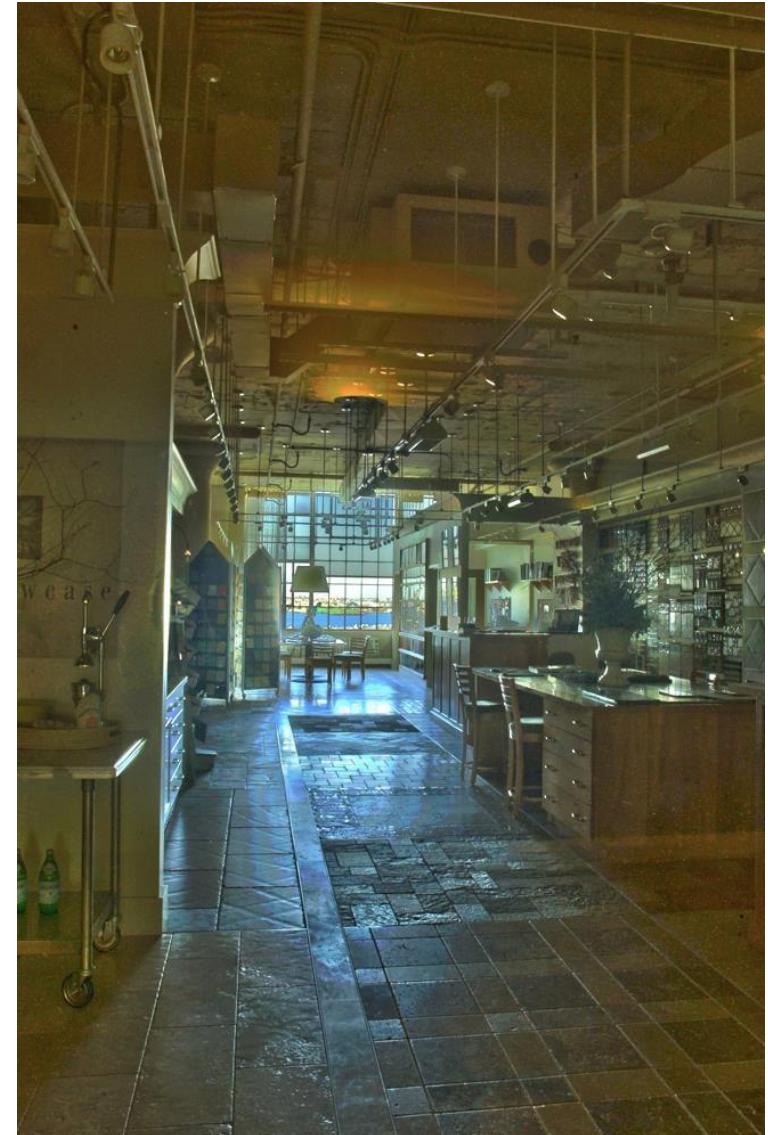
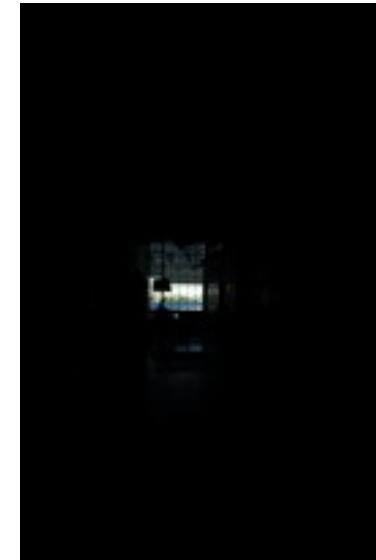
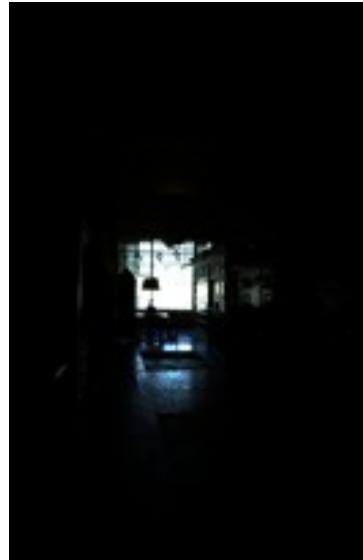
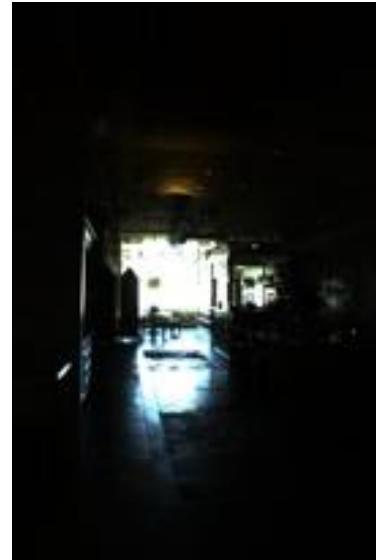
$t_2$



$t_1$



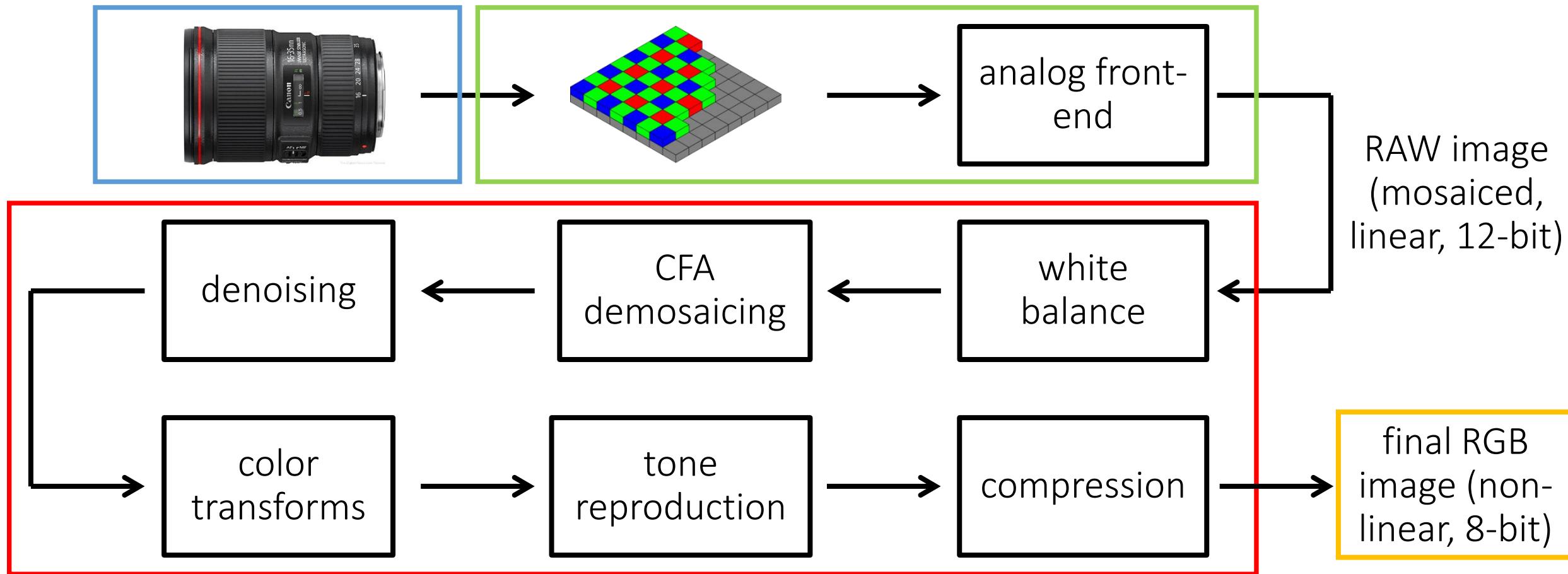
# Merging result (after tonemapping)



What if I cannot use raw?

# The image processing pipeline

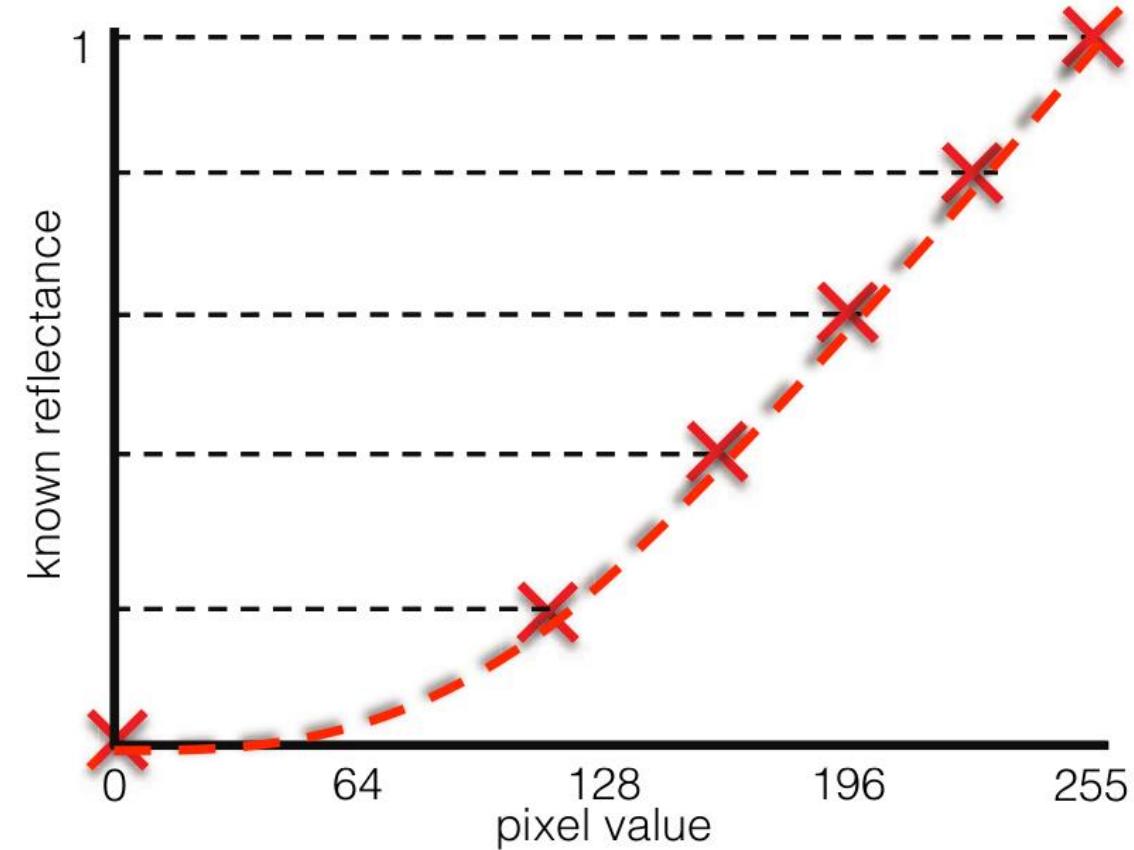
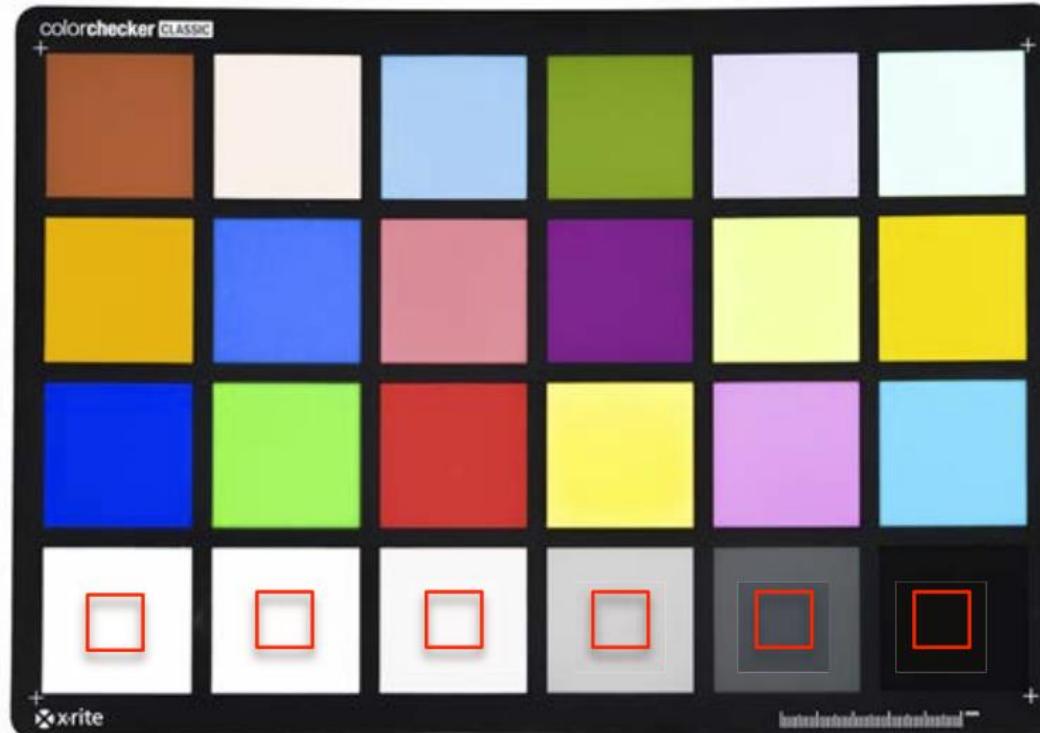
The sequence of image processing operations applied by the camera's image signal processor (ISP) to convert a RAW image into a "conventional" image.



# Processed images have a non-linear response curve

We must calibrate the response curve

e.g. JPEG

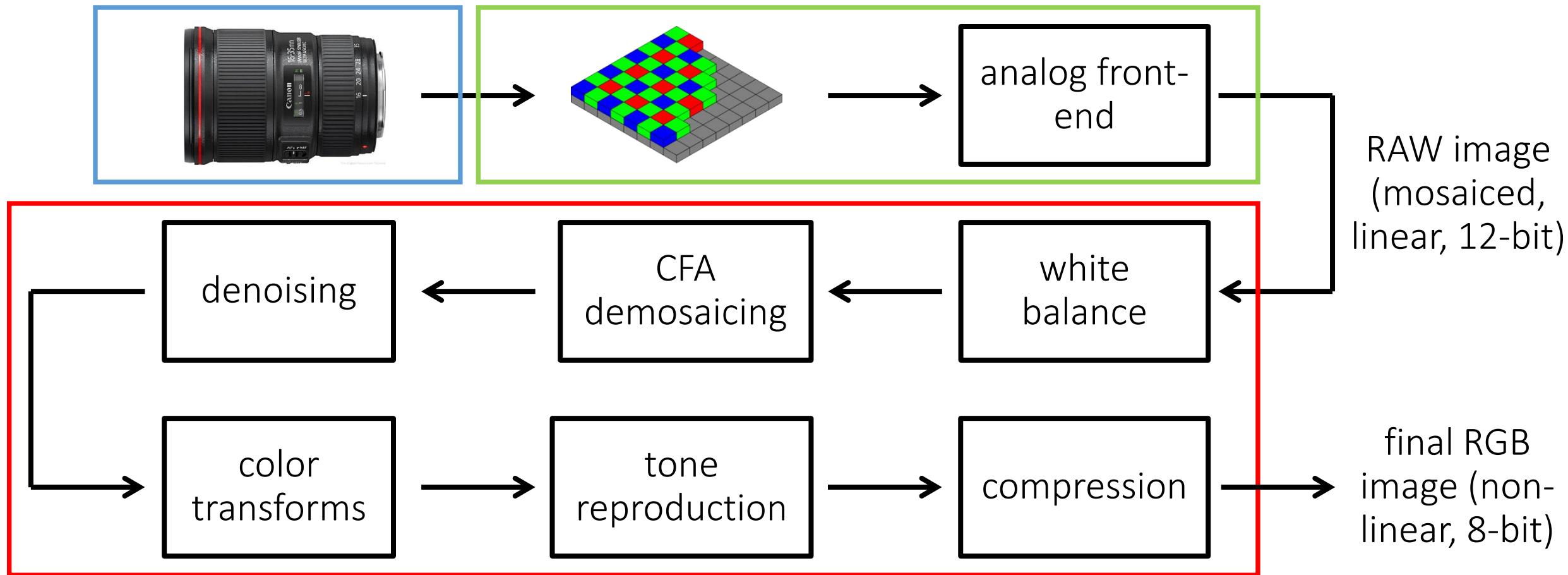


Calibration chart can be used for:

1. color calibration
2. radiometric calibration (i.e., response curve) using the bottom row

# The image processing pipeline

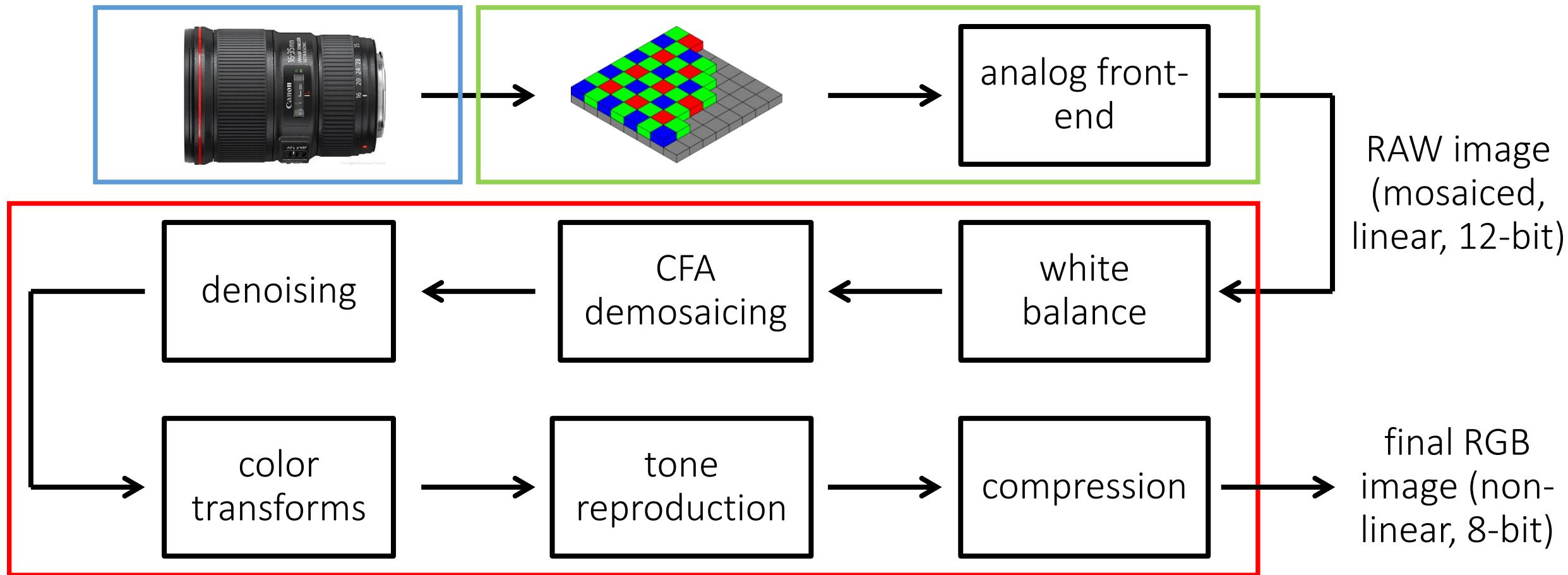
Which part of the pipeline does the non-linear response curve correspond to?



# The image processing pipeline

Which part of the pipeline does the non-linear response curve correspond to?

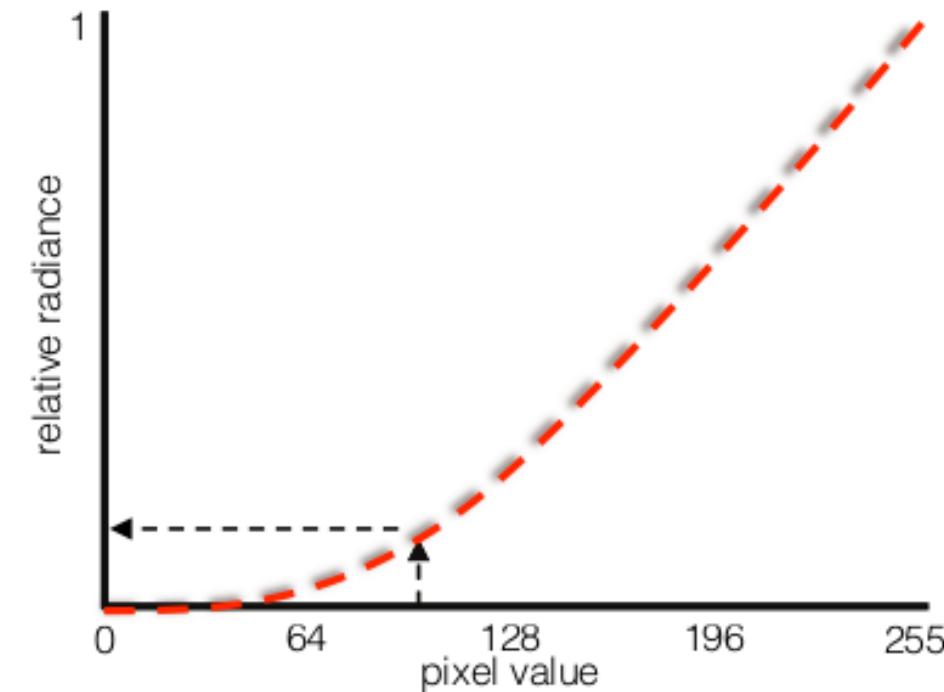
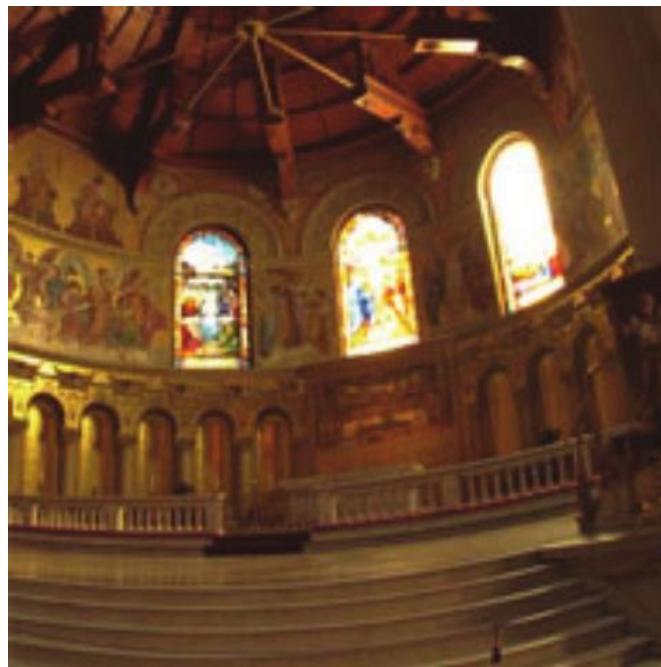
- The tone reproduction (mostly).



# Non-linear image formation model

Real scene radiance for image pixel  $(x, y)$ :  $L(x, y)$

Exposure time:  $t_i$



$$I_{\text{linear}}(x, y) = \text{clip}[ t_i \cdot L(x, y) + \text{noise} ]$$

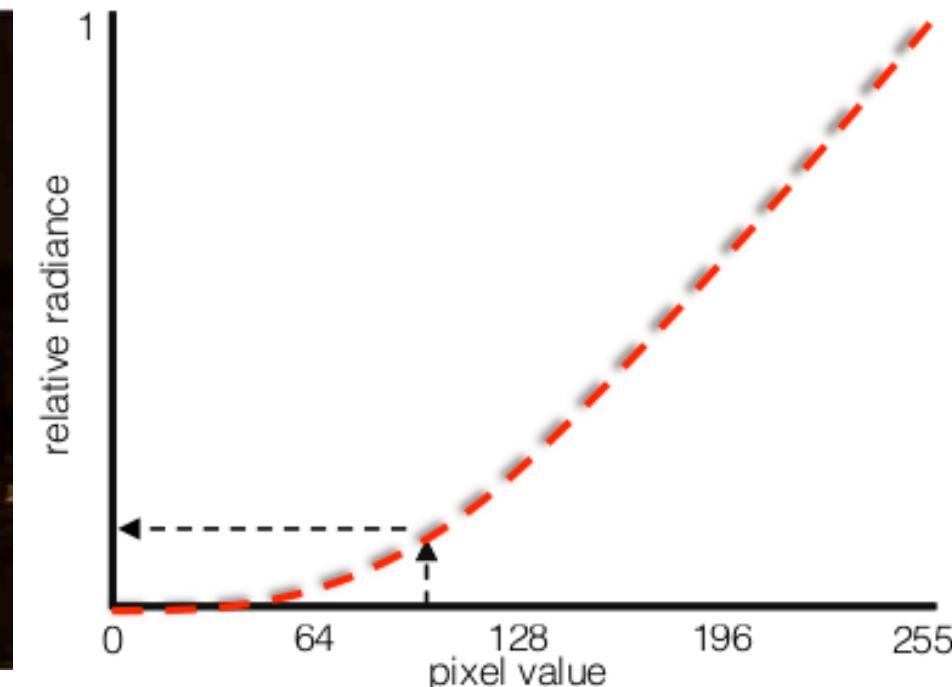
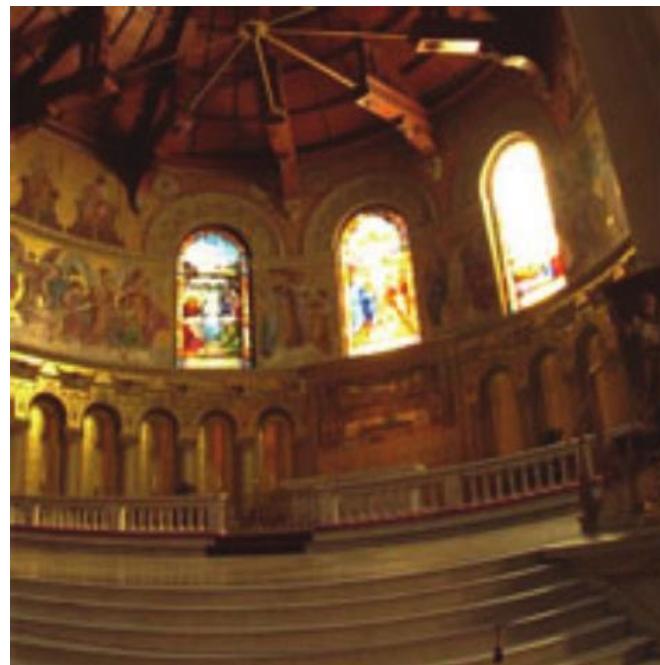
$$I_{\text{non-linear}}(x, y) = f[ I_{\text{linear}}(x, y) ]$$

How would you merge the non-linear images into an HDR one?

# Non-linear image formation model

Real scene radiance for image pixel  $(x, y)$ :  $L(x, y)$

Exposure time:  $t_i$



$$I_{\text{linear}}(x, y) = \text{clip}[ t_i \cdot L(x, y) + \text{noise} ]$$

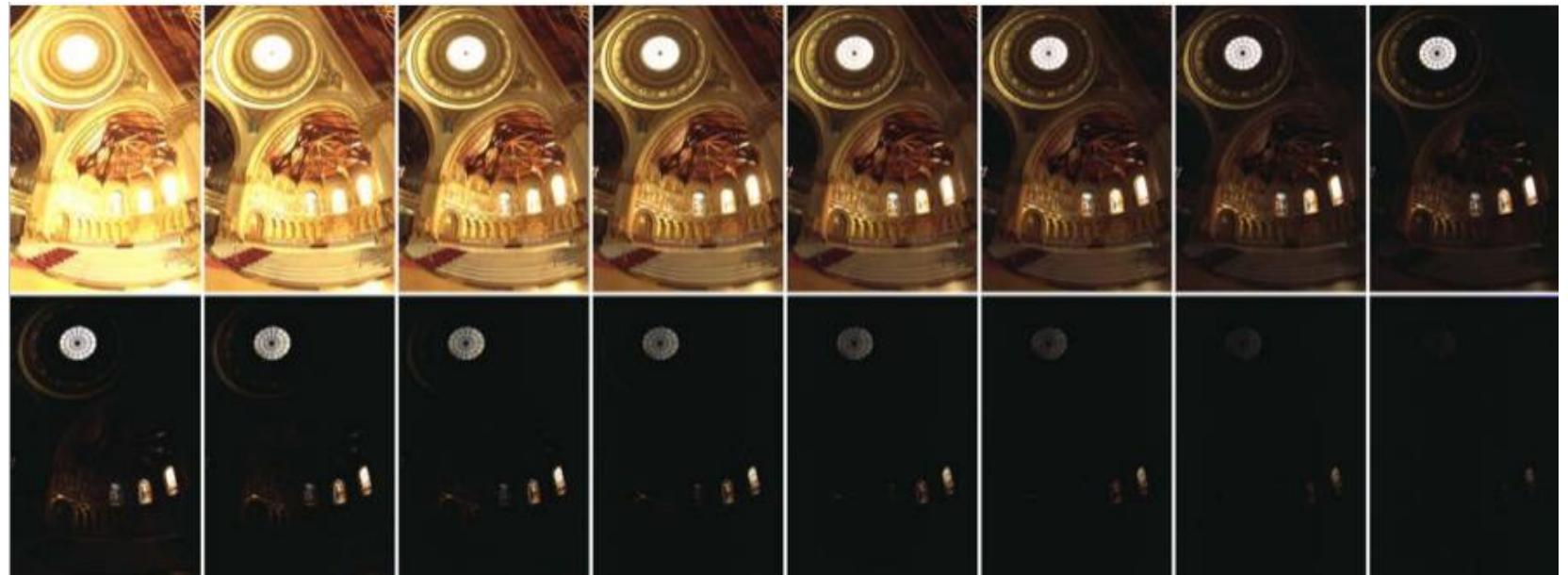
$$I_{\text{non-linear}}(x, y) = f[ I_{\text{linear}}(x, y) ]$$

$$I_{\text{est}}(x, y) = f^{-1}[ I_{\text{non-linear}}(x, y) ]$$

Use inverse transform to estimate linear image, then proceed as before

# Linearization

$$I_{\text{non-linear}}(x, y) = f[I_{\text{linear}}(x, y)]$$



$$I_{\text{est}}(x, y) = f^{-1}[I_{\text{non-linear}}(x, y)]$$



# Merging non-linear exposure stacks

1. Calibrate response curve
2. Linearize images

For each pixel:

3. Find “valid” images  $\leftarrow$  (noise)  $0.05 < \text{pixel} < 0.95$  (clipping)
4. Weight valid pixel values appropriately  $\leftarrow$   $(\text{pixel value}) / t_i$
5. Form a new pixel value as the weighted average of valid pixel values

Note: many possible weighting schemes

What if I cannot measure response curve?

# Tone reproduction curves

The exact tone reproduction curve depends on the camera.

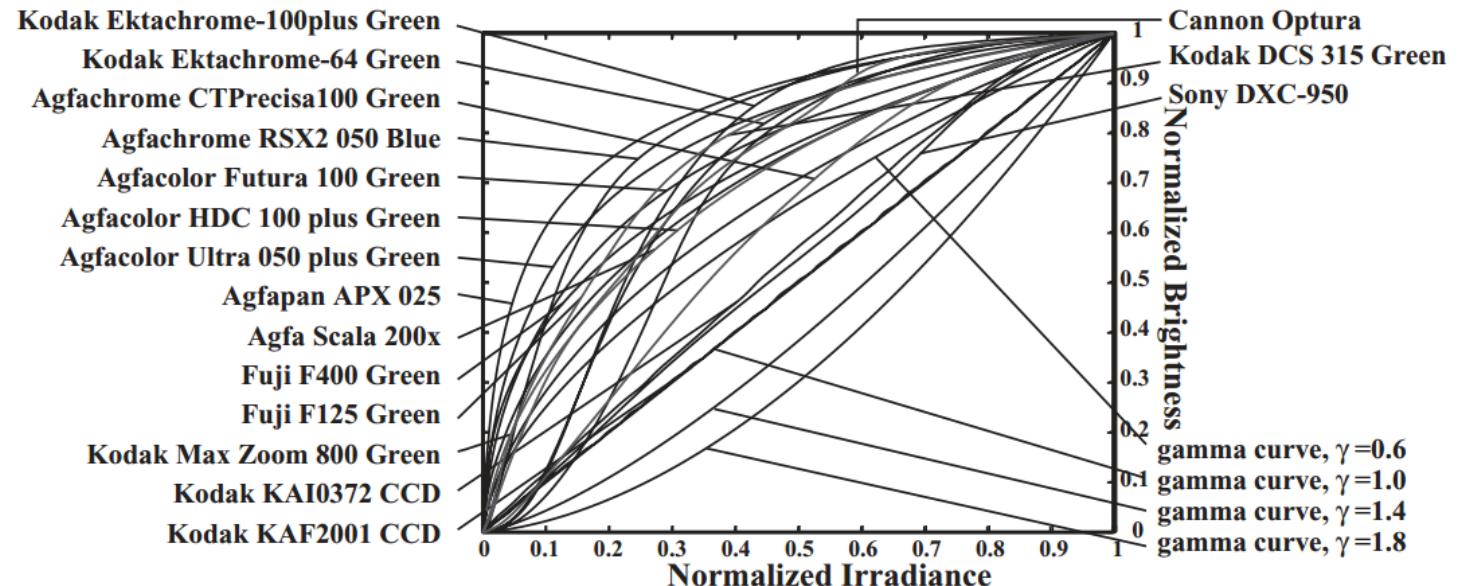
- Often well approximated as  $L^\gamma$ , for different values of the power  $\gamma$  (“gamma”).
- A good default is  $\gamma = 1 / 2.2$ .



before gamma



after gamma

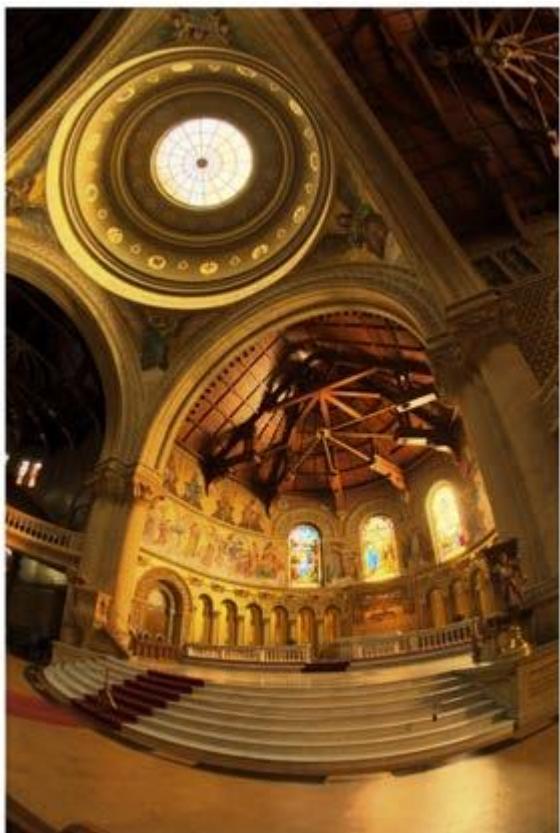


If nothing else, take the square of your image to approximately remove effect of tone reproduction curve.

# Relative vs absolute radiance

Final fused HDR image gives radiance only up to a global scale

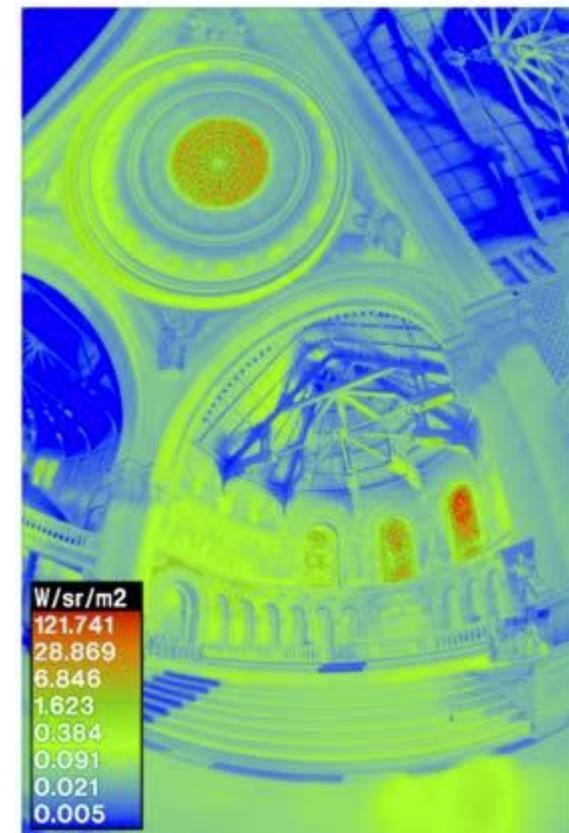
- If we know exact radiance at one point, we can convert relative HDR image to absolute radiance map



HDR image  
(relative radiance)



spotmeter (absolute  
radiance at one point)



absolute  
radiance map

# Basic HDR approach

1. Capture multiple LDR images at different exposures
2. Merge them into a single HDR image

Any problems with this approach?

# Basic HDR approach

1. Capture multiple LDR images at different exposures
2. Merge them into a single HDR image

Problem: Very sensitive to movement

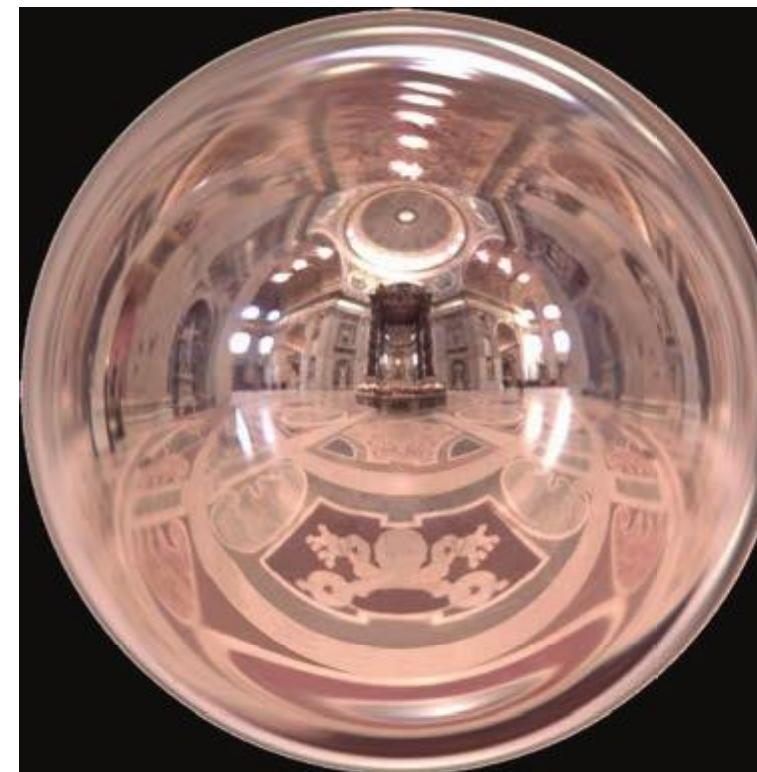
- Scene must be completely static
- Camera must not move

Most modern automatic HDR solutions include an alignment step before merging exposures

# Another type of HDR images

Light probes: place a chrome sphere in the scene and capture an HDR image

- Used to measure real-world illumination environments (“environment maps”)

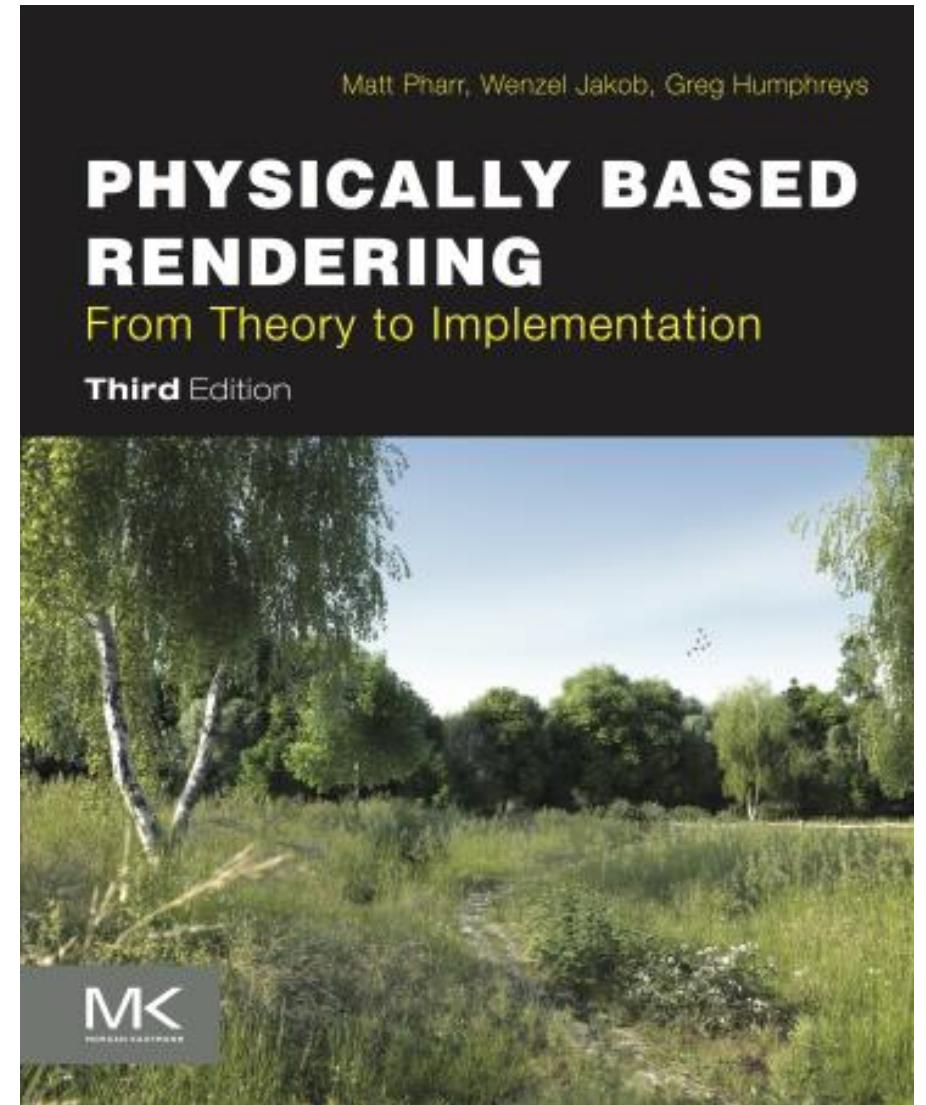


Application: image-based relighting  
(later lecture)

# Another way to create HDR images

Physics-based renderers simulate radiance maps  
(relative or absolute)

- Their outputs are very often HDR images



# A note about HDR today

- Most cameras (even phone cameras) have automatic HDR modes/apps
- Popular-enough feature that phone manufacturers are actively competing about which one has the best HDR
- The technology behind some of those apps (e.g., Google's HDR+) is published in SIGGRAPH and SIGGRAPH Asia conferences

## Burst photography for high dynamic range and low-light imaging on mobile cameras

Samuel W. Hasinoff  
Jonathan T. Barron

Dillon Sharlet  
Florian Kainz  
Google Research

Ryan Geiss  
Jiawen Chen  
Andrew Adams  
Marc Levoy



**Figure 1:** A comparison of a conventional camera pipeline (left, middle) and our burst photography pipeline (right) running on the same cell-phone camera. In this low-light setting (about 0.7 lux), the conventional camera pipeline underexposes (left). Brightening the image (middle) reveals heavy spatial denoising, which results in loss of detail and an unpleasantly blotchy appearance. Fusing a burst of images increases the signal-to-noise ratio, making aggressive spatial denoising unnecessary. We encourage the reader to zoom in. While our pipeline excels in low-light and high-dynamic-range scenes (for an example of the latter see figure 10), it is computationally efficient and reliably artifact-free, so it can be deployed on a mobile camera and used as a substitute for the conventional pipeline in almost all circumstances. For readability the figure has been made uniformly brighter than the original photographs.

### Abstract

Cell phone cameras have small apertures, which limits the number of photons they can gather, leading to noisy images in low light. They also have small sensor pixels, which limits the number of electrons each pixel can store, leading to limited dynamic range. We describe a computational photography pipeline that captures, aligns, and merges a burst of frames to reduce noise and increase dynamic range. Our system has several key features that help make it robust and efficient. First, we do not use bracketed exposures. Instead, we capture frames of constant exposure, which makes alignment more robust, and we set this exposure low enough to avoid blowing out highlights. The resulting merged image has clean shadows and high bit depth, allowing us to apply standard HDR tone mapping methods. Second, we begin from Bayer raw frames rather than the demosaiced RGB (or YUV) frames produced by hardware Image Signal Processors (ISPs) common on mobile platforms. This gives us more bits per pixel and allows us to circumvent the ISP's unwanted tone mapping and spatial denoising. Third, we use a novel FFT-based alignment algorithm and a hybrid 2D/3D Wiener filter to denoise and merge the frames in a burst. Our implementation is built atop Android's Camera2 API, which provides per-frame camera control and access to raw imagery, and is written in the HALide domain-specific language (DSL). It runs in 4 seconds on device (for a 12 Mpix image), requires no user intervention, and ships on several mass-produced cell phones.

**Keywords:** computational photography, high dynamic range

**Concepts:** •Computing methodologies → Computational photography; Image processing;

### 1 Introduction

The main technical impediment to better photographs is lack of light. In indoor or night-time shots, the scene as a whole may provide insufficient light. The standard solution is either to apply analog or digital gain, which amplifies noise, or to lengthen exposure time, which causes motion blur due to camera shake or subject motion. Surprisingly, daytime shots with high dynamic range may also suffer from lack of light. In particular, if exposure time is reduced to avoid

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SA '16 Technical Papers, December 05 - 08, 2016, Macao  
ISBN: 978-1-4503-4514-9/16/12  
DOI: <http://dx.doi.org/10.1145/2980179.2980254>

Color calibration  
(a.k.a., measuring your camera's color space)

# Many different spectral sensitivity functions

Each camera has its more or less unique, and most of the time *secret*, SSF.

- Makes it very difficult to correctly reproduce the color of sensor measurements.



Images of the same scene captured using 3 different cameras with identical sRGB settings.

# Linear color spaces

basis for retinal color  $\Leftrightarrow$  color matching functions  $\Leftrightarrow$  primary colors  $\Leftrightarrow$  color space

$$\begin{bmatrix} \mathbf{c}(\ell(\lambda)) \\ \vdots \end{bmatrix} = \underbrace{\begin{bmatrix} \mathbf{c}_1 & \mathbf{c}_2 & \mathbf{c}_3 \\ \vdots & \vdots & \vdots \end{bmatrix}}_{\mathbf{M}^{-1}} \begin{bmatrix} \int k_1(\lambda) \ell(\lambda) d\lambda \\ \int k_2(\lambda) \ell(\lambda) d\lambda \\ \int k_3(\lambda) \ell(\lambda) d\lambda \end{bmatrix}$$
$$\begin{bmatrix} \mathbf{c}(\ell_{435}) & \mathbf{c}(\ell_{545}) & \mathbf{c}(\ell_{625}) \\ \vdots & \vdots & \vdots \end{bmatrix} \mathbf{M}^{-1} \quad \begin{bmatrix} k_1(\lambda) \\ k_2(\lambda) \\ k_3(\lambda) \end{bmatrix} = \mathbf{M} \begin{bmatrix} k_{435}(\lambda) \\ k_{545}(\lambda) \\ k_{625}(\lambda) \end{bmatrix}$$

$$\begin{bmatrix} \mathbf{c}(\ell(\lambda)) \\ \vdots \end{bmatrix} = \begin{bmatrix} \mathbf{c}(\ell_{435}) & \mathbf{c}(\ell_{545}) & \mathbf{c}(\ell_{625}) \\ \vdots & \vdots & \vdots \end{bmatrix} \begin{bmatrix} \int k_{435}(\lambda) \ell(\lambda) d\lambda \\ \int k_{535}(\lambda) \ell(\lambda) d\lambda \\ \int k_{625}(\lambda) \ell(\lambda) d\lambda \end{bmatrix}$$

representation of retinal color in LMS space

change of basis matrix

$\mathbf{M}^{-1}\mathbf{M}$  can insert any invertible  $\mathbf{M}$

representation of retinal color in space of primaries

# Linear color spaces

Change of color space:

$$c' = H \cdot c$$

The diagram illustrates the mathematical equation  $c' = H \cdot c$  for color space transformation. The equation is centered at the top. Below it, two arrows point upwards from the text labels to the variables  $c'$  and  $c$  respectively. The left arrow originates from the text 'desired reference color space (i.e., XYZ)'. The right arrow originates from the text 'camera color space'.

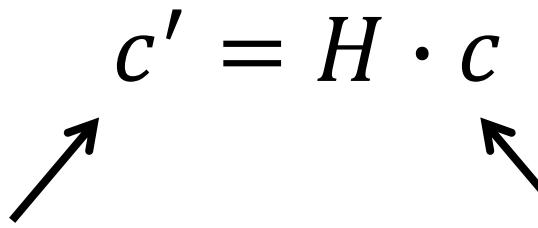
desired reference color space (i.e., XYZ)

camera color space

What does this look like?

# Linear color spaces

Change of color space:

$$c' = H \cdot c$$


desired reference color space (i.e., XYZ)

camera color space

What does this look like?

- It's a homography!

How do we compute homographies?

# Linear color spaces

Change of color space:

$$c' = H \cdot c$$


desired reference color space (i.e., XYZ)

camera color space

What does this look like?

- It's a homography!

How do we compute homographies?

- We use SVD and the DLT!

How many colors do we need to match?

# Linear color spaces

Change of color space:

$$c' = H \cdot c$$


desired reference color space (i.e., XYZ)

camera color space

What does this look like?

- It's a homography!

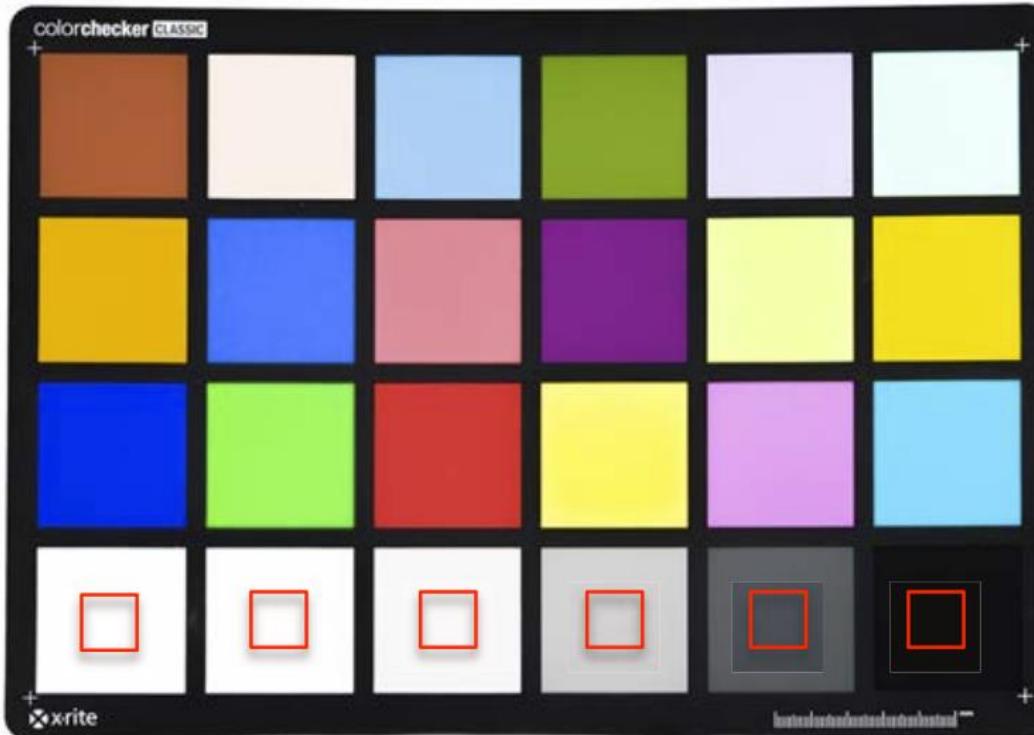
How do we compute homographies?

- We use SVD and the DLT!

How many colors do we need to match?

- We need at least four colors.

# Using (again) a color chart



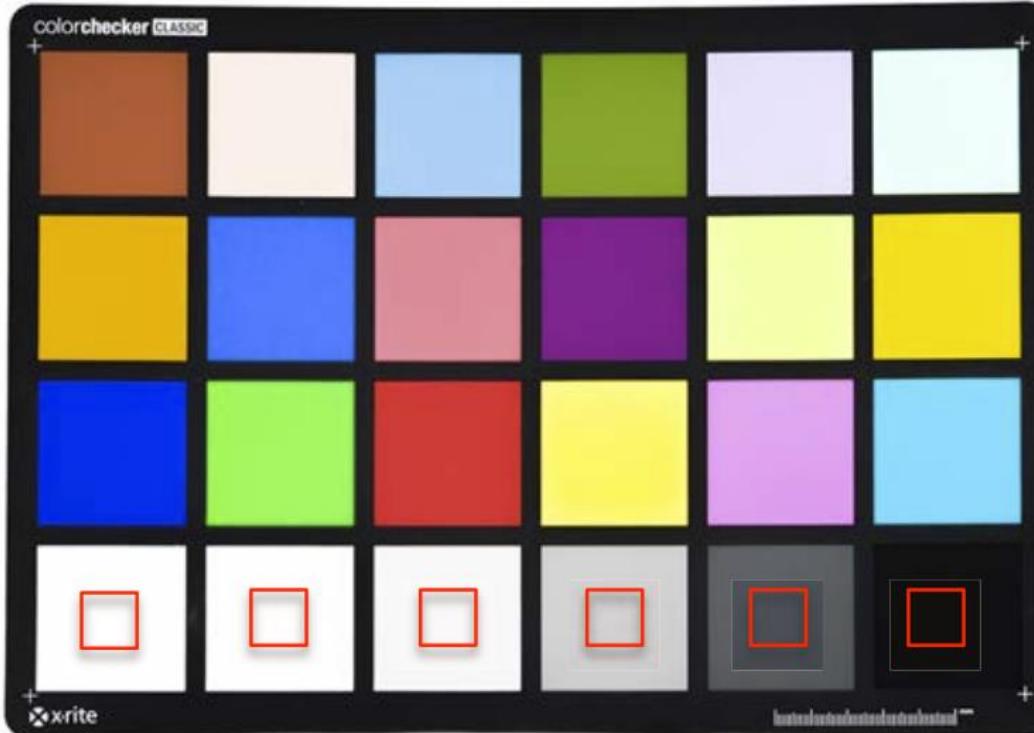
Color patches manufactured to have pre-calibrated XYZ coordinates.

Can we use any color chart image for color calibration?

Calibration chart can be used for:

1. color calibration
2. radiometric calibration (i.e., response curve) using the bottom row

# Using (again) a color chart



Color patches manufactured to have pre-calibrated XYZ coordinates.

Can we use any color chart image for color calibration?

- It needs to be a *linear* image!
- Do radiometric calibration first.

Calibration chart can be used for:

1. color calibration
2. radiometric calibration (i.e., response curve) using the bottom row

# An example



original

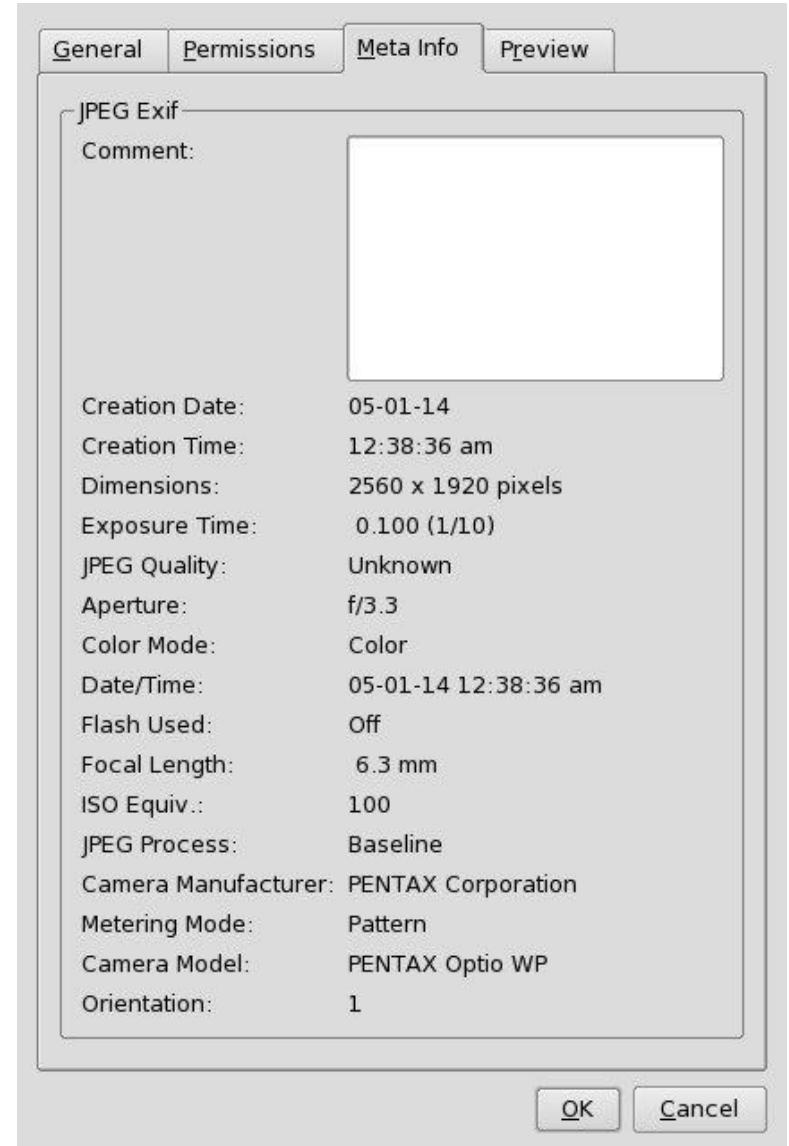


color-corrected

# Quick note

If you cannot do calibration, take a look at the image's EXIF data (if available).

Often contains information about tone reproduction curve and color space.



# Take-home messages

The values of pixels in a photograph and the values output by your camera's sensor are two very different things.

The relationship between the two is complicated and unknown, and we often need to account for it when doing computer vision.

# References

Basic reading:

- Szeliski textbook, Section 2.3
- Michael Brown, “Understanding the In-Camera Image Processing Pipeline for Computer Vision,” CVPR 2016, very detailed discussion of issues relating to color photography and management, slides available at: [http://www.comp.nus.edu.sg/~brown/CVPR2016\\_Brown.html](http://www.comp.nus.edu.sg/~brown/CVPR2016_Brown.html)
- Nine Degrees Below, <https://ninedegreesbelow.com/> amazing resource for color photography, reproduction, and management.