Radiosity

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Acknowledgements

- Tom Funkhouser
- Dutré et al. 2003, Chapter 6
- Keller, "Instant Radiosity," SIGGRAPH 97.

Radiosity

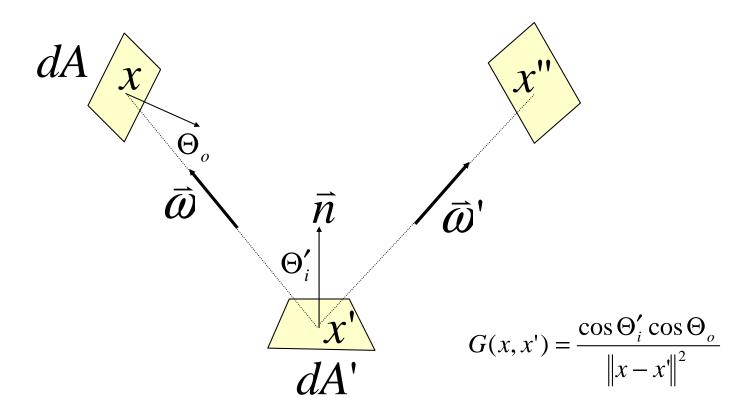


Overview

- Radiosity equation
- Solution methods
 - Computing form factors
 - Selecting basis functions for radiosities
 - Solving linear system of equations
 - Meshing surfaces into elements
 - Rendering images
- Discussion

Rendering Equation

$$L(x' \to x'') = L_e(x' \to x'') + \int_S f_r(x \to x' \to x'') L(x \to x') V(x, x') G(x, x') dA$$



Kajiya 1986

Radiosity Equation

$$L(x' \to x'') = L_e(x' \to x'') + \int_S f_r(x \to x' \to x'') L(x \to x') V(x, x') G(x, x') dA$$

Assume everything is Lambertian

$$\rho(x') = f_r(x \to x' \to x'')\pi$$

$$L(x') = L_e(x') + \frac{\rho(x')}{\pi} \int_{S} L(x)V(x, x')G(x, x')dA$$

Convert to Radiosities

$$B = \int_{\Omega} L_o \cos \theta d\omega \qquad L = \frac{B}{\pi}$$

$$L = \frac{B}{\pi}$$

$$B(x') = B_e(x') + \frac{\rho(x')}{\pi} \int_{S} B(x)V(x, x')G(x, x')dA$$

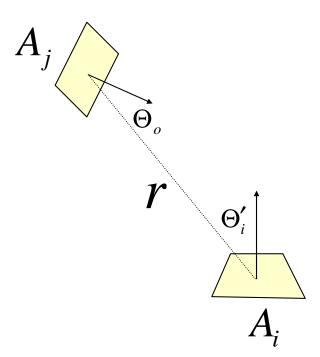
Radiosity Approximation

$$B(x') = B_e(x') + \frac{\rho(x')}{\pi} \int_{S} B(x)V(x, x')G(x, x')dA$$

Discretize the surfaces into "elements"

$$B_i = E_i + \rho_i \sum_{j=1}^{N} B_j F_{ij}$$

where
$$F_{ij} = \frac{1}{A_i} \int_{A_i} \int_{A_i} \frac{V_{ij} \cos \Theta_i' \cos \Theta_o}{\pi r^2} dA_j dA_i$$



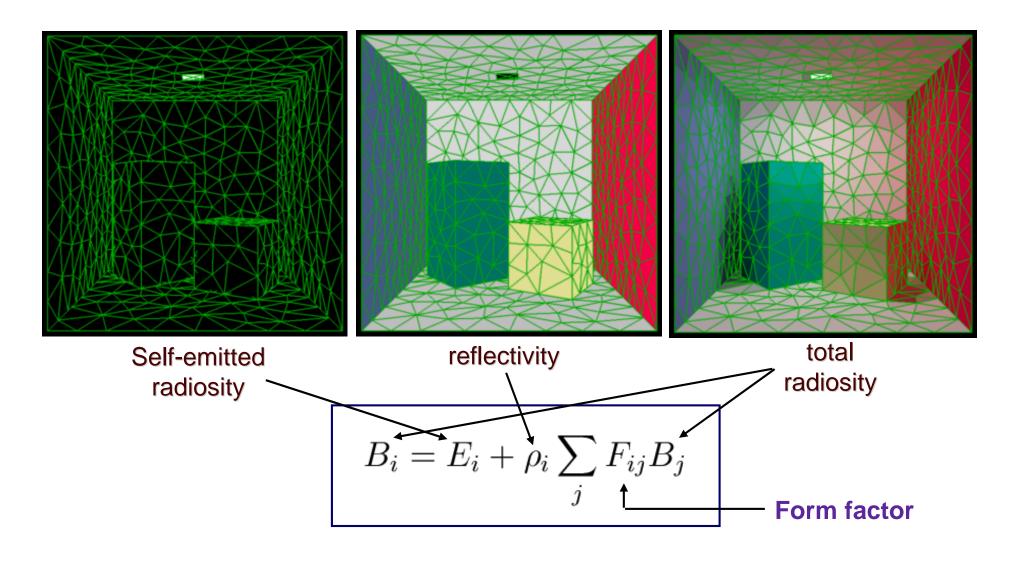


Fig 6.1, [Dutré et al. 2003]

System of Equations

$$B_i = E_i + \rho_i \sum_{j=1}^N B_j F_{ij}$$

$$E_i = B_i - \rho_i \sum_{j=1}^{N} B_j F_{ij}$$

$$B_i - \rho_i \sum_{i=1}^N B_j F_{ij} = E_i$$

$$B_{i} = E_{i} + \rho_{i} \sum_{j=1}^{N} B_{j} F_{ij}$$

$$E_{i} = B_{i} - \rho_{i} \sum_{j=1}^{N} B_{j} F_{ij}$$

$$B_{i} - \rho_{i} \sum_{j=1}^{N} B_{j} F_{ij} = E_{i}$$

$$\begin{bmatrix}
1 - \rho F & . & . & . & -\rho F_{n} \\
-\rho F & 1 - \rho F & . & . & -\rho F_{n} \\
. & . & . & . & . & . \\
-\rho_{n-1} F_{n-1,1} & . & . & . & . & -\rho_{n-1} F_{n-1,n} \\
-\rho_{n} F_{n,1} & . & . & . & 1 - \rho_{n} F_{n,n}
\end{bmatrix} \begin{bmatrix}
B_{1} \\
B_{2} \\
. \\
. \\
B_{n}
\end{bmatrix} \begin{bmatrix}
E_{1} \\
E_{2} \\
. \\
. \\
E_{n}
\end{bmatrix}$$

Intuition

$$B_i = E_i + \rho_i \sum_{j=1}^N B_j F_{ij}$$

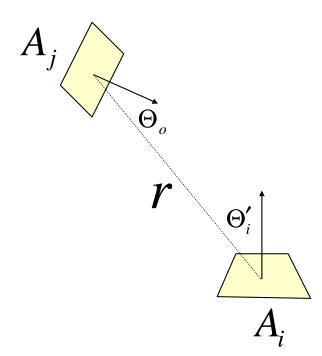
Multiply by A_i

$$B_i A_i = E_i A_i + \rho_i \sum_{j=1}^N B_j F_{ij} A_i$$

$$F_{ij}A_i = F_{ji}A_j$$

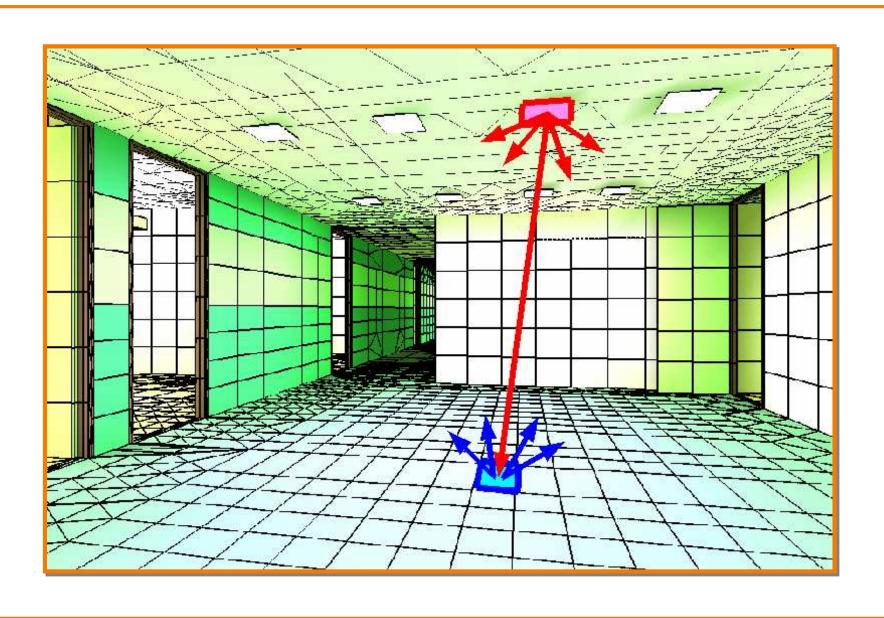
$$B_i A_i = E_i A_i + \rho_i \sum_{j=1}^N B_i F_{ji} A_j$$

$$B_i A_i = E_i A_i + \rho_i \sum_{j=1}^{N} F_{ji} B_j A_j \quad \leftarrow \text{energy balance}$$
equation



This is an equation

Intuition

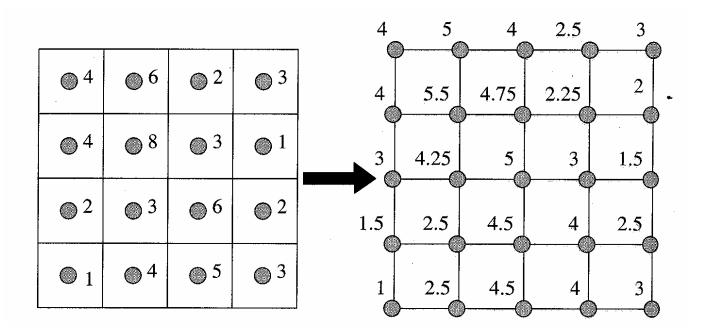


Overview

- Radiosity equation
- Solution methods
 - Computing form factors
 - Selecting basis functions for radiosities
 - Solving linear system of equations
 - Meshing surfaces into elements
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Displaying Radiosity

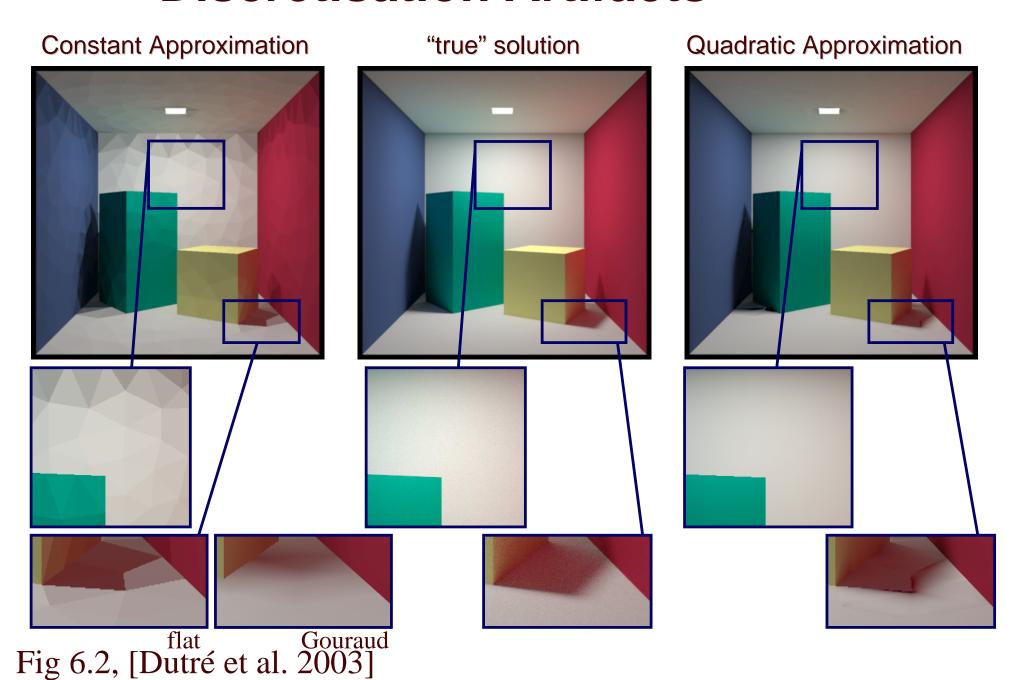
• Usually, Gouraud shading ...



computed

displayed

Discretisation Artifacts

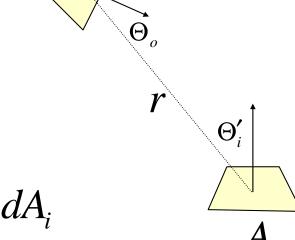


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Form Factor

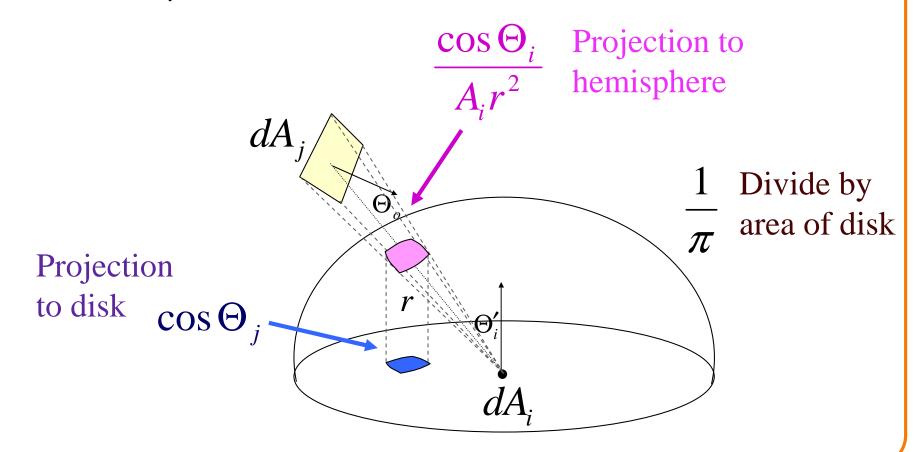
 Fraction of energy leaving element i that arrives at element j



$$F_{ij} = \frac{1}{A_i} \int_{A_i} \int_{A_i} \frac{V_{ij} \cos \Theta_i' \cos \Theta_o}{\pi r^2} dA_j dA_i$$

Form Factor Intuition

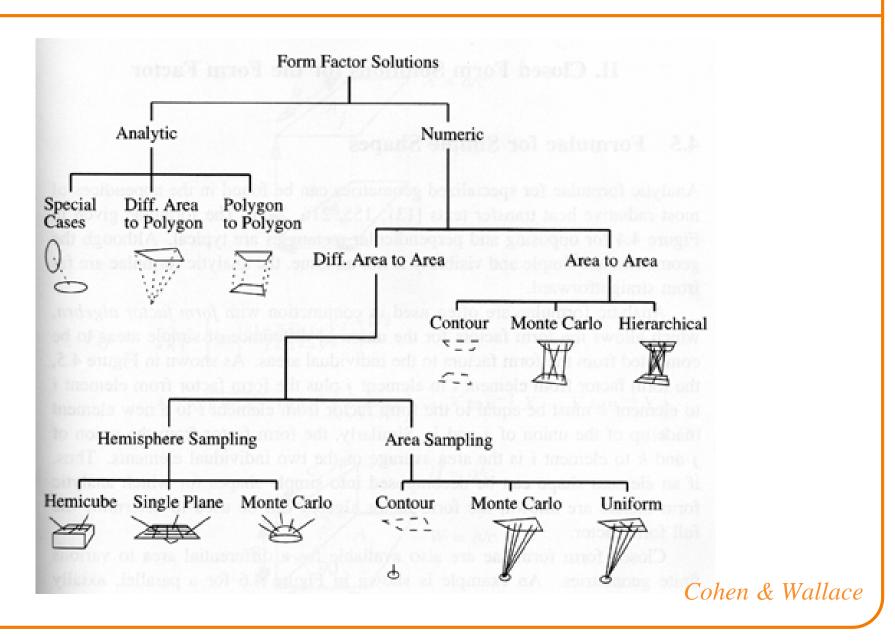
$$F_{di-dj} = \frac{1}{A_i} \frac{V_{ij} \cos \Theta_i \cos \Theta_j}{\pi r^2}$$



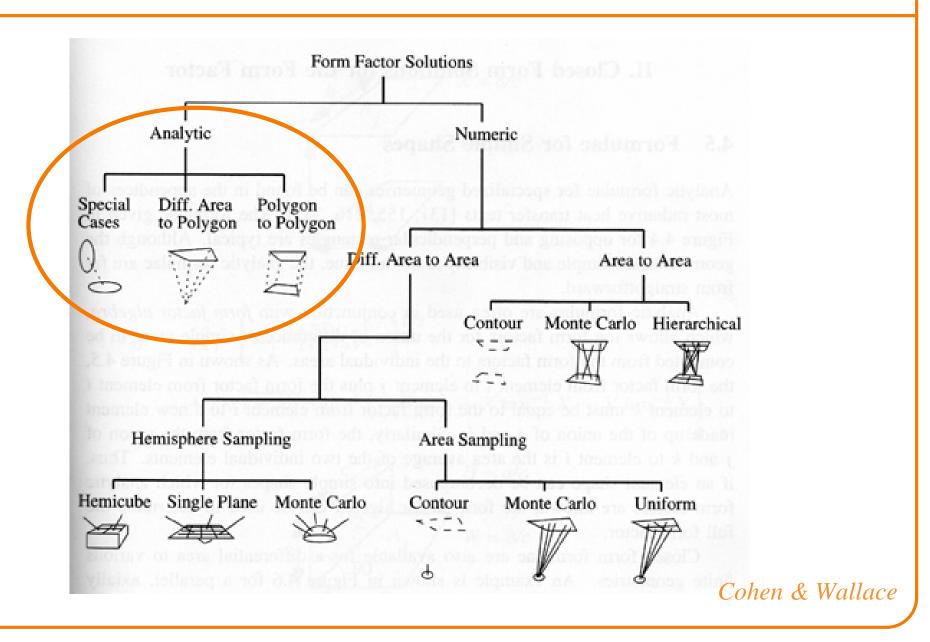
$F_{ij} = \frac{1}{A_i} \int_{S_i} \int_{S_j} G(x, y) dA_y dA_x$ Form Factor Singularities and $G(x,y) = \frac{\cos \theta_x \cos \theta_y}{\pi r_{xy}^2} \text{vis}(x,y).$ **Discontinuities DISTANCE OCCLUDED TENDS TO ZERO VISIBLE** DISCONTINUITIES

Fig 6.3, [Dutré et al. 2003]

Computing Form Factors

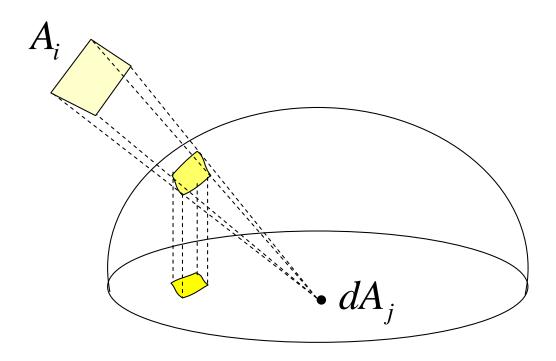


Computing Form Factors



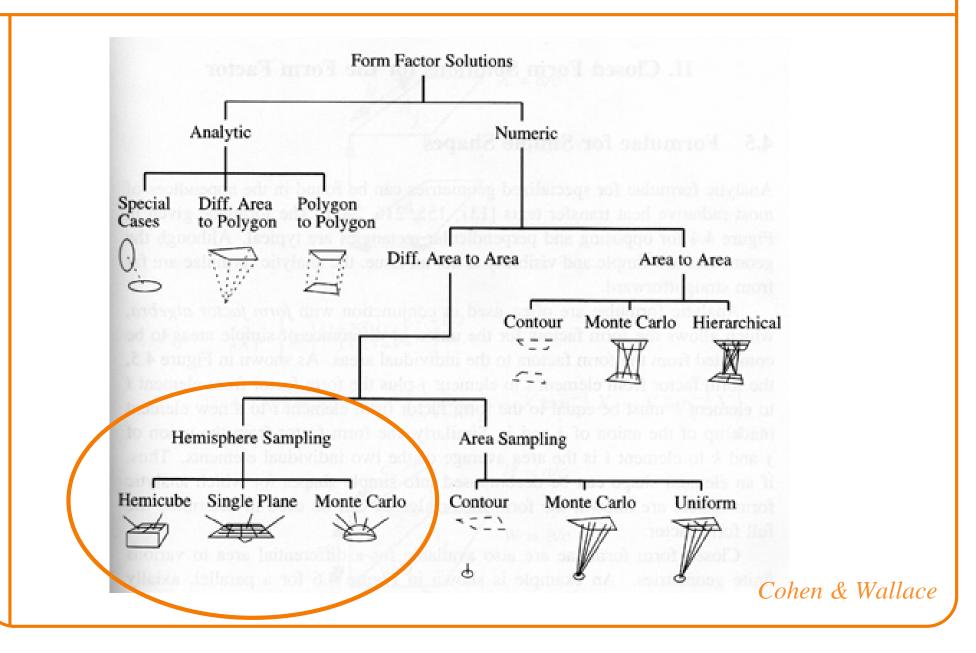
Analytic Form Factors

- Derive equation for projected area
 - Possible only for simple situations
 e.g., direct illumination of a point from a visible polygon



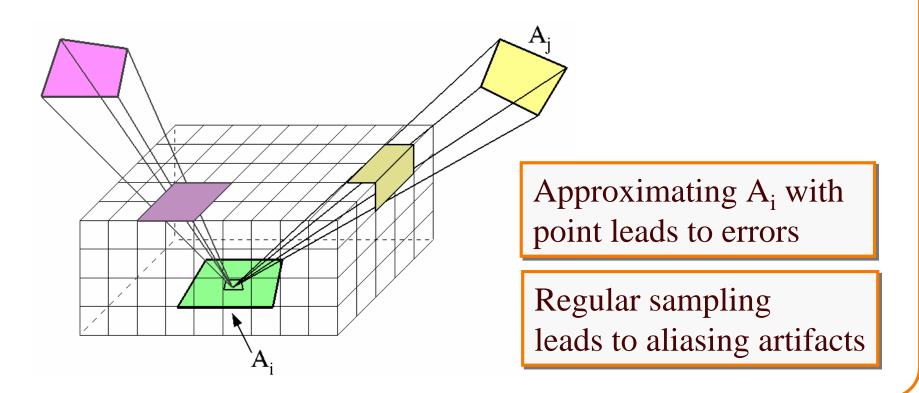
Partial visibility is problematic

Computing Form Factors

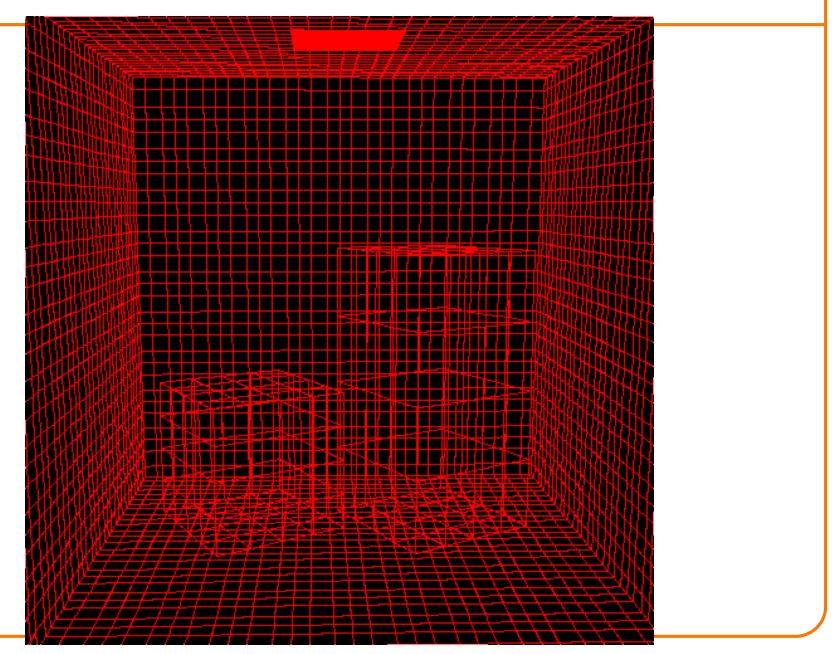


Hemicube [Cohen & Greenberg, 85]

- Compute form factor with image-space precision
 - Render scene from centroid of A_i
 - Use z-buffer to determine visibility of other surfaces
 - Count "pixels" to determine projected areas



Ex: Wireframe



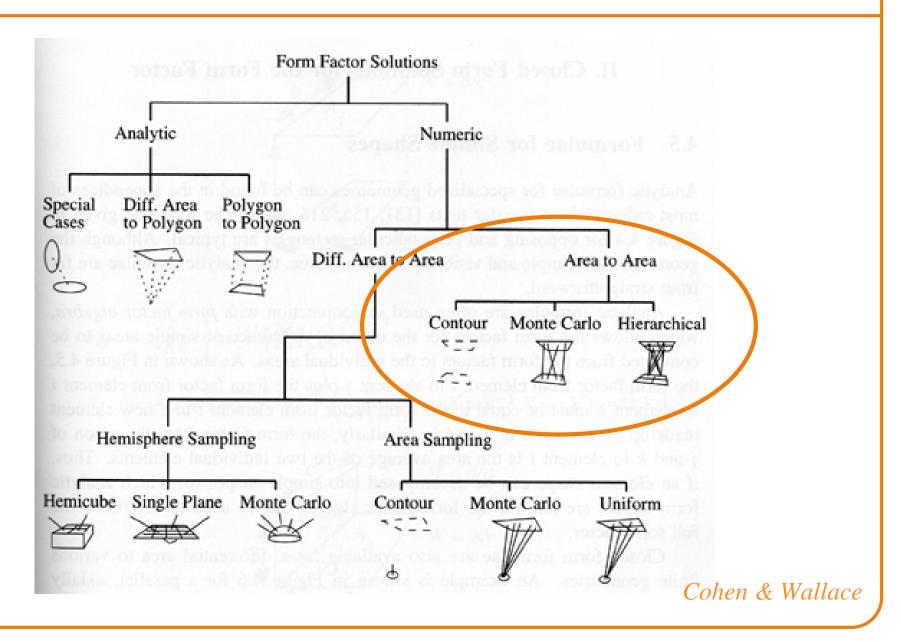
Ex: Classical HC (no interpolation)



Ex: Antialiased HC e.g. [Chandran et al. 2000]

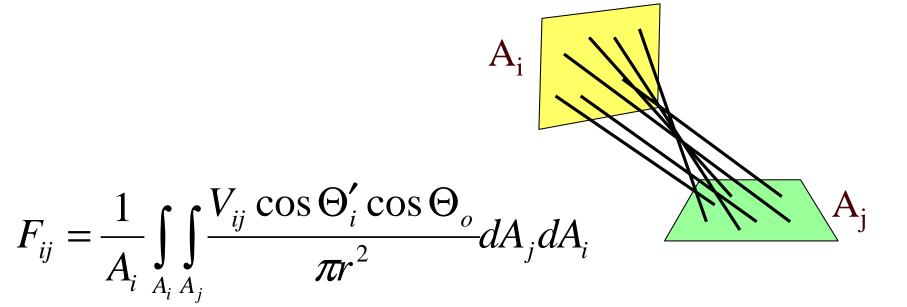


Computing Form Factors



Monte Carlo Sampling

- Compute form factor by random sampling
 - Select random points on elements
 - Intersect line segment to evaluate V_{ij}
 - Evaluate F_{ij} by Monte Carlo integration



Other approaches: Global Lines

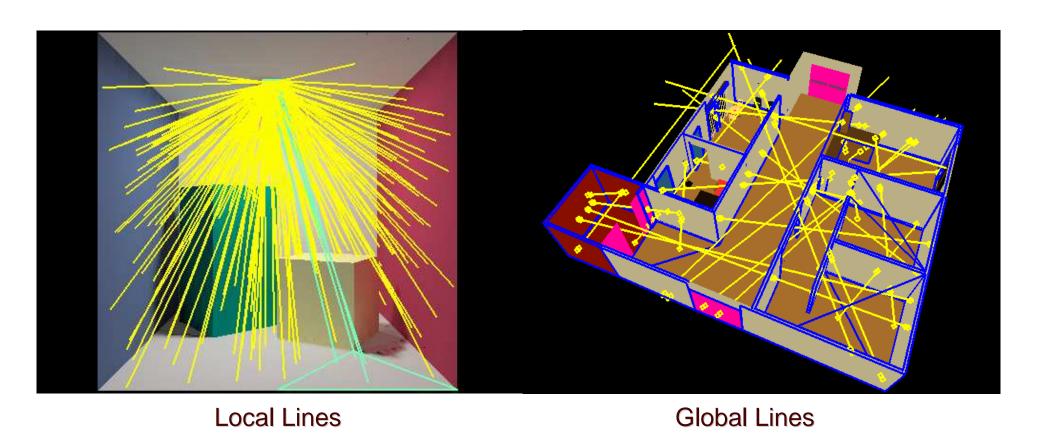


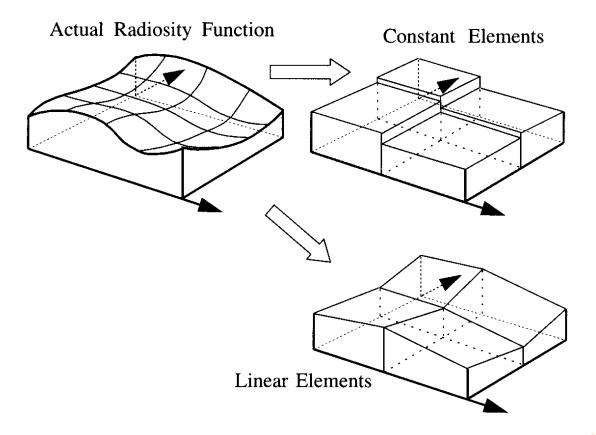
Fig 6.4, [Dutré et al. 2003]

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Selecting a Basis Function

- Store radiosity function on surface mesh
 - Piecewise-constant, piecewise-linear, wavelets, etc.



Cohen & Wallace

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Solving the System of Equations

Challenges:

- Size of matrix
- Cost of computing form factors
- Computational complexity

Α

x = b

Solving the System of Equations

Solution methods:

- Invert the matrix O(n³)
- ∘ Iterative methods O(n²)
- Hierarchical methods O(n)

Α

x = b

Gauss-Seidel Iteration

```
1 for all i

2 B_i = E_i

3 while not converged

4 for each i in turn

5 B_i = E_i + \rho_i \sum_{j \neq i} B_j F_{ij}

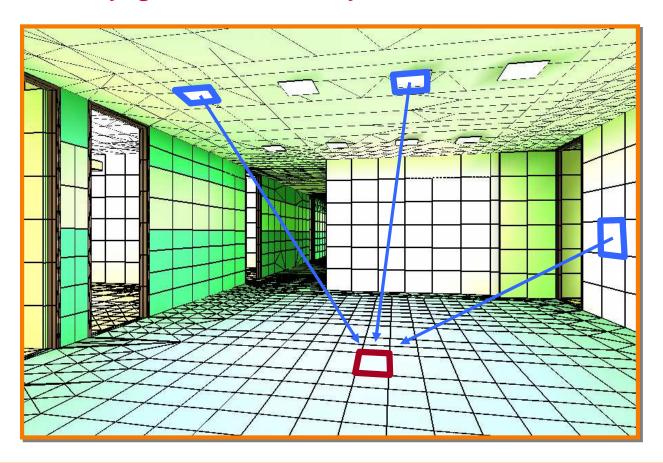
6 display the image using B_i as the intensity of patch i.
```

Gauss-Seidel Iteration

- Two interpretations:
 - Iteratively relax rows of linear system
 - Iteratively gather radiosity to elements

Gauss-Seidel Iteration

- Two interpretations:
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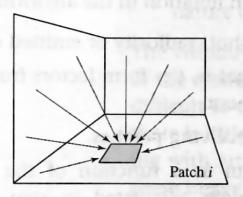
- Cohen et al., SIGGRAPH 1988
- Shoot light instead of gathering light
- Each iteration is O(n)
- May or may not keep F_{ij} after each iteration

Gathering & Shooting Radiosity

Gathering: a single iteration (k) updates a single patch i by gathering contributions from all other patches.

 $B_i^{(k+1)} = E_i + R_i \sum_{j=1}^{N} F_{ij} B_j^{(k)}$ $= \begin{bmatrix} + \end{bmatrix}$

Equivalent to gathering light energy from all the patches in the scene.

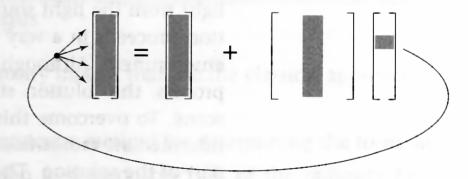


(a) Gathering

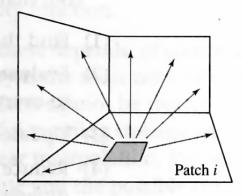
Shooting: a single step computes form factors from the shooting patch to all receiving patches and distributes (unshot) energy ΔB_i

for all
$$j$$
:

$$B_j^{(k+1)} = B_j^{(k)} + R_j F_{ji} \Delta B_i$$

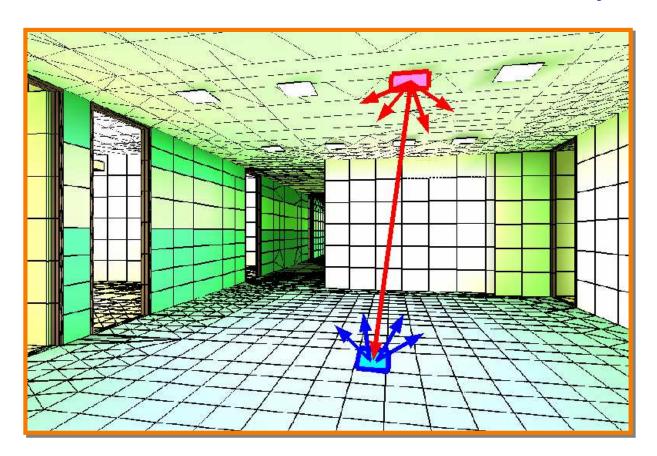


Equivalent to shooting light energy from a patch to all other patches in the scene.



(b) Shooting

- Interpretation:
 - Iteratively shoot "unshot" radiosity from elements
 - Select shooters in order of unshot radiosity

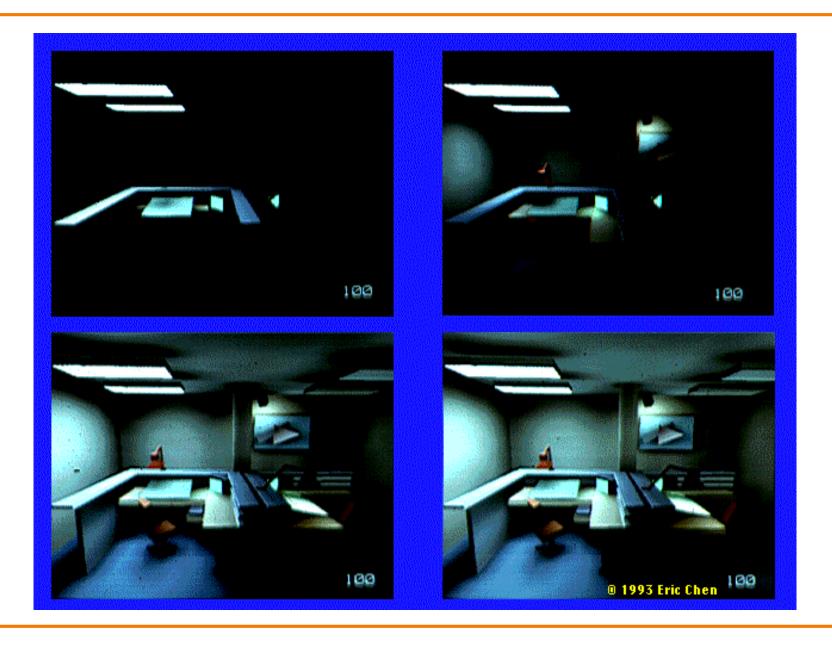


```
for all i
  B_i = E_i
3 \qquad \Delta B_i = E_i
   while not converged
5
     pick i, such that \Delta B_i * A_i is largest
      for every patch j
         \Delta rad = \Delta B_i * \rho_j F_{ji}
         \Delta B_i = \Delta B_i + \Delta rad
         B_i = B_i + \Delta rad
\Delta B_i = 0
11
      display the image using B_i as the intensity of patch i.
```



PROGRESSIVE SOLUTION

The above images show increasing levels of global diffuse illumination. From left to right: 0 bounces, 1 bounce, 3 bounces.





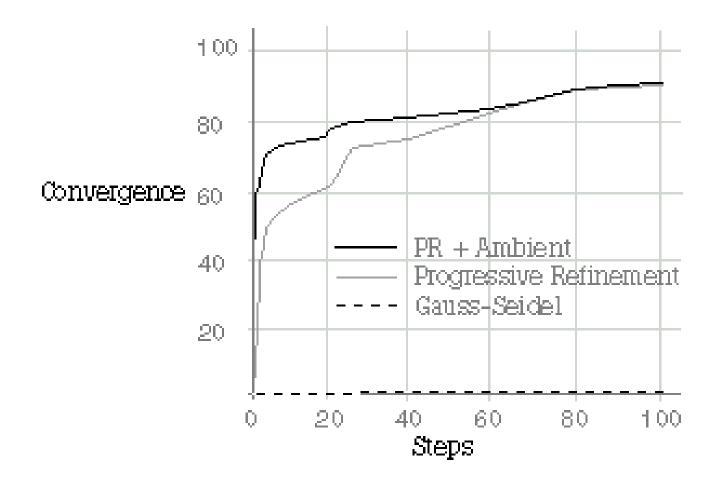
A radiosity image after 20, 250, and 5000 iterations of the progressive refinement method. From top to bottom for each column: (a) The radiosity solution as output from the iteration process. Each patch is allocated a constant radiosity. (b) The previous solution after it has been subjected to the interpolation process. (c) The same solution with the addition of the ambient term. (d) The difference between the previous two images. This gives a visual indication of the energy that had to be added to account for the unshot radiosity.

Effect of Ambient Light for Viewing



A radiosity image after 20, 250, and 5000 iterations of the progressive refinement method. From top to bottom for each column: (a) The radiosity solution as output from the iteration process. Each patch is allocated a constant radiosity. (b) The previous solution after it has been subjected to the interpolation process. (c) The same solution with the addition of the ambient term. (d) The difference between the previous two images. This gives a visual indication of the energy that had to be added to account for the unshot radiosity.

Adaptive refinement



PR: Some Special Cases

- Image after we have iterated through all light sources?
 - Shadows, but no interreflections
- Can incrementally display image while iterating
 - Add ambient light at each stage for visibility
 - Ambient shading if progressively refined
- Incremental form factor computation

Related: Stochastic Jacobi Radiosity

Algorithm 1 Incremental stochastic Jacobi iterative method.

- 1. Initialize total power $P_i \leftarrow P_{ei}$, unshot power $\Delta P_i \leftarrow P_{ei}$, and received power $\delta P_i \leftarrow 0$ for all patches i and compute total unshot power $\Delta P_T = \sum_i \Delta P_i$.
- 2. Until $\|\Delta P_i\| \leq \varepsilon$ or number of steps exceeds maximum, do
 - (a) Choose number of samples N.
 - (b) Generate a random number $\xi \in (0, 1)$.
 - (c) Initialize $N_{prev} \leftarrow 0$; $q \leftarrow 0$.
 - (d) Iterate over all patches i, for each i, do

i.
$$q_i \leftarrow \Delta P_i / \Delta P_T$$
.

ii.
$$q \leftarrow q + q_i$$
.

iii.
$$N_i \leftarrow |Nq + \xi| - N_{prev}$$
.

iv Do N_i times:

- A. Sample random point x on S_i .
- B. Sample cosine-distributed direction Θ at x.
- C. Determine patch j containing the nearest intersection point of the ray originating at x and with direction Θ , with the surfaces of the scene.
- D. Increment $\delta P_j \leftarrow \delta P_j + \frac{1}{N} \rho_j \Delta P_T$.

v.
$$N_{prev} \leftarrow N_{prev} + N_i$$
.

- (e) Iterate over all patches i, increment total power $P_i \leftarrow P_i + \delta P_i$, replace unshot power $\Delta P_i \leftarrow \delta P_i$, and clear received power $\delta P_i \leftarrow 0$. Compute new total unshot power ΔP_T on the fly.
- (f) Display image using P_i .

Stochastic Jacobi Radiosity

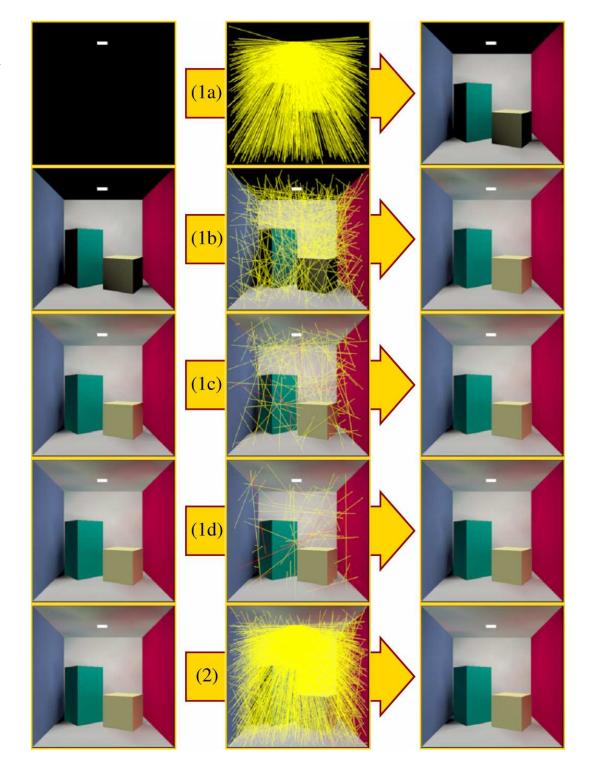
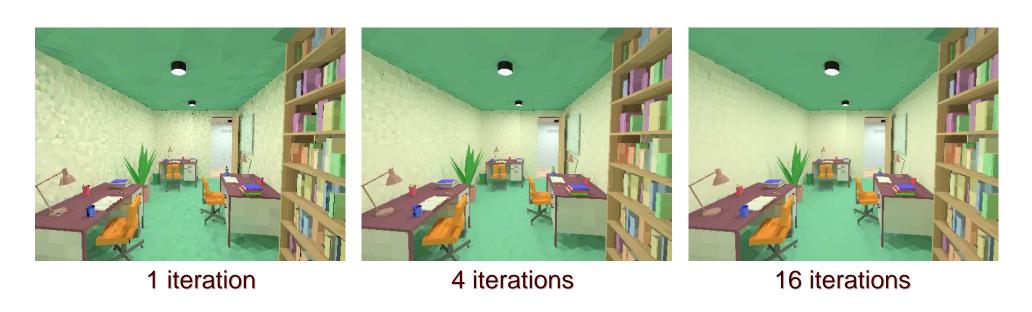


Fig 6.6, [Dutré et al. 2003]



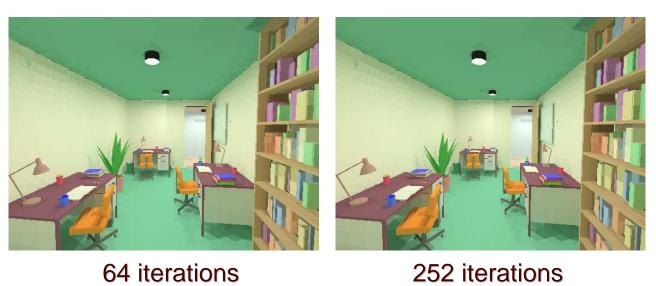


Fig 6.7, [Dutré et al. 2003]

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Surface Meshing

- Store radiosity across surface
 - Few elements
 - Represents function well
 - Few visible artifacts

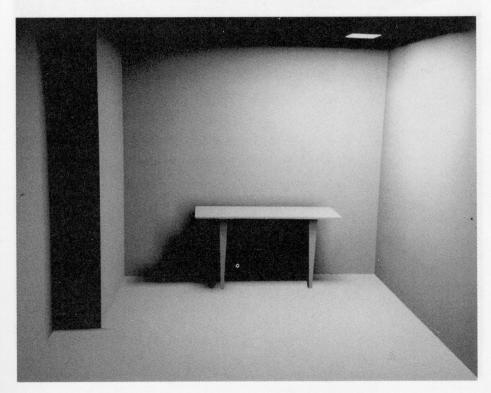
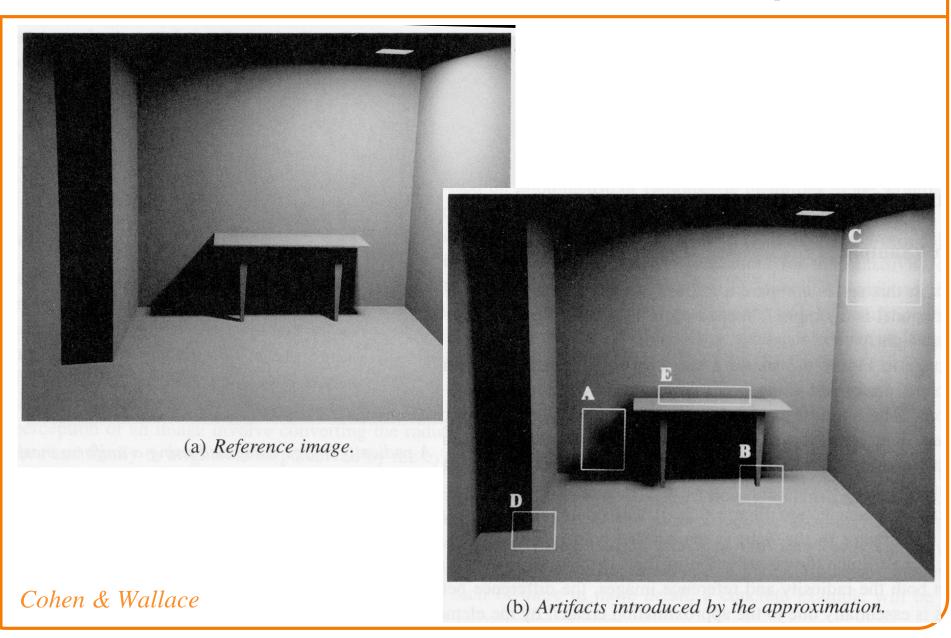
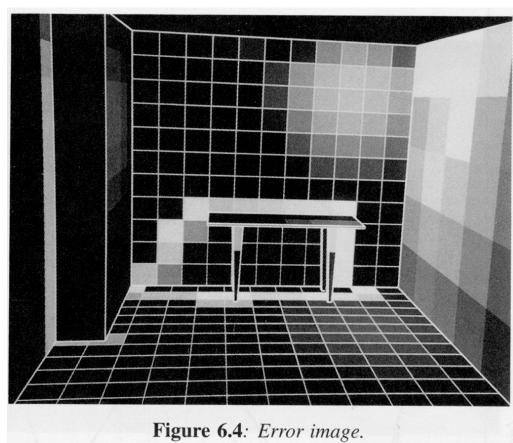


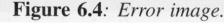
Figure 6.2: A radiosity image computed using a uniform mesh.

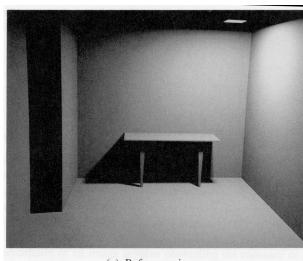
Artifacts of Bad Surface Meshing



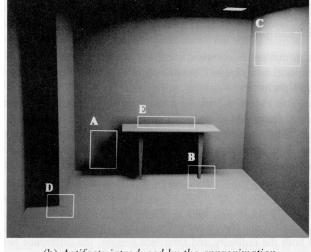
Error Image







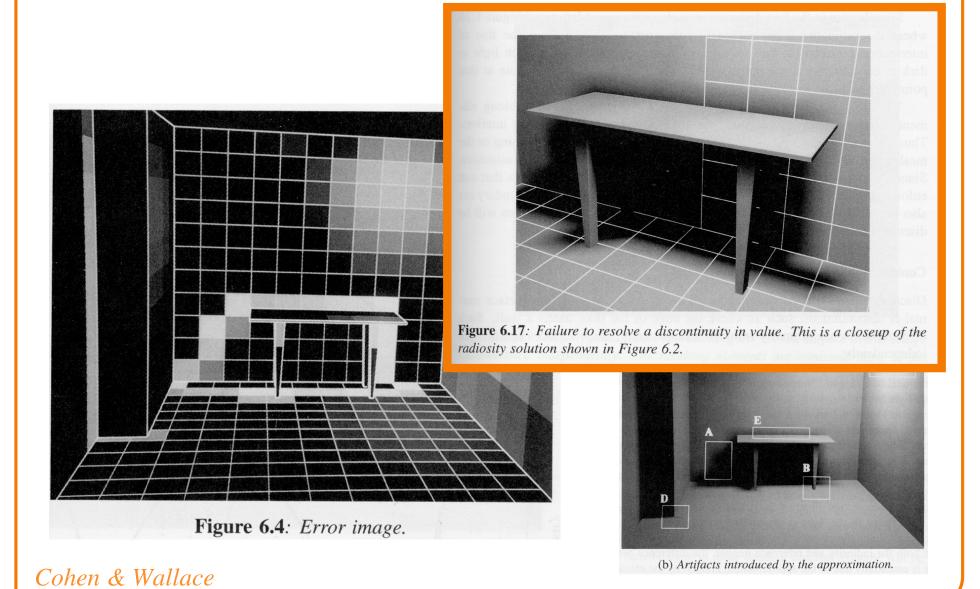
(a) Reference image.



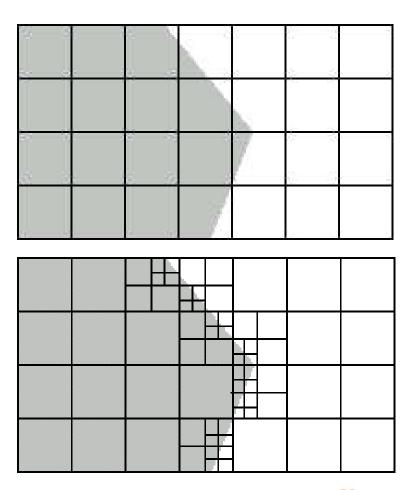
(b) Artifacts introduced by the approximation.

Cohen & Wallace

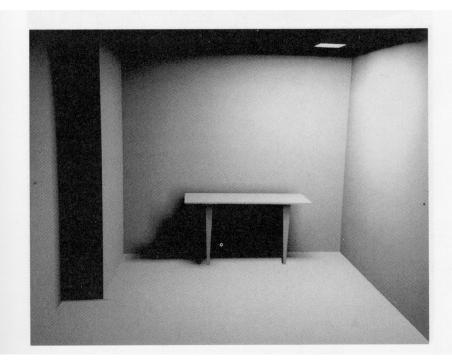
Error Image

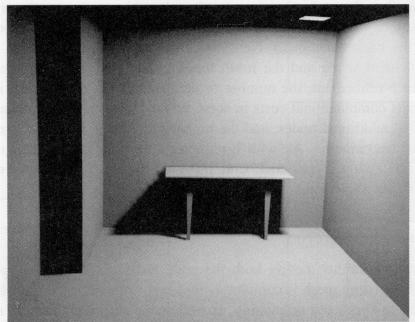


Refine mesh in areas of high residual



Yeap





Uniform mesh

Adaptive mesh

Cohen & Wallace

Error Comparison

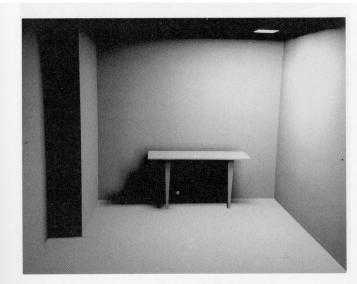
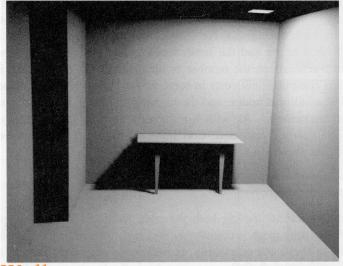


Figure 6.2: A radiosity image computed using a uniform mesh.



Cohen & Wighte 6.2. Adaptive subdivision. Compare to Figure 6.2.

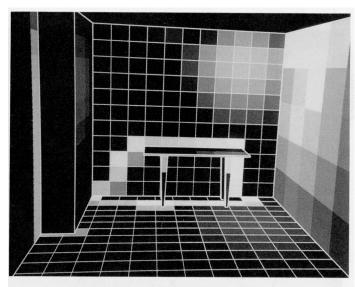


Figure 6.4: Error image.

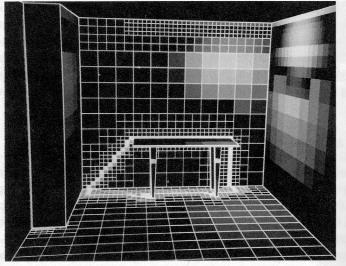
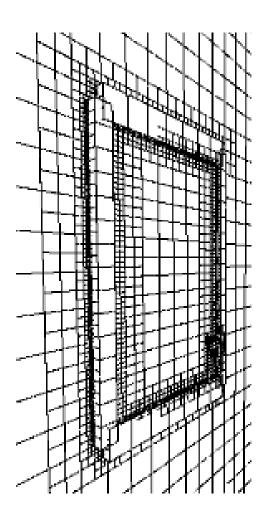
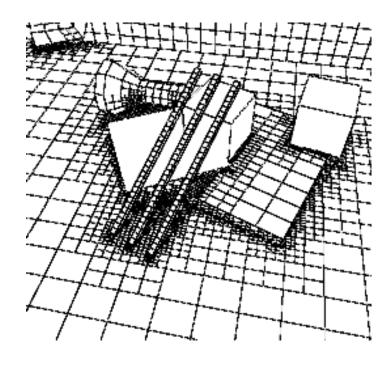


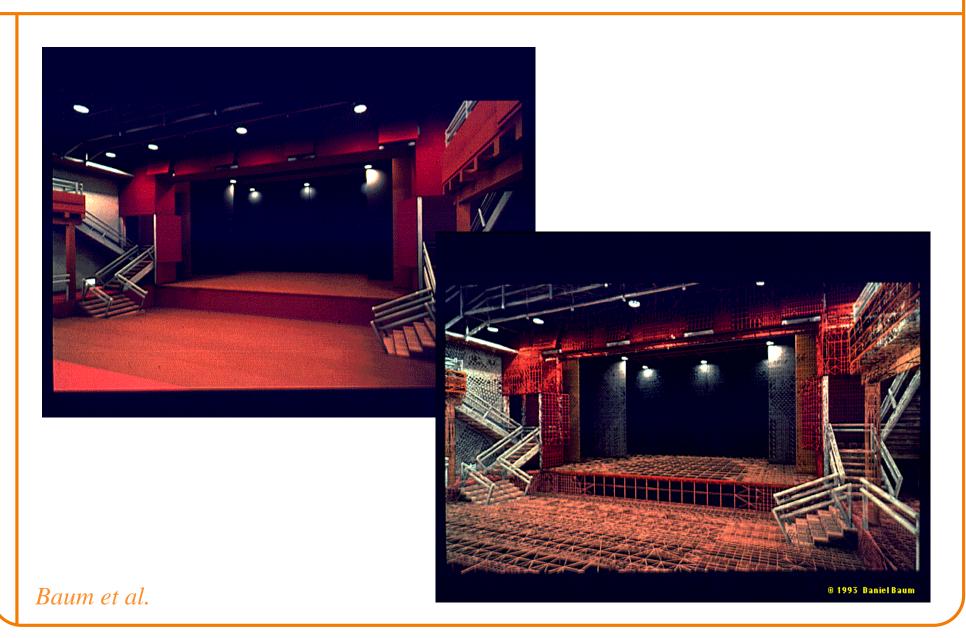
Figure 6.24: Error image for adaptive subdivision. Compare to Figures 6.4 and

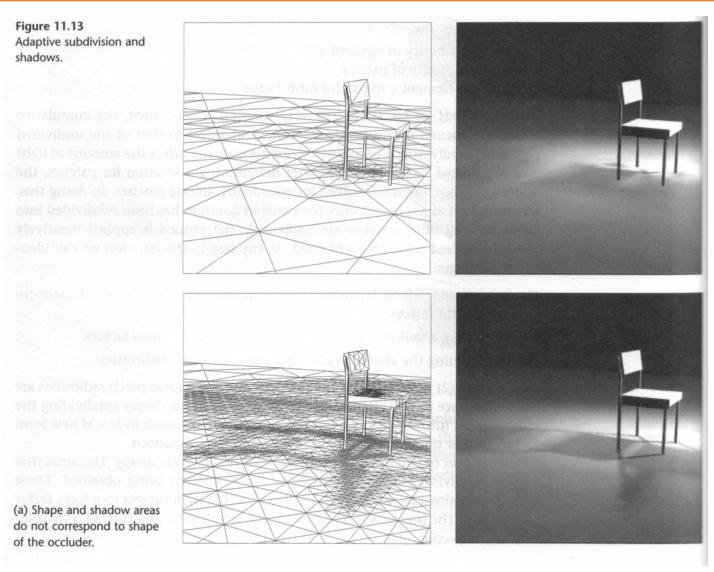








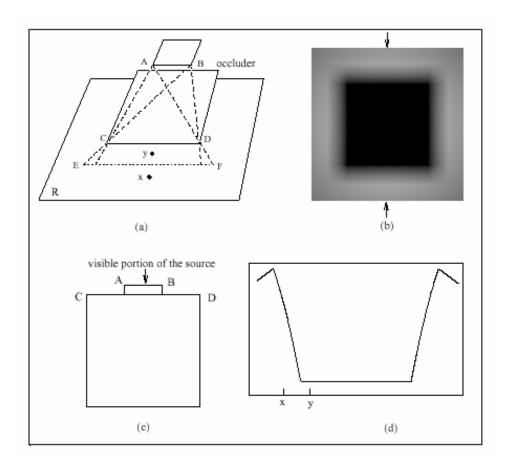




Watt, 3D CG, 3rd Ed, 2000

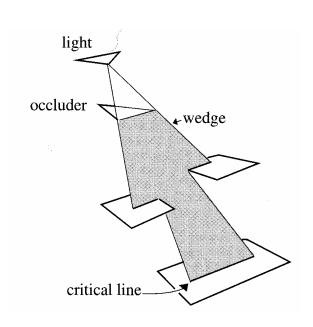
Discontinuity Meshing

 Capture discontinuities in radiosity across a surface with explicit mesh boundaries



Discontinuity Meshing

 Capture discontinuities in radiosity across a surface with explicit mesh boundaries



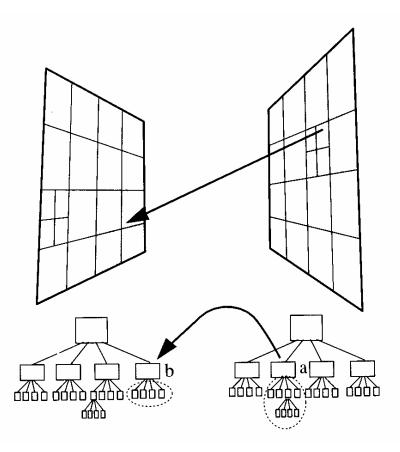


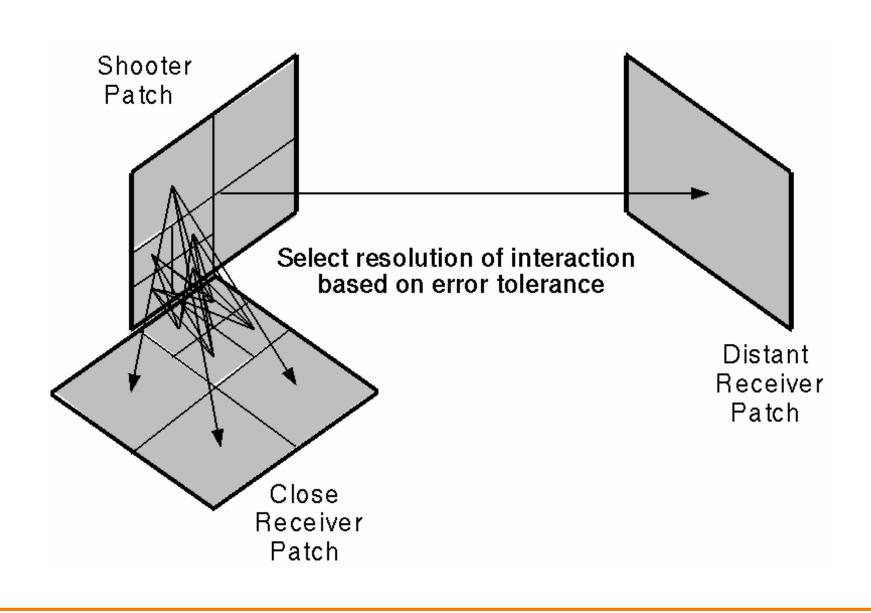


Discontinuity Mesh

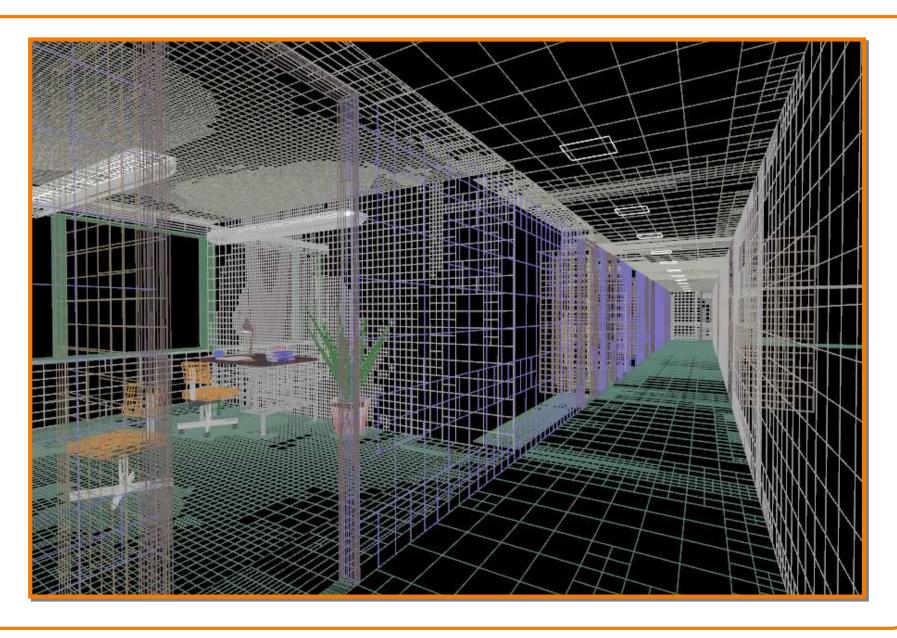
Lischinski et al.

- Refine elements hierarchically:
 - Compute energy exchange at element granularity satisfying a user-specified error tolerance









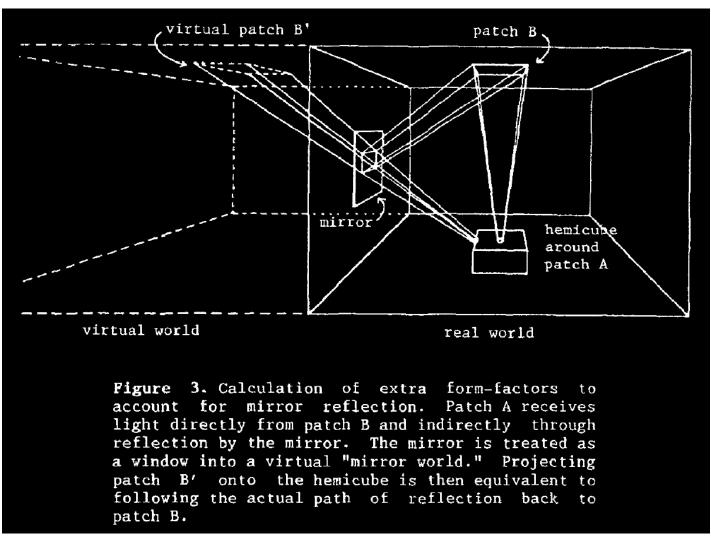


Extensions

- Non-diffuse environments
 - Directional radiosity functions
 - Extended form factors
 - Multipass methods
- Participating media
 - Path integrals in form factors
- Dynamic scenes
 - Incremental updates
- Parallel solvers
 - Decomposition
 - Scheduling



Example: Mirror Form Factors



Wallace, Cohen & Greenberg, 1987

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Overview

- Radiosity equation
- Solution methods
 - Computing form factors
 - Selecting basis functions for radiosities
 - Solving linear system of equations
 - Meshing surfaces into elements
 - Rendering images
- Discussion

Instant Radiosity

- Alexander Keller, SIGGRAPH 97
- "Particle-based" sampling method
 - Avoids discretization of underlying integral equation
 - Similar to photon mapping (density estimation) idea
- Idea: Hardware accelerated rendering of "photons" as point lights w/ shadows
- Limited number of photons due to sampling used
 - Deterministic quasi-random walk traces photons
 - » Jittered low-discrepancy sampling
- Fast and simple to implement

Basic Idea

 Render point-light sources into pixel (m,n) to evaluate integrals using graphics hardware

$$\langle L, \Psi \rangle := \int_{S} \int_{\Omega} L(y, \vec{\omega}) \Psi(y, \vec{\omega}) \cos \theta d\omega dy$$

$$\Psi_{mn}(y,\vec{\omega}) := \frac{\delta(\vec{\omega} - \vec{\omega}_{y_f})}{\cos \theta} \frac{1}{|P_{nm}|} \chi_{P_{mn}}(h(y,\vec{\omega}))$$

- How? Use M point lights $L(y) \approx \sum_{i=0}^{\infty} L_i \delta(y P_i),$
- → Integrals become trivial to evaluate

$$\overline{L}_{mn} := \langle L, \Psi_{mn} \rangle = \langle L_e, \Psi_{mn} \rangle + \langle T_{fr} L, \Psi_{mn} \rangle
= \langle L_e, \Psi_{mn} \rangle + T_{mn} L$$

Sampling: Choose the particles wisely

3.1 Quasi-Monte Carlo Integration

Similar to [Kel96b], we use the method of quasi-Monte Carlo integration for the evaluation of the integrals. For the approximation

$$\int_{[0,1)^s} f(x)dx \approx \frac{1}{N} \sum_{i=0}^{N-1} f(x_i)$$

we generate the N sample points $x_0, \ldots, x_{N-1} \in [0, 1)^s$ using the Halton sequence. Based on the radical inverse function

$$\Phi_b(i) := \sum_{j=0}^{\infty} a_j(i) \ b^{-j-1} \in [0,1) \Leftrightarrow i = \sum_{j=0}^{\infty} a_j(i) \ b^j,$$

the s-dimensional Halton low discrepancy sequence (see comparison in figure 1) is

$$x_i = (\Phi_{b_1}(i), \dots, \Phi_{b_s}(i)), i \in \mathbb{N}_0,$$

where b_j is the j-th prime number. Note that each segment $P_{N'}$ of

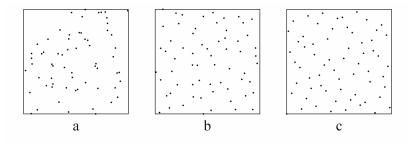


Figure 1: Two-dimensional uniform sampling patterns: a) random, b) jittered, and c) Halton for N=64 samples.

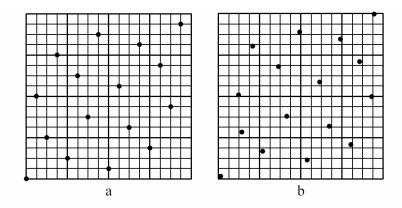


Figure 5: Grid structure of the a) Hammersley and b) jittered Hammersley sampling patterns for N=16.

Sampling

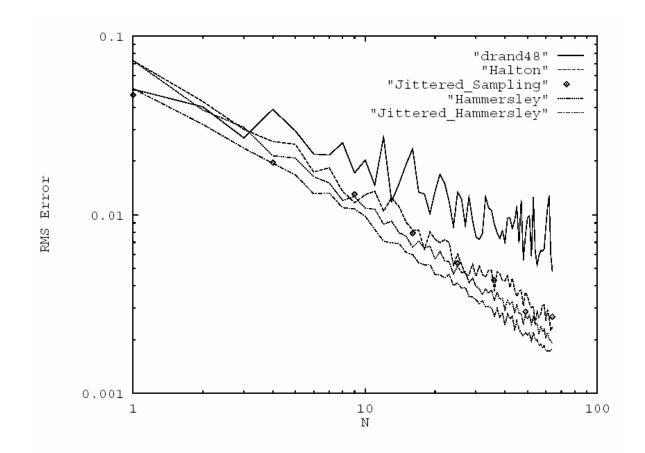


Figure 6: Convergence of different sampling patterns.

Quasi-Random Walk

$$L = (I - T_{f_r})^{-1} L_e = \sum_{j=0}^{\infty} T_{f_r}^j L_e$$

converges and can be used to solve the integral equation. Inserted in (1), after some transformations, for the radiosity setting we get

$$T_{mn}L = \frac{1}{|P_{mn}|} \sum_{j=0}^{\infty} \int_{P_{mn}} \int_{\Omega^j} \int_{S_e} p_j(y_0, \vec{\omega}_0, \dots, \vec{\omega}_j)$$
$$V(y_j, y') f_d(y') \frac{\cos \theta_j \cos \theta'}{|y_j - y'|^2} dy_0 d\omega_0 \cdots d\omega_j dP. \tag{3}$$

Here $S_e := \sup L_e \subseteq S$ is the support of the light sources, and $y' = h(y_f, P - y_f) \in S$ is the first point hit when shooting a ray from the eye at y_f through the point $P \in P_{mn}$ into the scene. $V(y_j, y')$ checks the mutual visibility of the points y_j and y', yielding 1 in case of visibility and 0 else. The radiance density

$$p_j(y_0, \vec{\omega}_0, \dots, \vec{\omega}_j) := L_e(y_0) \prod_{l=1}^j (\cos \theta_{l-1} f_d(y_l))$$

Quasi-Random Walk

 Assumption: mean reflectivity representative of local reflectivity in scene

$$\overline{\rho} := \frac{\sum_{k=1}^{K} \rho_{d,k} |A_k|}{\sum_{k=1}^{K} |A_k|} \approx ||T_{f_d}||,$$

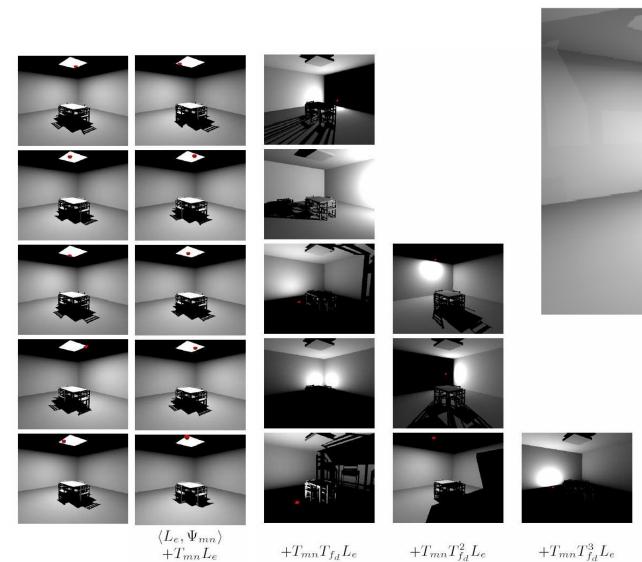
- Use mean reflectivity to estimate number of particles reflected at each generation
- Total number of particles is bounded by

$$M < \sum_{j=0}^{\infty} \overline{\rho}^{j} N = \frac{1}{1 - \overline{\rho}} N =: \overline{l} N$$

Pseudocode

```
void InstantRadiosity(int N, double \overline{\rho})
     double w, Start; int End, Reflections = 0;
     Color L; Point y; Vector \vec{\omega};
     Start = End = N;
     while (End > 0)
          Start *= \overline{\rho};
          for (int i = (int) Start; i < End; i++)
               // Select starting point on light source
               y = y_0(\Phi_2(i), \Phi_3(i));
               L = L_e(y) * supp L_e;
               w = N;
               // trace reflections
               for(int j = 0; j \le Reflections; j++)
                    glRenderShadowedScene (\frac{N}{|w|}L, y);
                    glAccum (GL_ACCUM, \frac{1}{N});
                    // diffuse scattering
                    \vec{\omega} = \vec{\omega}_d(\Phi_{b_{2j+2}}(i), \Phi_{b_{2j+3}}(i));
                    //trace ray from y into direction \vec{\omega}
                    y = h(y, \vec{\omega});
                    // Attenuate and compensate
                    L \star = f_d(y);
                    w *= \overline{\rho};
          Reflections++;
          End = (int) Start;
     glAccum(GL_RETURN, 1.0);
```

Algorithm (N=10, M=20)



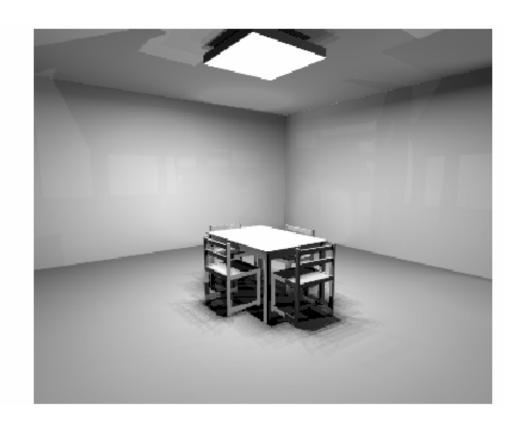


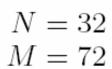




 $+T_{mn}T_{f_d}^4L_e$

$$\begin{aligned} N &= 10 \\ M &= 20 \end{aligned}$$







$$N = 64$$
$$M = 147$$

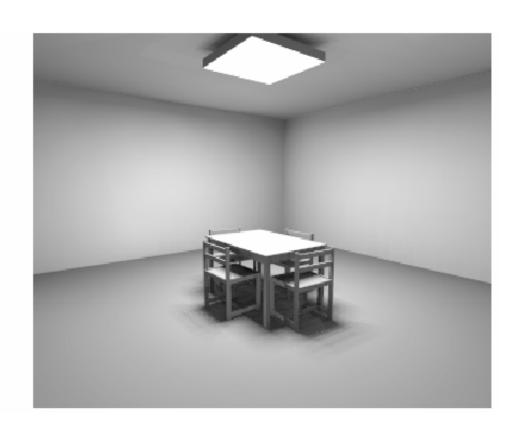




Figure 8: Specular effects of the standlight on the floor by using the full BRDF f_r in T_{mn} for N=128.

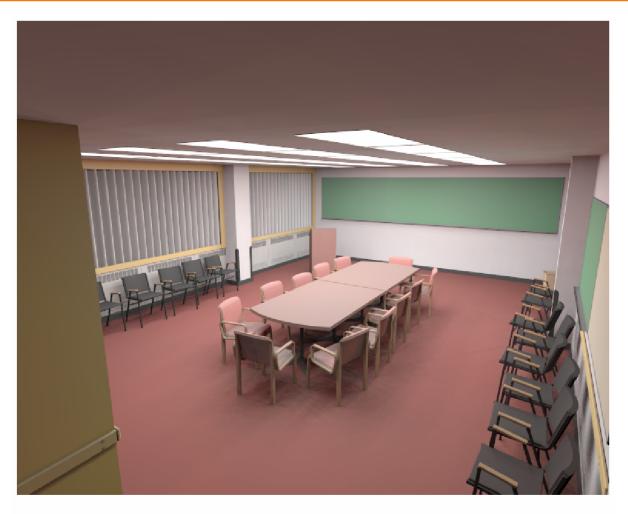


Figure 9: Conference room image for N=128.